

1 **Characterizing the evolution of mass flow properties and dynamics through analysis of**  
2 **seismic signals: Insights from the 18 March 2007 Mt. Ruapehu lake-breakout lahar**

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4 Braden Walsh<sup>1</sup>, Charline Lormand<sup>2</sup>, Jon Procter<sup>3</sup>, Glyn Williams-Jones<sup>1</sup>

5 <sup>1</sup>Department of Earth Sciences, Simon Fraser University, Burnaby, British Columbia, Canada

6 <sup>2</sup>Department of Earth Sciences, University of Durham, Durham, DH1 3LE, UK

7 <sup>3</sup>Volcanic Risk Solutions, Institute of Agriculture and Environment, Massey University,

8 Palmerston North, New Zealand

9

10 Corresponding Author: Braden Walsh ([braden\\_walsh@sfu.ca](mailto:braden_walsh@sfu.ca))

11

12 **Abstract**

13 Monitoring for lahars on volcanoes can be challenging due to the ever-changing landscape  
14 which can drastically transform the properties and dynamics of the flow. These changes to the  
15 flows require the need for detection strategies and risk assessment that are tailored not only  
16 between different volcanoes, but at different distances along flow paths as well. Being able to  
17 understand how a flow event may transform in time and space along the channel is of utmost  
18 importance for hazard management. While visual observations and simple measuring devices in  
19 the past have shown how lahars transform along the flow path, these same features for the

20 most part have not been described using seismological methods. On 18 March 2007 Mt.  
21 Ruapehu produced the biggest lahar in New Zealand in over 100 years. At 23:18 UTC the tephra  
22 dam holding the Crater Lake water back collapsed causing  $1.3 \times 10^6 \text{ m}^3$  of water to flow out and  
23 rush down the Whangaehu channel. We describe here the seismic signature of a lake-breakout  
24 lahar over the course of 85 km along the Whangaehu river system using three 3-component  
25 broadband seismometers installed <10 m from the channel at 7.4, 28, and 83 km from the  
26 crater lake source. Examination of 3-component seismic amplitudes, frequency content, and  
27 directionality combined with video imagery and sediment concentration data were used. The  
28 seismic data shows the evolution of the lahar as it transformed from a highly turbulent out-  
29 burst flood (high peak frequency throughout), to a fully bulked up multi-phase  
30 hyperconcentrated flow (varying frequency patterns depending on the lahar phase) to a slurry  
31 flow (bedload dominant). Estimated directionality ratios show the elongation of the lahar with  
32 distance down channel, where each recording station shows a similar pattern, but for differing  
33 lengths of time. Furthermore, using directionality ratios shows extraordinary promise for lahar  
34 monitoring and detection systems where streamflow is present in the channel.

### 35 1. Introduction

36 Volcanic mass flows (e.g. debris flows, pyroclastic density currents, debris avalanches) are one  
37 of the greatest threats to communities, industry, recreation, etc. on and around volcanoes.  
38 These volcanic mass flows are particularly dangerous as they are fast moving turbulent flows  
39 that can occur without any warning or an eruption transpiring (Capra et al., 2010). These flows  
40 can move a sizable amount of liquid and debris great distances that can critically impact

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43 locations hundreds of kilometers from the volcano or source. Lake-breakout or outburst flood  
44 events can be particularly destructive because they tend to be larger and can cause long lasting  
45 changes to the landscape and surrounding ecosystems (O'Connor et al., 2013; Procter et al.,  
46 2021). Furthermore, unlike eruption or rain triggered mass flows, outburst floods have very  
47 little to no warning. Eruption sources can be prepared for by the onset of the eruption and/or  
48 the monitoring of the volcano through various methods (e.g. seismology, infrasound, gravity,  
49 gas and water chemistry). Likewise, for rain-induced flows using techniques such as the amount  
50 or intensity of rain (e.g. Capra et al., 2010; 2018) or by monitoring the amount of available  
51 material (e.g. Iguchi, 2019) can help forecast when an event may occur.

52 In New Zealand, there have been numerous cases of large damaging mass flows in modern  
53 times. In October 2012, a lake-breakout lahar originating from Te Maari, destroyed hiking trails  
54 and forestry, eventually flowing over 4.5 km to damage and block off Highway 46 (Procter et al.,  
55 2014; Walsh et al., 2016). Moreover, on 24 December 1953, the deadliest lahar in New Zealand  
56 history occurred killing 151 people when a lahar struck a train crossing at the Tangiwai Rail  
57 Bridge, 39.8 km from the Crater Lake on top of Mt. Ruapehu (O'Shea, 1954). The ability to  
58 predict and investigate the changing dynamics and properties of large volcanic mass flows as  
59 they progress down channel is the first step in beginning to understand flow mechanisms  
60 better, and ultimately address the hazards involved to mitigate the risk.

61 In order to better characterize and understand these flow events, many applications and  
62 instruments have been used in the past (e.g. trip wires, stage gauge, load cells, pore pressure).  
63 While many of these tools can yield quick assessments and provide ample warning (e.g. current

64 meters, trip wires), they can sometimes be at risk of false detections, equipment damage or  
65 loss, and/or lack the capability to evaluate multiple pulses or flow events. Geophysical  
66 instruments (e.g. seismometers, geophones, infrasound) on the other hand can be installed at a  
67 safe distance away from the channel and have shown signs of not only being capable warning  
68 systems (e.g. Coviello et al., 2019), but have the ability to accurately estimate flow properties  
69 (e.g. Arattano and Marchi, 2005; Doyle et al., 2010; Schimmel et al., 2021), as well as flow  
70 dynamics (e.g. Gimbert et al., 2014; Coviello et al., 2018; Walsh et al., 2020). However, in order  
71 to fully utilize these instruments, improved interpretation, comprehension, assessment, and  
72 universality is needed. One technique to increase the ability to predict, warn, and estimate the  
73 properties and dynamics of flow events is to use all three components of the seismic recording.  
74 Recently, several studies have shown that using all three components is effective in  
75 characterizing flow events (e.g. snow-slurry lahars, Cole et al., 2009; snow avalanches, Kogelnig  
76 et al., 2011; streamflow, Roth et al., 2016; landslides, Surinach et al., 2005; lahars, Walsh et al.,  
77 2020; rockfalls, Kuehnert et al., 2021; hyperconcentrated flows, Walsh et al., 2016). Using the  
78 horizontal components along with the vertical component can yield additional information  
79 about the flow that is not utilized if only the vertical component is used. Notably, directionality  
80 analysis (e.g. Doyle et al., 2010; Walsh et al., 2020) can provide information about wetted  
81 perimeter, sediment concentration, and number of particle collisions. Furthermore, differing  
82 energies and frequency outputs from channel parallel and channel perpendicular signals can  
83 point to specific changes within the flow (Burtin et al., 2010; Roth et al., 2016) that can provide  
84 insights into the internal dynamics.

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## 85 1.1 Anatomy of lahars

92 When a lahar is created from a lake-breakout or outburst flood event, the transition from flood  
93 or streamflow torrent depends on the erosivity of the channel and the supply of sediment being  
94 entrapped within the flow (e.g. Scott, 1988; Doyle et al., 2011). An event may start as a highly  
95 turbulent low sediment flow, then transform into a hyperconcentrated flow, and may even  
96 eventually 'bulk up' to exhibit characteristics of a debris flow with the possibility of laminar or  
97 plug-like behavior (Scott, 1988, Pierson et al., 1990). At Mt. Ruapehu, the propagational  
98 differences of lahars down channel have been observed and characterized in the past (e.g.  
99 Cronin et al., 1996; Cronin et al., 1999; Cronin et al., 2000; Manville et al., 2000; Procter et al.,  
100 2010a; Lube et al., 2012). From these studies, models of how lahars bulk up and transition  
101 throughout the run-out distance have been postulated. For the lahars in the Whangaehu  
102 channel, Cronin et al. (1999) created three 4-phase conceptual models based on source  
103 distances of 23.5 km, 42 km, and >55 km. The first two models are for lahar regimes, whereas  
104 the third model described a lahar almost at its peak run-out distance. In each model, the first  
105 phase consists of a super charged streamflow pulse that flows ahead of the head of the flow.  
106 This phenomenon has also been noted for debris flows interacting with streamflow (Arattano  
107 and Moia, 1999). Furthermore, discharge is maximum at the transition between phase 1 and  
108 phase 2 (Cronin et al., 1999). Phase 2 is described as a mixing zone between streamflow and  
109 increasing sediment content, where the peak sediment concentration usually occurs at the end  
110 of phase 2 or at the beginning of phase 3 (e.g. Pierson and Scott, 1985). Cronin et al. (1999)  
111 defined phase 3 as the lahar body, which has the least amount of the original streamflow  
112 contained within. Phase 3 is also characterized by coarse sediment suspensions and is the most  
113 likely location for debris flow rheology. Finally, phase 4 is the tail of the lahar where debulking

114 and dilution occurs transforming the lahar back into a hyperconcentrated, mixed, or  
115 streamflow.

