

1 Long-term magnetic anomalies and its possible relationship to the latest greater 2 Chilean earthquakes in the context of the seismo-electromagnetic theory

3
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13 14 Abstract

15
16 Several magnetic measurements and theoretical development from different research groups have shown
17 certain relationships with worldwide geological processes. Secular variation of geomagnetic cutoff rigidity,
18 magnetic frequencies, or magnetic anomalies have been linked with spatial properties of active convergent
19 tectonic margin or earthquakes occurrences during recent years. These include the rise of similar
20 fundamental frequencies in the range of micro hertz before Maule 2010, Japan 2011, and Sumatra 2004
21 earthquakes and the dramatic rise of the cumulative number of magnetic anomalous peaks before several
22 earthquakes as Nepal 2015 and Mexico 2017, among others. Currently, all of these measurements have
23 been physically explained by the microcrack generation due uniaxial stress change in rock experiments.
24 The basic physics of these experiments have been used to describe the lithospheric behavior in the context
25 of the seismo-electromagnetic theory. Due to the dramatic increase in experimental evidence, physical
26 mechanism and theoretical framework, this paper analyses vertical magnetic behavior close to the latest
27 three main earthquakes in Chile: Maule 2010 (Mw8.8), Iquique 2014 (Mw8.2), and Illapel 2015 (Mw8.3).
28 The FFT, Wavelet transform and daily cumulative number of anomalies methods were used during quiet
29 space weather time during one year before and after each earthquake in order to filter space influence. FFT
30 Method confirms the rise of power spectral density in the mHz range before one month each earthquake,
31 which decreases to lower values after some months after earthquake occurrence. The cumulative anomalies
32 method exhibited an increase previous to each Chilean earthquake (50-90 days prior earthquakes) similar
33 to those found for Nepal 2015 and Mexico 2017. The wavelet analyses also show as similar properties as
34 FFT analysis. However, the lack of physics-based constraints in the wavelet analysis do not allow
35 conclusions as strong as FFT and cumulative methods. By using these results and previous research, it
36 could be stated that these magnetic features could give seismic information from impending events.
37 Additionally, these results could be related to the Lithosphere-Atmosphere-Ionosphere coupling (LAIC
38 effect) and the growth of microcracks and electrification in rocks described by the seismo-electromagnetic
39 theory.
40

1. Introduction

As earthquakes are geological events that might cause great destruction, studies about their preparation stage and generation mechanism are a matter of concern. That is why scientific studies offering new information, evidence or insights about different physical mechanisms or activation insights during the seismic cycle, improve our understanding of earthquakes occurrences. Currently, one of the most controversial physical mechanisms that is being studied is the lithospheric electromagnetic variations as earthquake's precursory signals. Nevertheless, the study of magnetic and geological relationships is not something new. For example, the decadal variations of the geomagnetic field have been associated with an irregular flow of the outer core (Prutkin, 2008). Thus, the secular variation of the magnetic field can be interpreted as the response of the movement of the fluid outer core interacting with the topography of the lower mantle. Then, as that topography in the core-mantle boundary corresponds to a projection of the topography of the earth's surface (Soldati et al., 2012), it was not surprising that Cordaro et al. (2018) and Cordaro et al. (2019) found significant variations of geomagnetic cutoff rigidity R_c at relevant geological places in the Chilean margin.