### 116 1.2 18 March 2007 lake-breakout event

117 Mt. Ruapehu (2797 asl) is the largest stratovolcano in the central North Island of New Zealand  
118 (Figure 1) which sits at the southwestern end of the Taupō Volcanic Zone (TVZ). The volcano has  
119 a volume of 110 km<sup>3</sup> which is composed of several overlapping cone building formations and  
120 surrounding ring plain volcanoclastics (Carrivick et al., 2009; Pardo et al., 2012). On top of the  
121 volcano, above the currently active vent sits a 1x10<sup>7</sup> m<sup>3</sup> acidic crater lake (Procter et al., 2010a).

122 The Whangaehu channel is the preferred outlet for Crater Lake water and lahars in recent  
123 history (Procter et al., 2012; Procter et al., 2021). The Whangaehu channel is on the eastern  
124 flank of Mt. Ruapehu where it runs down across the volcanic ring plane where it eventually  
125 heads southwest for ~200 km reaching the Tasman Sea.

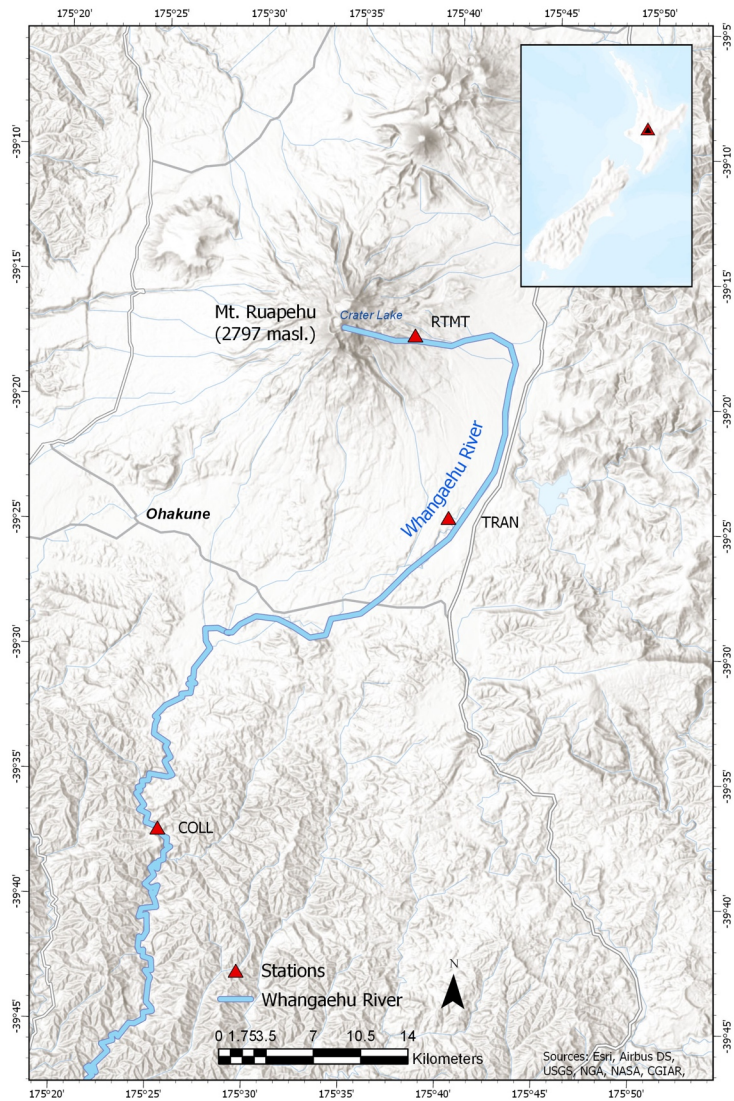
126 Prior to the events that took place in the morning local time on 18 March 2007, a heavy  
127 rainstorm occurred accumulating about 256 mm of water over the 10 hours prior to the dam  
128 breach that led to the outburst flood (Massey et al., 2010). The intense rain caused the Crater  
129 Lake to rise an extra 6.4 m above the natural lava formation ledge, which started to cause  
130 seepage and extra water entering the Whangaehu gorge (Carrivick et al., 2009). At ~11:18 NZT

131 (UTC +12), the tephra dam collapsed causing 1.3x10<sup>6</sup> m<sup>3</sup> of water to flow out of the lake and  
132 into the Whangaehu channel (Procter et al., 2010a). The dam was eroded and undercut in  
133 multiple stages resulting in a series of retrogressing landslides along with the main debris  
134 flow/lahar channel.

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139 Figure 1 Map of Mt. Ruapehu and the surrounding area located on the central North Island of New Zealand. Blue  
 140 outline represents the Whangaehu channel and the path the 18 March 2007 lahar traveled down. Red triangles  
 141 denotes the three monitoring stations along the Whangaehu channel at 7.4, 28, and 83 km.

142 Since the lahar was caused by lake-breakout dynamics and thus an abundance of water, the  
143 event was classified as a hyperconcentrated streamflow rather than a sediment-filled debris  
144 flow (Procter et al., 2010b). At ~8.0 km from source velocity measurements recorded the flow  
145 at ~9.5 m/s and had an estimated 6 m of downcutting showing the ability for the lahar to  
146 deposit and erode massive amounts of material (Procter et al., 2010a,b). Furthermore, the 18  
147 March 2007 lahar was one of the most thoroughly monitored lahars ever (Manville and Cronin  
148 2007). In total there were 21 monitoring locations setup to measure various lahar properties,  
149 (e.g. flow monitor, camera, stage height, flow sampling, pore-pressure, seismic, etc.) along the  
150 channel (Keys and Green, 2008; Lube et al., 2012), with the lahar taking over 16 hours to  
151 eventually travel out to the New Zealand coast, ~200 km from the original crater lake source.

152 Here, we delve into the properties of a lake-breakout hyperconcentrated streamflow that  
153 bulked up to a volume of  $\sim 4.4 \times 10^6 \text{ m}^3$  (Procter et al., 2010a) over the course of 83 km that  
154 occurred on 18 March 2007 along the Whangaehu channel originating from Mt. Ruapehu, New  
155 Zealand. The combination of seismic analysis (frequency and directionality) with supplementary  
156 measurements (e.g. video, sediment concentration) show how a lahar transforms over time and  
157 distance and how using these seismic techniques can help monitor the ever changing dynamics  
158 and properties of a flow event. Furthermore, we examine previous models of the evolution of a  
159 lahar and compare the model with the seismic data available.

## 160 2. Data

161 The seismic data for the 18 March 2007 lahar was recorded on three seismometers installed at  
162 various distances (7.4, 28, 83 km) along the Whangaehu channel (Figure 1). The three 3-

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165 component broadband Guralp 6T sensors (COLL, RTMT, TRAN) recorded data at 100 Hz  
166 sampling and had GPS time stamps. For each site, the seismometers axes were installed to true  
167 North and the recorded data were rotated to align North as flow parallel (P) and East as the  
168 cross-channel direction (T). Furthermore, the seismometers were installed normal to horizontal  
169 to lessen the degree of vertical energy transfer to the horizontal components. The monitoring  
170 station Round the Mountain Track (RTMT), was installed 4 m from the channel and 7.4 km  
171 downstream from the source of the lahar. The lahar arrived at RTMT at 23:36 UTC and had an  
172 average velocity of 9.3 m/s (Figure 2a). The Trans Rail Gauge (TRAN) station was installed 28 km  
173 from source and 10 m from the channel, which also included a video camera that captured an  
174 image every 30 seconds. The lahar arrived at TRAN at 24:35 UTC with an average velocity of 5.6  
175 m/s (Figure 2d). The Colliers Bridge (COLL) station was installed 10 m from the channel and 83  
176 km from source. The lahar arrived at COLL at 04:13 UTC and had an average velocity of 4.8 m/s  
177 (Figure 2f). Arrival times are based off of images and eye witnesses at each of the monitoring  
178 stations. The flow velocity at RTMT and COLL were estimated from imagery and at TRAN from a  
179 flow meter.

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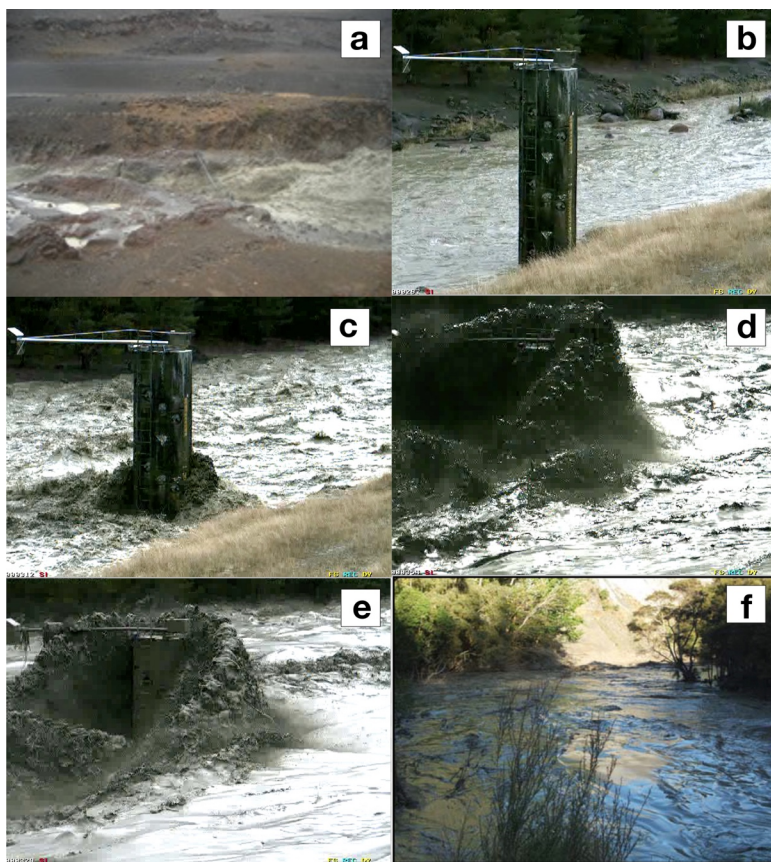
### 180 3. Results

181 To examine the multi-component dynamics of the 18 March lake-breakout event along the  
182 Whangaehu channel at three monitoring locations, the data were corrected for instrument  
183 response and split into 10 s time windows. At each recording location, peak spectral frequency  
184 (PSF) amplitude, root mean squared (RMS) amplitude, and directionality ratios (DR) are  
185 estimated for each of the 10 s time windows. At each monitoring station the first hour of the

189 lahar including five minutes prior to the arrival are shown in all the results except when  
190 indicated.

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193

194 *Figure 2 Images from the 18 March 2007 lake break-out lahar from RTMT (a), TRAN (b, c, d, e), and COLL (f). Note*  
195 *the transformation of the lahar at TRAN from streamflow (b), increased discharge pre-lahar phase 1 pulse (bow*  
196 *wave) (c), head of the lahar (d), and low PSF beginning of lahar body (e).*

197 **3.1 Frequency analysis**

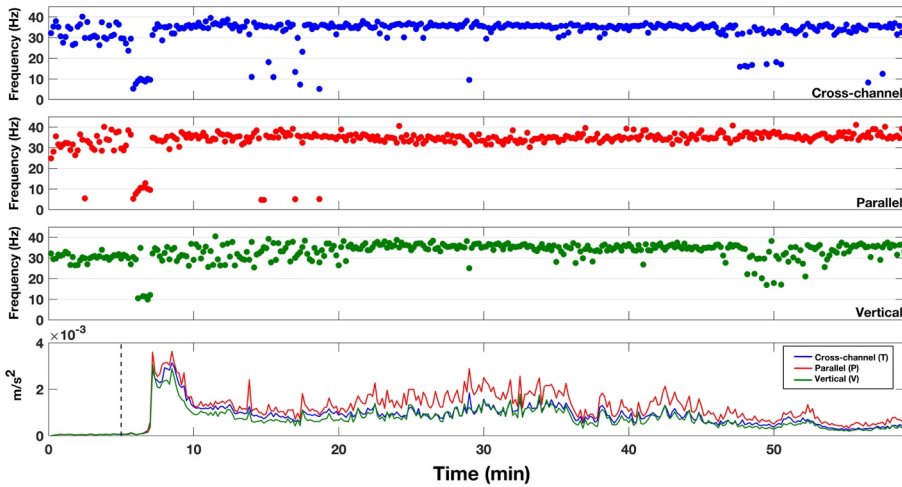
198 In order to examine the PSFs for all three components at each site along the channel, we use  
199 the frequency recorded at the maximum amplitude of the frequency spectra for each 10 s  
200 running time window. The PSF for RTMT (7.4 km from source) shows similar patterns between  
201 all three components (Figure 3). ~~The five~~ minutes prior to the arrival of the head (peak seismic  
202 amplitude) of the lahar ~~are~~ characterized by scattered PSFs between 20-40 Hz for the cross-  
203 channel (Figure 3, blue dots) and parallel (Figure 2, red dots) directions, while in the vertical  
204 direction (Figure 3, green dots) the PSF is ~30 Hz. When the front of the lahar arrives at the  
205 station, the PSF in all three components decreases to ~5-10 Hz for about 1 min before  
206 increasing again to higher frequencies. After ~~the~~ front of the lahar passes the station and when  
207 the head arrives, the PSF in the cross-channel and parallel directions remain between 30-40 Hz  
208 for the rest of the recording window. In the vertical component, the PSF is scattered between  
209 20-40 Hz for ~15 min after the arrival of the head of the lahar and then becomes narrower,  
210 similar to both the cross-channel and parallel components with PSFs between 30-40 Hz.

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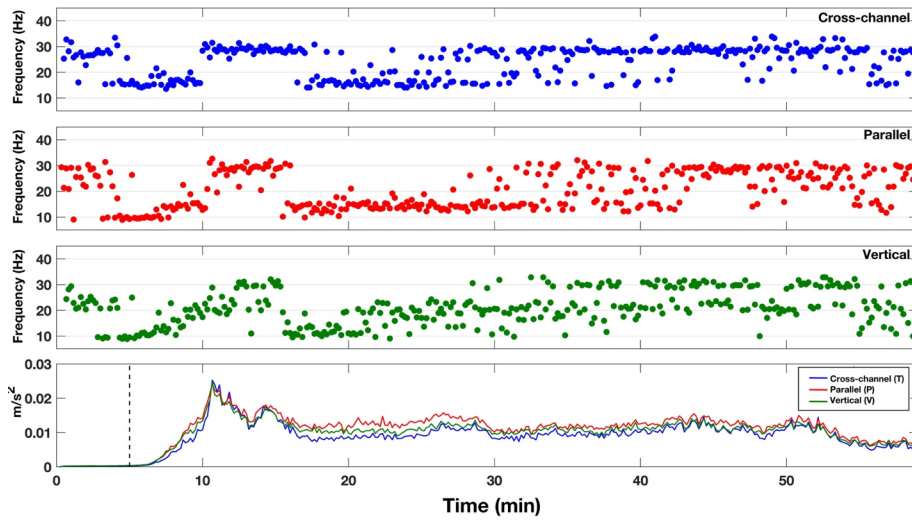
216 Figure 3 Peak spectral frequencies for RTMT (7.4 km from source) for cross-channel (blue dots), flow parallel (red  
 217 dots), and vertical (green dots) directions. Bottom row depicts the RMS amplitude of the lahar, color coded to the  
 218 same colors as the PSF. The dashed vertical line marks the timing of the lahar front passing the monitoring station.  
 219 All PSFs and RMS amplitudes were calculated using 10 s time windows.

220 Further down the channel at station TRAN (28 km from source), the PSFs for all three  
 221 components show a similar overall pattern (Figure 4). The pre-lahar PSF distribution in all three  
 222 components is between 20-32 Hz. Like RTMT higher up the channel, the PSFs for the front of  
 223 the lahar at TRAN first drops down to around 10 Hz and when the lahar head arrives (10 min,  
 224 Figure 4) the PSF increases to ~30 Hz for parallel (Figure 4, red dots) and cross-channel (Figure  
 225 4, blue dots) directions and between 20-30 Hz in the vertical component. This decrease to  
 226 lower frequencies before the head of the lahar at TRAN lasts for about 5 min. After the head of  
 227 the lahar passes the recording station the PSF content decreases for ~15 min to 10-20 Hz for  
 228 the parallel and cross-channel components and between 10-25 Hz for the vertical (Figure 4,  
 229 green dots) components. The PSF after the 30 minute mark in Figure 4 displays a bimodal

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232 pattern with frequencies between 10-35 Hz, with PSF time windows concentrating most at ~30  
233 Hz.