Regarding earthquakes, many attempts to determine the location, date and magnitude of seismic movements have been made in the past (e.g. Jordan et al., 2011), but these historical efforts have failed to conclude that it is possible to use seismological data as a predictive tool (Geller, 1997). Besides, when less classical methods (e.g., electromagnetic methods) have been used since some decades ago. First attempts on this topic can be found in the work of Varotsos - Alexopoulos - Nomicos (VAN) (see Varotsos et al., 1984, and the references therein). Techniques for seismic-electrical signals associated with the VAN method has been considerably improved and applied in several contexts (see Varotsos et al., 2019, and the references therein). Also, some debates about this theory can be found in the work of Hough (2010). Recently, electromagnetic methods have risen with relevant and concluding evidence. Specifically, it is because physical mechanism, based on the Zener-Stroh mechanism, links micro cracks to magnetic anomalies and fault's friction is currently available (e.g., Stroh, 1955, Slifkin, 1993, Venegas-Aravena et al., 2019, Venegas-Aravena et al., 2020), wide frameworks are being studies (holistic interaction between lithosphere, ionosphere and atmosphere, e.g. De Santis et al., 2019, Yu et al., 2021). Moreover, different electromagnetic theories related to earthquakes have been implemented. For example, De Santis et al. (2017) and De Santis et al. (2019b), showed the method of magnetic anomalies in which long-term magnetic data from different satellites (ionosphere level) were considered during quiet or no disturbed periods due to the space weather. After removing a known magnetospheric process from data as daily variation, the remaining magnetic perturbation or anomaly could be considered as lithospheric origin. This method allowed them to study magnetic measurements mostly free of external perturbation prior and after 16 worldwide earthquakes of magnitude approximately greater than Mw 6.5. When satellites covered areas close to each earthquake locations, they found an increase in the number of magnetic anomalies prior (1-3 months) to the occurrence of these earthquakes and a decrease after the earthquake (De Santis et al., 2017, Marchetti and Akhoondzadeh, 2018, Marchetti et al., 2019a, b, De Santis et al., 2019b).

Other methodologies also support certain statistical correlations to earthquake's preparation phase. For instance, the rise of magnetic signals characterized by a wide range of ultra-low frequencies (5-100 mHz and 5.68 – 3.51 μ Hz) or the ionospheric disturbs before several earthquakes have been widely and intensively reported during a couple of decades (Hayakawa and Molchanov, 2002, Pulnits and Boyarchuk, 2004, Varotsos, 2005, Balasis and Manda 2007, Foppiano et al., 2008, Molchanov and Hayakawa, 2008, Liu, 2009, Hayakawa et al., 2015, Contoyiannis et al., 2016, Potirakis et al., 2016, Villalobos et al., 2016, De Santis et al., 2017, Oikonomou et al., 2017, Cordaro et al., 2018, Marchetti and Akhoondzadeh, 2018, Potirakis et al., 2018, Ippolito, et al., 2020, Varotsos et al., 2019, Florios, et al., 2020, Pulnits et al., 2021, among others).

The magnetic phenomena not only have risen during decadal or preparation state, but also during the fast coseismic stage. For example, small magnetic variations (~ 0.8 nT) at ~ 100 km were measured during the Tohoku 2011 Mw9.0 earthquake (Utada et al., 2011). Similar findings were shown by Johnston et al. (2006) during the Parkfield 2004 Mw6.0 earthquake (~ 0.3 nT) at ~ 2.5 km. In addition, peaks of ~ 0.9 nT were measured at ~ 7 km during the Loma Prieta 1989 Mw7.1 earthquake (Fenoglio et al., 1995, Karakeliana et al., 2002). The above-mentioned reports have shown strong evidence of the presence of magnetic signals during the seismic preparation stage and during the rupture process itself. Up to this date, there are several experiments and theoretical models that identify and explain the physical mechanism of different magnetic variations related to geological properties (e.g. Freund, 2010, Scoville et al., 2015,

1 Yamanaka et al., 2016, Venegas-Aravena et al., 2019, Vogel et al. 2020, Yu et al., 2021). According to
2 experiments, the rise of electrical current flux within rocks is due the movement of imperfections and the
3 suddenly growth of microcracks when rock samples are being uniaxially stressed in the semi-brittle regime
4 (Anastasiadis et al., 2004, Stavrakas et al., 2004, Ma et al., 2011, Cartwright-Taylor et al., 2014, among
5 others). The applied external stress generates the internal collapse of rock, which imply the fast growth of
6 microcracks and the increase of electrical currents that flows throughout the crack right before the failure
7 of rock samples (e.g. Triantis et al., 2008, Pasiou and Triantis, 2017, Stavrakas et al., 2019). These currents
8 created by this mechanism are known as pressure stimulated currents (PSC) and its rise occurs mainly when
9 the rock samples abandon linearity (see Triantis et al. (2020) and references therein for further details). This
10 pre-failure indicator has been used as the experimental base for theoretical descriptions of impending
11 earthquakes at lithospheric scale (Tzanis and Vallianatos, 2002, Vallianatos and Tzanis, 2003, Venegas-
12 Aravena et al., 2019, Venegas-Aravena et al., 2020). This seismo-electromagnetic theory has explained the
13 frequency range, the cumulative number of anomalies, the coseismic signals, friction states at fault, and the
14 b-value time evolution by considering fast stress changes in the fault surrounding area. This area of fast
15 stress changes was theorized by Dobrovolsky et al. (1979) and it could cover thousands of kilometers.
16 Similarly, Venegas-Aravena et al. (2019) also found that the growth of microcracks and magnetic signals
17 are hosted by these stress conditions within this large area. Recently, large areas of fast stress and strain
18 changes, that surround the impending earthquakes, has been also confirmed by GPS analysis (Bedford et
19 al., 2020).