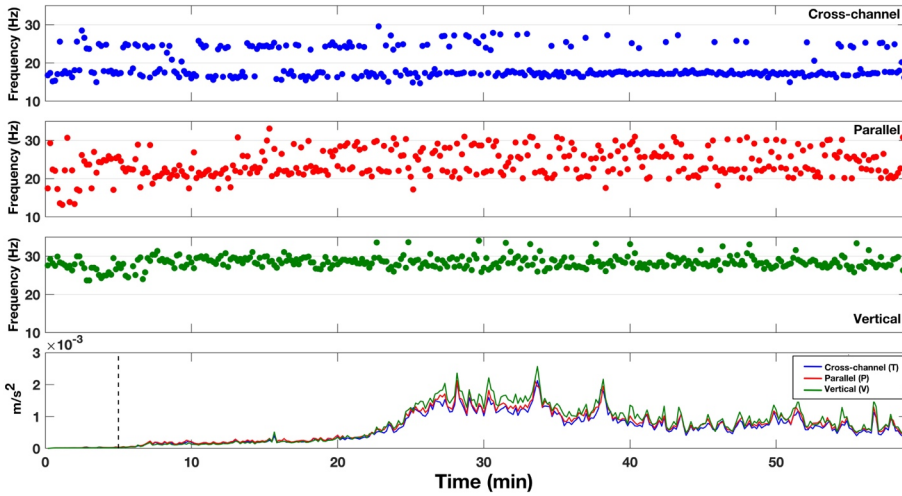
234 At the COLL recording station (83 km from source), the PSF distribution shows differing patterns  
235 for all three components (Figure 5). The PSF in the cross-channel direction (Figure 5, blue dots)  
236 depicts a bimodal pattern throughout with a strong lower concentration of time windows at  
237 ~18 Hz and a higher PSF at ~25 Hz. For the parallel component (Figure 5, red dots), the pre-  
238 lahar signal has a wide PSF range between 12-30 Hz. When the lahar arrives, the PSF becomes  
239 concentrated at ~22 Hz for ~8 min before transforming into a bimodal pattern similar to that of  
240 the cross-channel PSF, with frequencies between 20-30 Hz. In the vertical component (Figure 5,  
241 green dots), the PSF remains concentrated around ~28 Hz, only varying just prior to the arrival  
242 of the lahar and during the highest energy stage of the lahar (25-40 min).



243

244 Figure 4 Peak spectral frequencies for TRAN (28 km from source) for cross-channel (blue dots), flow parallel (red  
 245 dots), and vertical (green dots) directions. Bottom row depicts the RMS amplitude of the lahar color coded to the  
 246 same colors as the PSF. The dashed vertical line marks the timing of the lahar front passing the monitoring station.  
 247 All PSFs and RMS amplitudes were calculated using 10 s time windows.

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250 Figure 5 Peak spectral frequencies for COLL (83 km from source) for cross-channel (blue dots), flow parallel (red  
 251 dots), and vertical (green dots) directions. Bottom row depicts the RMS amplitude of the lahar color coded to the  
 252 same colors as the PSF. The dashed vertical line marks the timing of the lahar front passing the monitoring station.  
 253 All PSFs and RMS amplitudes were calculated using 10 s time windows.

254 **3.2 Directionality**

255 When recording mass flows with 3-component sensors, the directionality may be examined due  
 256 to the sensor being able to record signals in the two horizontal directions. The directionality  
 257 ratio allows for the determination of which horizontal component has stronger energy over the  
 258 course of the recording window. This is possible because, in channel side deployments for mass  
 259 flow monitoring systems, the sensor is either installed so that the North component is aligned  
 260 to be parallel to the flow or can be rotated during the data processing stage to align with the  
 261 channel orientation. Furthermore, with the channel side installations, attenuational factors can  
 262 mostly be ignored due to the close proximity to the channel and energy output of the flow  
 263 event. The directionality ratio (DR) can be defined as the cross-channel amplitude divided by

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268 the flow parallel amplitude. A DR > 1 indicates that the cross-channel amplitude is larger than  
269 that of the flow parallel, and vice-versa for a DR < 1. Directionality ratios have been used in the  
270 past to show rheology changes within flows for warning purposes, where the DR increases  
271 when streamflow transitions into a lahar (Walsh et al., 2020), and have been hypothesized to  
272 be an indicator for flow properties such as sediment concentration, wetted perimeter, and/or  
273 amount of particle collisions within a lahar (Doyle et al., 2010).

274 The directionality ratios for 10 s running time windows at each seismic station for the 18 March  
275 2007 lake-breakout lahar are shown in Figure 6. The DR for RTMT (Figure 6a) displays a DR  
276 around 1 pre-lahar, then decreases right before the lahar arrives at the recording station  
277 (Figure 6, dashed line), then as soon as the lahar passes, the DR increases to above DR = 1 for  
278 ~2 min. After the initial lahar flood pulse passes RTMT, the DR then proceeds to decrease below  
279 a DR = 1 for the rest of the recording window. Similar to RTMT, the DR for TRAN starts out with  
280 a DR < 1 (0.7-0.8) and as the lahar front passes, the DR similarly decreases to 0.6-0.7 before  
281 increasing to a DR > 1 for ~5 min when the lahar is at peak energy output starting at about the  
282 10 min mark (Figure 6d, red line). After the passing of the peak energy, the DR for TRAN  
283 decreases below 1 again for the remainder of the recording window. Further down the channel  
284 at COLL (Figure 6c), the DR before the lahar arrives has a wide range of values between 0.8-1.2.  
285 When the front of the lahar passes (Figure 6, dashed line), the DR stabilizes between 0.8-1,  
286 before increasing slightly when the peak energy of the lahar passes the monitoring site at about  
287 the 25 min mark.

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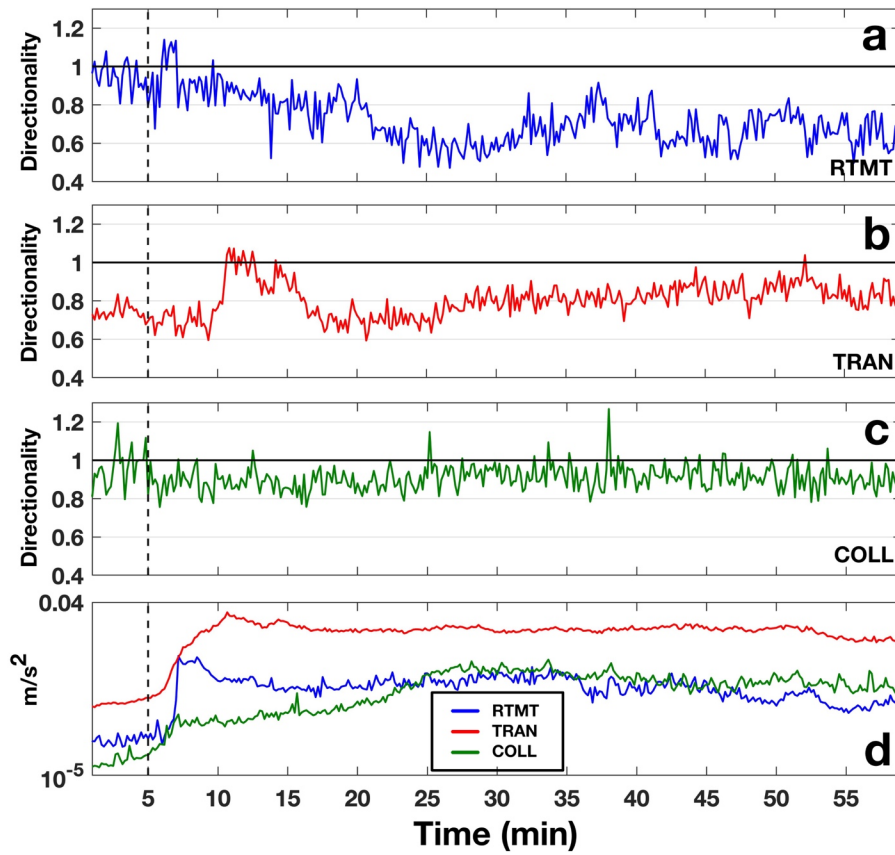
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294 *Figure 6 Directionality ratio plots over time for RTMT (a), TRAN (b), and COLL (c). Vertical RMS seismic signals for*  
295 *the three stations are plotted in (d) where blue is RTMT, red is TRAN and green represents COLL. The dashed*  
296 *vertical lines mark the timing of the lahar front passing the monitoring station. All DRs and RMS amplitudes were*  
297 *calculated using 10 s time windows.*