20
21 Despite the abovementioned evidences, still there are no reports of cumulative anomalies in one
22 the most actives margin: the Chilean margin (e.g., Vigny et al., 2011, Pedrera et al., 2014, Carvajal et al.,
23 2017, Zhang et al., 2017, Abad et al., 2020, Satake et al., 2020). In Figure 1 one can observe the historical
24 strong earthquakes across the Chilean margin. That is why this works present a wide study of magnetic
25 signals which include spectral (Fourier and Wavelet analysis), cumulative, and space weather analysis one
26 year before and after the latest three main megathrust earthquakes in Chile: Maule 2010 (Mw8.8), Iquique
27 2014 (Mw8.2), and Illapel 2015 (Mw8.3). The space weather and general magnetic conditions are found in
28 Section 2. The main magnetic and frequency analysis are defined and performed in Section 3. The relation
29 between results and physical mechanism from the seismo-electromagnetic theory is Section 4. Finally,
30 discussions and conclusions are in Section 5.

31 **2. Data processing consideration regarding the space weather and magnetic conditions**

32
33 In order to perform a clear interpretation of the results, any methods and data processing must
34 answer the classic questions: 1) What is actually being measured? 2) Where do the disturbances come from?
35 3) How should the disturbing data be removed? Here, the proper way to answer the above-mentioned
36 questions is by the recognition of the physical process, which generates external disturbances on
37 measurements. Then, the uses of the standard index convention that identify disturbed times, and statistical
38 analysis. This led us to implement four filters before working. These filters are:

39
40
41 1.- DST filter: Eliminate periods of high solar magnetic activity. That is, the data within these periods are
42 useless since the terrestrial cannot be distinguished from the solar.

43
44 2.- Daytime filter (or quiet time): Eliminate daytime data as they reflect the interaction of the solar wind
45 with the magnetosphere.

46
47 3.- Stochastic filter: Moving averages eliminate the low-frequency variations associated with the usual flow
48 of the ionosphere. Experiments and theory show that electrification of rocks prior to failure occurs mainly
49 in the millihertz range (e.g., Triantis et al., 2012). Thus, moving averages filter the lower frequencies by
50 using residuals methods to contain error propagation, that is, the difference between the signal and its
51 smoothed signal.

52
53 4.- Recurrence filter: This filter controls the failure of the other 3. Specifically, by means, the definition of
54 anomalous residual implies that any magnetic anomaly must be uncommon. Thus, the probability of finding
55 an anomalous residual within a given period should be tended to zero. In other words, few anomalies should
56 be measured in the period. This indicates that if the number of detected anomalies dramatically increases,
57 they are more common, and thus, not all of them could be considered anomalous. This contradiction could
58 arise in two scenarios: a) if a large number of anomalies occurs during the entire period, it means that the
59 threshold should increase. b) If a large number of anomalies occurs during a short period of time, let say

1 during one single day out of several months or years, then, it means that filters 1, 2 and 3 were insufficient
2 to filter that specific day which implies that day cannot be considered.
3

4 Finally, the resulting filtered variations are potentially generated in the lithosphere, not in the space
5 or ionospheric environment. It is important to note that the remaining data is almost unchanged since the
6 analysis studies the applicable periods. With this added to records of several years, we eliminate one of the
7 most significant scientific community's concerns: the origin of disturbances, propagation of errors, and false
8 positives. After this process, spectrograms or other methods can be used. That is, it requires a very
9 sophisticated preparation to discern and identify problematic disturbances in the records. As this is a
10 sophisticated filter process will be detailed in the following Sections.
11