298 **4. Discussion**

#### 4.1 Frequency constraints

In order to obtain an understanding if PSFs are able to properly describe the lahar dynamics (i.e. the weight of the spectral amplitude at the PSF), frequency constraints must be analyzed. To complete this, normalized spectrograms along with spectral centroidal frequency (SCF) and spectral spreads are computed. The normalized spectrograms are estimated by normalizing the spectral amplitude for each 10 second time window of the lahar individually. By normalizing each time window, ranges of dominant frequencies can be visualized. SCFs are used because they represent the weighted average of the spectra, and yield the location (i.e. frequency) of the center of the spectral mass. The SCF of each time window is estimated similar to that of Saló et al. (2018), in which:

$$SCF = \frac{\sum_{f_1}^{f_2} f * A(f)}{\sum_{f_1}^{f_2} A(f)} \quad (1)$$

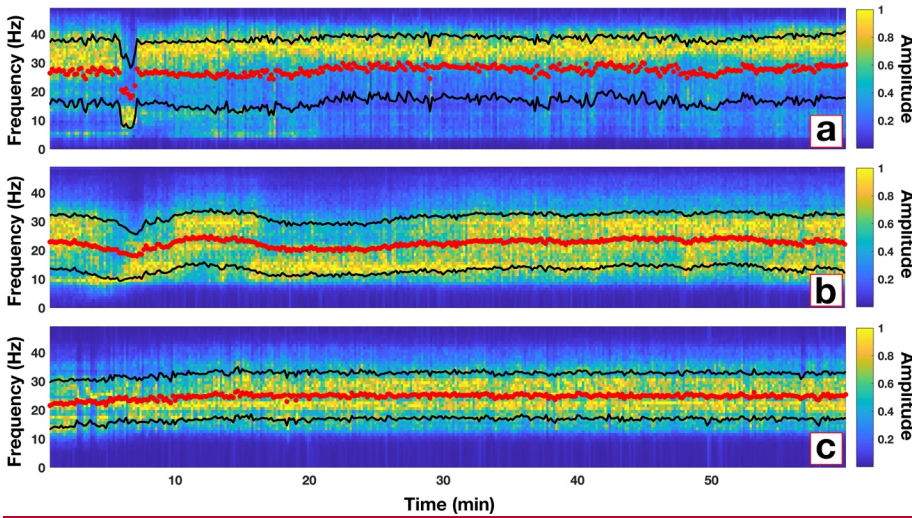
where  $f$  is the frequency and  $A(f)$  is the spectral amplitude associated with each frequency bin. The spectral spread measures the width of the spectral energy around the SCF, thus yielding information about the quality of the PSFs. Spectral spread can be estimated by:

$$S = \sqrt{\frac{\sum_{f_1}^{f_2} (f - SCF)^2 * A(f)}{\sum_{f_1}^{f_2} A(f)}} \quad (2)$$

The computed normalized spectrograms along with SCFs and spectral spreads for each of the three monitoring stations are shown in Figure 7. For simplicity and comparison, only the flow parallel data are shown. The normalized spectrograms for every station and component can be seen in Figures S1-S3.

318 The normalized spectrogram for RTMT (Figure 7a) yields very similar results to that of the PSF  
319 (Figure 3b), where most of the higher spectral amplitudes are at the same frequencies as those  
320 of the PSF. Notably, the low ~10 Hz signal immediately before the arrival of the head of the  
321 lahar is not only seen in the dominate normalized spectra, but also through the decrease in SCF.  
322 Additionally, the PSFs at these time windows are contained within the spectral spread (Figure  
323 7a, black lines). For TRAN, the normalized spectrogram (Figure 7b) is again, very similar to the  
324 PSF in Figure 4b. The SCF mirrors the pattern of the PSF with higher frequencies for the  
325 streamflow, a decrease for the front of the lahar, increase for the head of the lahar, decrease  
326 after the passing of the head, and finally a slight increase later in the lahar body. The  
327 normalized spectra yields this same pattern, with the late lahar body displaying the only  
328 timeframe with increased spectral amplitude distributed throughout the spectral spread (Figure  
329 7b, after 30 min). This most likely explains the bimodal distribution of PSFs for TRAN in Figure 4  
330 after the ~30 min mark. Continuing, the normalized spectrogram for COLL (Figure 7c), also  
331 shows similarities to that of the PSFs in Figure 5b. The PSFs for COLL range between ~20-30 Hz  
332 with a slight bimodal pattern. This same pattern can be seen where the higher spectral  
333 amplitudes are located (Figure 7c). Furthermore, the SCF for COLL splits the PSF range and stays  
334 at ~25 Hz during the bimodal phase of the PSF. Overall, with the analysis of the normalized  
335 spectrograms, SCFs and spectral spreads, we confirm that the use of PSFs to describe mass flow  
336 dynamics is concise for the 18 March 2007 lake-breakout lahar.

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338 *Figure 7 Normalized spectrograms for the flow parallel direction for each of the three monitoring sites along the*  
 339 *Whangaehu channel. Red dots represent the spectral centroidal frequency and black lines show the range of the*  
 340 *spectral spread. Note, normalized spectrograms for the other directions can be seen in Figures S1-S3.*

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341 **4.2 Evolution of lahar signals**

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342 A lahar propagating down channel can bulk up by collecting material from erosion or through  
 343 the coalescing of multiple pulses to shorten the total length of the lahar (Procter et al., 2010;  
 344 Doyle et al., 2011). Lahars can also debulk by depositional means or by the natural elongation  
 345 of the lahar as it progresses down channel (Doyle et al., 2011; Lube et al., 2012). Considering  
 346 the 18 March 2007 lake-breakout lahar was a large pulse of water that only mixed with the  
 347 existing streamflow and contained no juvenile material, examining the seismic signatures along  
 348 the flow path can be used to characterize the evolution and transformation of a lake-breakout  
 349 event from outburst flood to hyperconcentrated flow and beyond. At RTMT (Figure 3), the  
 350 seismic signature is dominated by the flow parallel direction with > 30 Hz PSF. The exception to

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353 this is the timeframe immediately before the head of the lahar passes when the PSF decreases  
354 to ~10 Hz. This low frequency signal can be seen at TRAN (Figure 4) and in the flow parallel  
355 component at COLL (Figure 5, red dots). However, at COLL the PSF is ~20 Hz instead of 10 Hz as  
356 in RTMT and TRAN due to flow properties at 83 km from source. This low PSF before the head  
357 of the lahar arrives at the station could represent the supercharged stream flow pulse (bow  
358 wave) that is pushed in front of the head of the lahar as described by Cronin et al. (1999) where  
359 they noticed these same pulses in front of lahar heads for three lahars on Mt. Ruapehu in 1995.  
360 Conversely, this frontal pulse could be from the uplift of streamflow from the faster moving  
361 underflow of the lahar (Manville et al., 2000). Furthermore, the low frequency zone before the  
362 head of the flow lengthens as the lahar progresses downstream, suggesting that lahar  
363 elongation can also be seen in the seismic frequency domain. The ~10 Hz PSF may be explained  
364 by flow processes (Schmandt et al., 2013; Barriere et al., 2015; Bartholomaeus et al., 2015) and  
365 could be due to the flow at this stage being more sensitive to discharge (Gimbert et al., 2014;  
366 Schmandt et al., 2017; Anthony, et al., 2018) or in the case of the underflow hypothesis,  
367 frictional sliding on the channel bed (Huang et al., 2004). The frontal surge or phase 1 of the  
368 lahar can be seen in the DR (Figure 6) as well. For every station along the channel the DR has a  
369 slight drop when phase 1 passes the recording station (Figure 6, dashed line). The elongation of  
370 phase 1 can also be seen, where the dip in the DR is only ~1 min for RTMT, ~5 min for TRAN,  
371 and approximately 20 min for COLL. The reason the DR decreases during phase 1 for the 2007  
372 lahar could be due to the parallel component being more sensitive to flow processes than  
373 bedload forces (Barriere et al., 2015; Roth et al., 2016). During phase 1 discharge increases,  
374 sediment concentration is low, and streamflow dominates resulting in a low DR ([e.g. Doyle et](#)

375 [al., 2010](#)). The low DR can also be seen before the arrival of phase 1, due to streamflow already  
376 occurring in the channel. The higher flow parallel amplitude over cross-channel amplitude for  
377 streamflow has also been noted in the past for lahars at Volcán de Colima, Mexico (Walsh et al.,  
378 2020).

379 Following the low PSF phase 1 (i.e. front of the lahar), the peak seismic amplitude occurs. The  
380 peak seismic amplitude for RTMT is accompanied by an increase to higher PSFs > 30 Hz (Figure  
381 [3, 7a](#)). PSFs > 30 Hz have been shown in the past to be either dominated by turbulence or  
382 bedload transport (e.g. Gimbert et al., 2014; Roth et al., 2016). The 2007 lake-breakout lahar  
383 has been described as a hyperconcentrated streamflow (e.g. Procter et al., 2010b) with low  
384 sediment concentration, especially early on before the lake water captured enough material to  
385 bulk up into a full 4-phase lahar ([see section 1.1](#)). At RTMT, which was only 7.4 km from source,  
386 the lahar had not fully bulked up yet and was in a net depositional regime (Procter et al.,  
387 2010a). Due to the conditions of the lahar at RTMT, we surmise the higher PSF content for the  
388 peak seismic amplitude is dominated by turbulent-flow-induced noise. Furthermore, the higher  
389 PSF content at RTMT compared to TRAN and COLL (~30 Hz) could be due to the angle of the  
390 slope at the recording stations. Gimbert et al. (2014) noted that turbulence noise will dominate  
391 over bedload-induced noise on steeper slopes. Further down the channel at TRAN, the PSF for  
392 the peak seismic amplitude is ~30 Hz for all three components. Again, this high PSF can be  
393 attributed to turbulence as seen by the images taken at TRAN (Figure 2d). The difference at  
394 TRAN is the length of the higher PSF, where at RTMT the high PSF stays throughout the entirety  
395 of the recording window, at TRAN the high PSF and seismic amplitude only last for ~5 min. The  
396 difference at TRAN could be from the evolution of the lahar. By time the lahar reached the

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399 monitoring station at TRAN (28 km from source) the lahar was fully bulked up and had the  
400 properties of a traditional four phase lahar as described by Scott (1988) or Cronin et al. (1999).  
401 By time the lahar reached COLL 82 km from source, the peak seismic amplitude is associated  
402 with PSFs between 15-30 Hz, with bimodal patterns in the horizontal components and a tighter  
403 spread in the vertical component (~27-29 Hz). At COLL, the lahar had converted into a plug-like  
404 flow with lower turbulence and hence the higher PSFs are most likely associated with bedload  
405 transport (Figure 2f). Furthermore, Burtin et al. (2010) and Roth et al. (2016) noted that when  
406 the vertical component has greater seismic amplitudes than the horizontal components,  
407 bedload dominates. This same amplitude feature can be seen at COLL (Figure 5, bottom panel)  
408 where the vertical energy is greater than each of the horizontal components. The bimodal  
409 pattern of the horizontal components is likely to be the recording of both turbulence or flow  
410 properties (lower PSF) and bedload transport (higher PSF). This also explains why the vertical  
411 component does not show the same bimodal pattern. Barriere et al. (2015) described the  
412 parallel component as being more sensitive to flow properties, and Doyle et al. (2010) noted  
413 that the cross-channel component is likely dominated by turbulence, thus the reasoning behind  
414 the differing PSF patterns between components. This PSF feature is similar to the lahars  
415 recorded by Walsh et al. (2020), where the cross-channel PSF is confined within a narrow band  
416 around 15-20 Hz and the flow parallel PSF is more bimodal. At COLL, the cross-channel PSF is  
417 dominated by PSFs at ~18 Hz (lower than vertical component at ~28 Hz), with the flow parallel  
418 between 20-30 Hz.

419 The DR at the peak seismic amplitude for all three recording stations increases (Figure 6). The  
420 DR for both RTMT and TRAN increases to  $DR > 1$ . Doyle et al. (2010) noted that higher wetted

421 perimeters will increase the DR, which is true for the 18 March 2007 lake-breakout lahar (Figure  
422 6, peak DR/RMS amplitude). Conversely, the DR decreases after the peak seismic amplitude  
423 while the wetted perimeter is still high. Also, at COLL the DR only increases slightly with the  
424 seismic amplitude. While the wetted perimeter may be a factor in increasing cross-channel  
425 energy and thus the DR, the more likely explanation for the 18 March 2007 lahar might be the  
426 higher level of particle collisions at the peak seismic amplitude. More particle collisions would  
427 increase the DR (e.g. Doyle et al., 2010) due to more lateral excitation within the flow and  
428 against the channel walls. The increase in collisional energy also relates well with the PSF, as  
429 higher PSF correlates to an increase in the amount of interflow collisions as shown by Huang et  
430 al. (2004) and may also explain why DRs correlate well with PSF (Figure 8a). The DR for COLL  
431 during this same timeframe probably is not due to the amount of particle collisions due to the  
432 plug-like flow (Figure 2f), but rather the increase in sediment concentration (Figure 8c). As the  
433 sediment concentration increases at COLL the DR starts to increase as well (Figure 8b). Similar  
434 to Doyle et al. (2010), COLL yields a correlation between DR and sediment concentration (Figure  
435 8b), where higher DRs indicate higher concentrations of sediment contained in the flow. Lastly,  
436 as noted above, DRs may correlate with PSF or at least indicate differing processes taking place  
437 within the flow (Figure 8a). Lower PSF would produce lower DRs because low PSF are more  
438 sensitive to flow processes (hence higher parallel energy) whereas higher PSFs would produce  
439 higher DRs due to higher PSF being dominated by collisions and turbulence (higher cross-  
440 channel energy) (Figure 8a).

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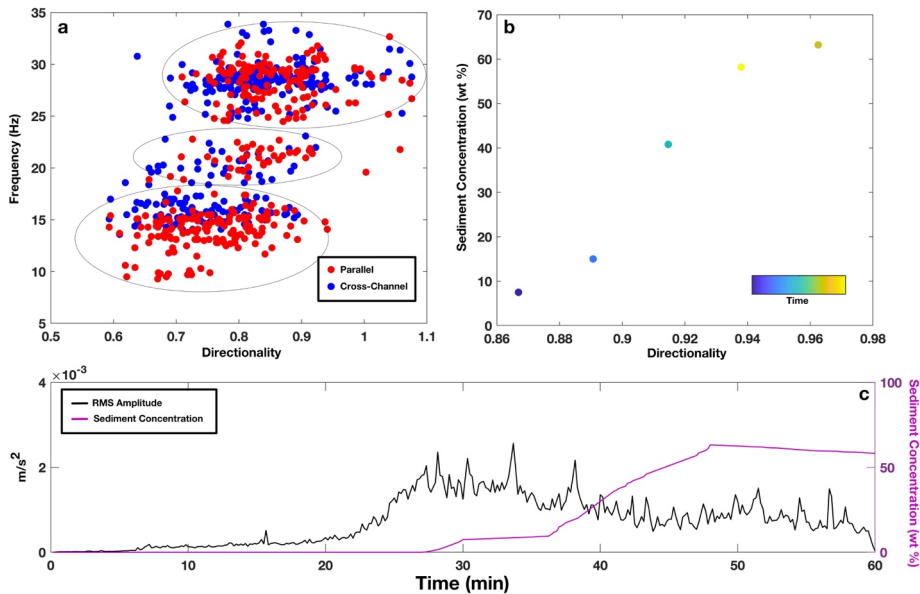
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448 *Figure 8. Plots of (a) correlation between PSF and DR at TRAN, (b) sediment concentration and DR at COLL, and (c)*  
 449 *seismic amplitude (black line) with sediment concentration (purple line) depicting the lag in sediment at COLL. Note*  
 450 *on (a) parallel (red dots) and cross-channel (blue dots) PSF display three different zones (black circles). Also note*  
 451 *that at COLL the first sediment concentration measurement did not occur until the 30 min mark.*

452 While the lahar at RTMT was a large outburst flood/sediment-laden flow, and at COLL a plug-  
 453 like flow, at TRAN the 18 March 2007 lahar was a dynamic bulked up “traditional” lahar. The  
 454 evidence for this is in the PSF content for TRAN (Figure 4) compared to the other two  
 455 monitoring sites. At TRAN the PSF has a step-up step-down pattern for the first 30 min of the  
 456 lahar passing, and then transitions to a bimodal or wide PSF range for the rest of the recording  
 457 window. As noted above, the low PSF preceding the lahar head arrival is thought to be due to a  
 458 sensitivity to water transport properties (Figure 2c). The increase to higher PSFs during the peak  
 459 seismic amplitude may be from particle collisions and/or higher turbulence (Figure 2d). After  
 460 the maximum seismic amplitude at TRAN, the PSF decreases to 10-20 Hz. This drop in PSF after

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462 the highest stage and amplitude could be from a more water transport dominated regime,  
463 which can be seen in the increased parallel amplitude (Figure 4, [8a](#)). The decrease may also be  
464 from greater frictional sliding on the channel bed (Huang et al., 2004). Furthermore, this PSF  
465 range could simply indicate a decrease in turbulence (Figure 2e). After the decrease to 10-20 Hz  
466 PSFs, the PSF displays a bimodal or wide frequency range at ~28 min (Figure 4, [7b](#)). As  
467 aforementioned for COLL, this PSF pattern could be from both bedload- and water-transport-  
468 induced noise. This timeframe is also where the peak sediment concentration would be, as  
469 noted by Cronin et al. (1999), and thus the PSF would show more bedload high PSF. This  
470 hypothesis also compares well with the DR (Figure 6b), where the cross-channel energy  
471 increases starting at ~25 min indicating that the sediment concentration may be increasing  
472 (Doyle et al., 2010). Finally, the wide PSF range later in the recording window (Figure 4) could  
473 also result from the lahar having two distinct layers as described by Cronin et al. (2000), where  
474 there is a wide more dilute finer grain top layer and a channelized sediment-rich layer on the  
475 bottom. The two layer model can apply to TRAN because the lahar at this monitoring station  
476 overtook the channel (Figure 2d,e) and proceeded to flow horizontally outward forming the  
477 surface layer described by Cronin et al. (2000).

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#### 478 **4.3 Implications for monitoring**

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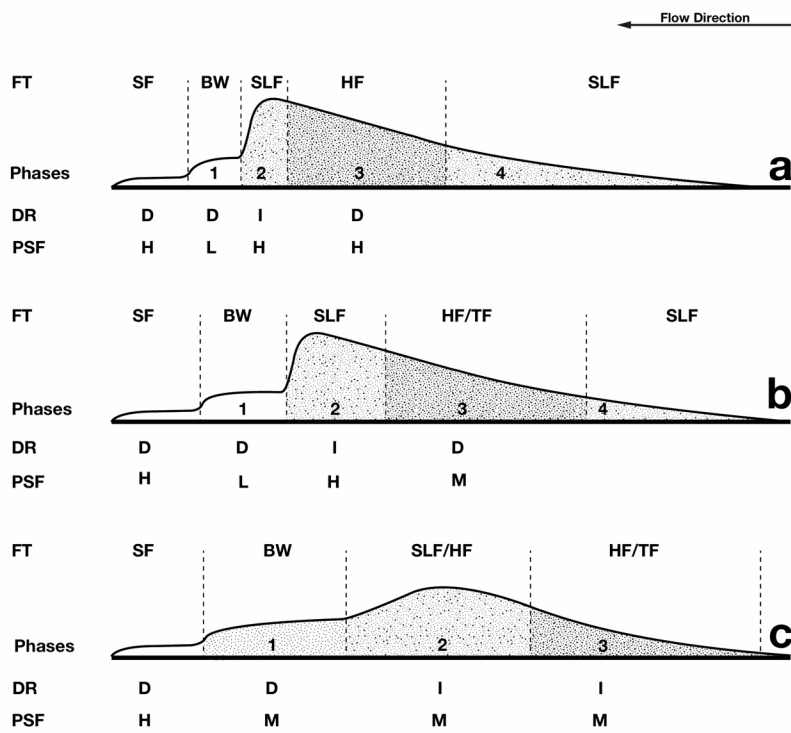
479 The main goal of this research is to contribute in defining better monitoring criteria for  
480 dangerous mass flow events. The data described above is part of a larger collection of  
481 monitoring data collected over the entire length of the Whangaehu channel consisting of 21  
482 monitoring sites and years of preparation (e.g. Manville and Cronin, 2007; Keys and Green,

485 2008). Due to this, the ability to accurately estimate the properties of the lahar at various  
486 stages along its path is possible. When it comes to flow events of any size, the ability to  
487 understand how the dynamics change with distance along the channel is important for warning  
488 and future hazard mitigation. We show here that a lake-breakout event can start out as an  
489 outburst flood, bulk up into a hyperconcentrated flow, then eventually elongate and entrap  
490 enough sediment to transform into a plug-like slurry flow. Each of these flow types yields  
491 differing PSF ranges and patterns due to the relationship between the channel geometry,  
492 sediment concentration, turbulence, and bedload transport. While the lahar at different  
493 stations along the channel may have differing PSF content, we also show that the lahar  
494 elongates and a predictable model (e.g. Cronin et al., 1999) can be used with and shown in the  
495 seismic data. Being able to apply such a model may yield some relevance of universality in  
496 terms of warning systems at different distances away from the mass flow source. Whereas  
497 shown above, the flow phases at each monitoring station can be seen, but at differing lengths  
498 and times in the seismic signal (e.g. Figure 6). To better visualize this concept, conceptual  
499 models based off of the Cronin et al., 1999 models are created for each of the three seismic  
500 stations for the 18 March 2007 lahar (Figure 9). In the conceptual models for the 2007 lahar,  
501 the aforementioned elongation of the frontal pulse or bow wave (phase 1) and head of the  
502 lahar (phase 2) is shown, along with the differences and similarities between the properties of  
503 the lahar at the three seismic monitoring sites.

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508 *Figure 9 Conceptual models for the 18 March 2007 lahar at each of the three monitoring stations along the*  
 509 *Whangaehu channel depicting flow type and the estimated seismic properties at each flow phase. a) RTMT 7.4 km*  
 510 *from source, b) TRAN 28 km from source, and c) COLL 83 km from source. Flow types (FT) are as followed;*  
 511 *streamflow (SF), bow wave streamflow (BW), hyperconcentrated flow (HF), Transitional flow (TF), and sediment-*  
 512 *laden streamflow (SLF). Note, decreased (D), increased (I), high (H), low (L), and mixed (M) are notations for*  
 513 *directionality ratios and peak spectral frequency estimates.*

514 Another implication for future warning is the implementation of 3-component sensors and the  
 515 use of DRs for channels that have streamflow. Walsh et al. (2020) showed for lahars flowing in  
 516 La Lumbre channel at Volcán de Colima that the DR for streamflow is <1 and then increases  
 517 when the lahar arrives. This same feature can be seen at each of the three monitoring sites for

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519 the 18 March 2007 event (Figure 6) indicating differing flow types will still show this DR pattern  
520 within the same flow and at other channels. To further show this, there were three natural non-  
521 lake-breakout eruption-based lahars that occurred in the Whangaehu channel in September  
522 2007 (for more details on the lahars see Cole et al., 2009; Kilgour et al., 2010) and recorded on  
523 the seismometer at RTMT. The DR for the September events starts with streamflow with a DR <  
524 1 and when the first lahar arrives the DR increases to >1 and as the lahar fully passes, the DR  
525 decreases to <1 again (Figure 10a). As the second lahar arrives at RTMT (Figure 10, second  
526 dashed line), the DR increases to >1 again. After the second lahar passes the DR decreases once  
527 again back below DR<1. Finally, as the third lahar arrives (Figure 10, third dashed line) the DR  
528 yet again increases above 1 for the entirety of the event.

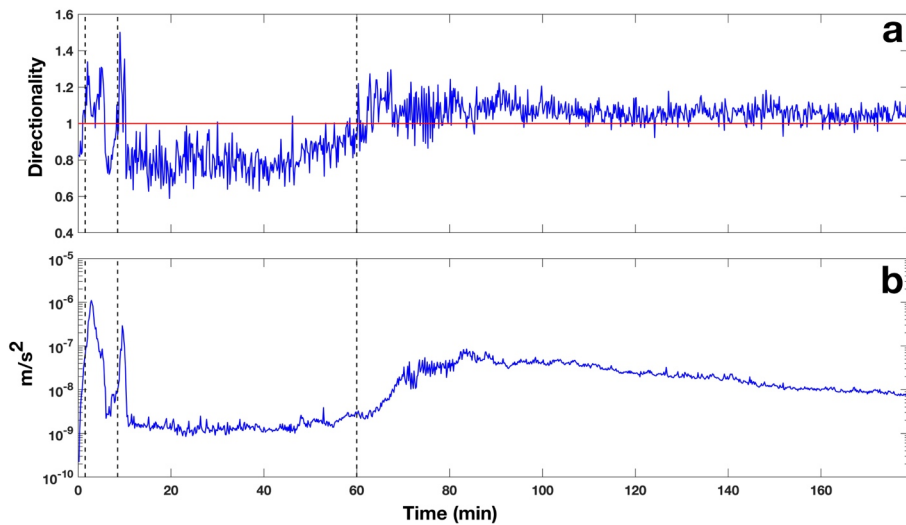
529 For many mass flows and especially those that flow into channels with preexisting streamflow,  
530 the peak seismic amplitude does not always coincide with the arrival of the mass flow, and thus  
531 may not be the most reliable for event detection or warning. These observations may be due to  
532 a frontal surge, the lag in sediment concentration or differences in peak amplitude with peak  
533 discharge. Phase 1 (frontal streamflow surge) of the model proposed by Cronin et al. (1999) was  
534 based on a hyperconcentrated flow interacting with streamflow, but has also been shown for  
535 debris flows as well (e.g. Arratano and Moia, 1999). Arratano and Moia (1999) showed at  
536 Moscardo Torrent, Italy, through a hydrograph that there was a precursory surge ahead of the  
537 debris flow that was not seen in the seismic record. Similarly, at Ruapehu, for the 18 March  
538 2007 lahar, at each of the three stations there is little evidence in the seismic amplitude that  
539 there was a precursory surge or phase 1 (Figures 3-5, bottom panel). Conversely, the surge  
540 ahead of the lahar can be seen in both the PSF analysis (drop to low frequencies) and in the DR

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544 (decrease in DR) right before the peak seismic amplitude arrives. This shows that when  
545 monitoring for future events that not only the amplitude should be used, but other analysis  
546 (e.g. PSF, DR) as well, otherwise there could be a delay in the detection of an event.



547

548 *Figure 10 (a) Directionality ratio for the time sequence of the three lahars that occurred on 25 September 2007. (b)*  
549 *RMS amplitude of the seismic record at RTMT during the timing of the three September lahars. Note that the black*  
550 *dashed lines represent the timing of each lahar arriving at the monitoring site.*

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551 Using all three components of the seismometer can be very beneficial in lahar monitoring. The  
552 above-mentioned DR analysis can only be completed with horizontal recording, and analyzing  
553 PSF in each component can yield critical information about the flow properties and dynamics.  
554 Examining the seismic amplitude differences can generate significant discoveries, for example,  
555 when the vertical component is stronger than the horizontal components, bedload dominates  
556 over turbulence noise (Burtin et al., 2010). Greater flow parallel signals may indicate higher  
557 water transport noises (Barrier et al., 2015) and higher cross-channel signals could be caused by

559 increased interflow particle collisions and flow-channel wall interactions (Doyle et al., 2010).  
560 While using the differences in each component can be useful, there are also some concerns.  
561 Channel geometry and bed conditions can alter the seismic signal (e.g. Coviello et al., 2019;  
562 Marchetti et al., 2019). Additionally, the flow parallel direction can be influenced by the lahar  
563 that has already passed, the lahar at the station and the lahar arriving. Furthermore, the tilt of  
564 the seismometer may play a large role in determining which component is larger (e.g. Anthony  
565 et al., 2018). In the case of the 18 March 2007 lahar a large pulse of water passed the stations  
566 which may explain why the parallel component is stronger than the other two components at  
567 RTMT and TRAN. At COLL, the lahar had elongated, lost energy, and thus shows more  
568 decreased flow parallel energy compared to the previous two stations. In the cross-channel  
569 direction, if a flow overtops the channel, the amplitude would presumably be dampened. This  
570 may be the case at TRAN where both the flow parallel and vertical directions are more  
571 energetic than the cross-channel amplitude after the passing of the head and breaking out of  
572 the channel occurred (Figures 2d, 4 bottom row). Another concern when using the horizontal  
573 components of a seismometer are the effects shallow layers may have on the site response of  
574 the sensor. This is especially true when a sensor is installed on soft or loose sediment (e.g. soil,  
575 fluvial/alluvial deposits). To test for potential effects by shallow layer fundamental frequencies,  
576 H/V analysis of ambient noise (streamflow dominant) was conducted (see Supplementary  
577 material). For RTMT, the H/V results depict a broad frequency peak between 5-15 Hz with a  
578 local maximum at ~8 Hz (Figure S4a). Comparing the H/V frequency with the PSF of RTMT  
579 (Figure 3), the only overlap is immediately before the front of the lahar passes the station  
580 where the PSF decreases for ~1 minute before the head of the lahar arrives. The H/V analysis

581 for TRAN has a multi- broad-peak shape, with frequency peaks at ~14 and ~28-35 Hz (Figure  
582 S4b). While these frequencies are similar to PSF values for TRAN (Figure 4), the H/V analysis has  
583 no distinguishable fundamental frequency, contains large error, and no frequency peak has a  
584 H/V amplification > 2. In order for a H/V frequency peak to be considered ideal, generally the  
585 amplification must to be greater than 2 and the standard deviation lower than a factor of 2  
586 (SESAME, 2004). The H/V amplification for COLL displays a broad frequency peak between 13-  
587 18 Hz, with a local maximum at ~18 Hz (Figure S4c). Comparing the PSFs at COLL (Figure 5), only  
588 the cross-channel direction has significant PSF values in the same frequency range (~18 Hz  
589 band). With all three stations not yielding distinct H/V frequency peaks, we surmise that the  
590 PSF content for the 18 March 2007 lake-breakout lahar is most likely dominated by the large  
591 flow passing by the seismic sensor rather than large site amplification effects from a shallow  
592 layer. While this may be the case, there is still the possibility that some of the PSF values could  
593 be due to local effects and should not be considered in the lahar analysis, e.g., the low PSFs at  
594 RTMT between 15-20 min (Figure 3), at TRAN contributing to some of the “jumping” in PSF  
595 content (Figure 4), or in the mostly dominant 15-20 Hz PSF in the cross-channel direction at  
596 COLL. Conversely, SCF values at each station do not reside in the broad H/V frequency range at  
597 any station (Figure 7), which may further support the hypothesis that almost all of the recorded  
598 frequencies are indeed produced by the lahar. With the use of horizontal components  
599 becoming common in mass flow monitoring, future 3-component analyses of mass flows should  
600 consider estimating H/V ratios or use other site response methods (e.g. spectral ratio analysis)  
601 in order to identify whether near-surface structures may affect the recorded flow data. Overall,  
602 all these concerns can and should be tested to estimate potential error in 3-component

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605 methods. Nevertheless, using all three components of the seismometer can enhance the  
606 productivity of warning systems, and if possible, should be used instead of single component  
607 sensors.

## 608 5. Conclusions

609 At 23:18 UTC on 18 March 2007 Mt. Ruapehu produced the biggest lahar in New Zealand in  
610 over 100 years causing  $1.3 \times 10^6$  m<sup>3</sup> of water to flow out of the Crater Lake and rush down the  
611 Whangaehu channel flowing for over 200 km to the Tasman sea. Seismic analysis at three  
612 monitoring locations along the channel (7.4, 28, and 83 km) yielded an understanding of how  
613 flow type and processes of the lahar evolve with distance. The proximal lahar was a highly  
614 turbulent outburst flood, which generated high PSF content in all three components. Further  
615 along the channel after the lahar had bulked up and transformed into a multi-phase  
616 hyperconcentrated flow, the PSF content was variable and showed changes in the flow  
617 regime/phase. Finally, at the most distal monitoring station, the lahar had lost energy and  
618 transformed into a slurry-type flow where the PSF content became more bedload-dominant.  
619 Additionally, directionality ratios from all three sites along with data from additional lahars  
620 yielded strong evidence that DRs can be used for warning systems when there is streamflow  
621 present in the channel. Furthermore, PSF and DRs show evidence of a pre-lahar water pulse  
622 that is concealed in the raw seismic data, but has been observed visually. Ultimately, the use of  
623 3-component broadband seismic analysis for the 18 March 2007 lahar at Mt. Ruapehu may lead  
624 to more accurate and advanced real-time warning systems for mass flows through the use of  
625 frequency and directionality around the world.

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628 *Author Contribution*

629 BW performed seismic analysis and drafted the manuscript, CL organized and prepared data,  
630 and JP created the visual location representation of the event. All participating authors  
631 contributed to the discussions and editing of the draft of the manuscript, as well as approving  
632 the final edition.

633 *Competing Interests*

634 The authors declare that they have no conflict of interest

635 *Acknowledgements*

636 This work was supported by the Resilience to Natures Challenges – New Zealand National  
637 Science, volcano program of research. We would also like to thank all the people from Massey  
638 University, Horizons Regional Council, NIWA, and the Department of Conservation that  
639 collected data and set up monitoring locations all along the channel in preparation for and  
640 during the lahar. A final special thanks to Kate Arentsen for editorial support.

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