12 **2.1 External magnetic disturbances**

13
14
15 Before going to the study of the magnetic field and its temporal variations, it should be
16 remembered that the rate of change of the magnetic field is influenced by the rate of variation of the spatial
17 particle count. These are different cases of irregular and regular phenomena of the nearby space climate.
18 Regular magnetic variation creates periodic fluctuations in the interplanetary magnetic field in a wide range
19 of periods, from few day periods up to seasonal variations (Moldwin, 2008, Blagoveshchensky et al., 2018,
20 Yeeram, 2019). Irregular variations occur when sudden increases of incoming solar particles are recorded
21 across the geomagnetic field. This particle disturbance induces a 10% to 20% decrease in magnetic field
22 intensity because of the change in pressure that extraterrestrial particles exert on the magnetosphere, an
23 effect that can last from a couple of hours to several days (Russel et al., 1999). One explication is that
24 particles following the magnetic field lines, in the turbulent magnetic reconnection that is present in the
25 diurnal variation and the regular variations (Priest and Forbes, 2000, Kulsrud, 2004, Cordaro, et al., 2016,
26 Lazarian et al., 2020). Other minor irregular magnetic fields as auroral events and electric current in the
27 ionosphere are not considered for this paper (see Diego et al. (2005) for detailed description for these
28 phenomena).
29

30 Some indices are used in order to measure the space disturbances and its manifestation in the
31 geomagnetic field. For example, the Kp index measure the influence of geomagnetic storms in the
32 horizontal magnetic field (Dieminger et al., 1996), while Dst index is interpreted as a measure of the
33 magnetospheric ring-current strength which is proportional to the particle's kinetic energy (e.g. Silva et al.,
34 2017). Usually, Dst index could increase dozens or hundreds of nT during magnetic storms ($K_p > 4$), that
35 is why it is important to incorporate these indices to create reliable magnetic models.
36

37 **2.2 Secular variation in the Chilean convergent margin**

38
39 The magnetic response to these disturbances requires a reference model that allows to discriminate
40 earth's magnetic features from disturbs that spreads throughout interplanetary magnetic fields. One of those
41 features corresponds to the magnetic shielding against incoming turbulent particles which is known as
42 geomagnetic cutoff rigidity R_c (Pomerantz, 1971). The rigidity R_c is defined as the product of the force of
43 the magnetic field and the curvature radius of the incident particle r_g and it can be estimated globally by
44 using the Tsyganenko magnetic field model (for details see: Smart et al., 2000, Smart and Shea, 2001,
45 Tsyganenko, 2002a, 2002b). The R_c variations describes geomagnetic secular variations which could be
46 related to geological features in the Chilean margin (Pomerantz, 1971, Shea and Smart, 2001, Smart and
47 Shea, 2005, Herbst et al., 2013, Cordaro et al., 2018, Cordaro et al., 2019). For example, regarding to
48 latitudinal effect (Pomerantz, 1971), Cordaro et al. (2019) found that the highest variation rate of effective
49 R_c values were obtained at 46.5°S, 76°W and at and at 52°S, 76.5°W (Figure 1). The first one is in the
50 Taitao Peninsula, Chile which corresponds to the triple junction point of three tectonic plates: Nazca, South
51 America, and Antarctica. The second one is close to Puerto Natales in the Strait of Magellan area, also a
52 triple junction point of three tectonic plates: South America, Antarctica and Scotia (Figure 1). There are
53 other geological and geomagnetic links as the flat slab in the Chilean convergent margin (Cordaro et al.
54 2018, Cordaro et al. 2019). However, these results are not surprising because changes in R_c represent
55 secular variations that represent magnetic secular variations created at the outer core (Bloxham et al., 2002,
56 McFadden and Merrill, 2007, Sarson, 2007, Finlay, 2007, Herbst et al., 2013). Specifically, 3D models of
57 core mantle boundary (CMB) topology based on the velocities of seismic waves (Simmons et al., 2010)
58 show the existence of positive topography in upthrust regions and negative topography in subduction zones
59 (Yoshida, 2008, Lassak et al., 2010, Soldati et al., 2012). Let us remark that the intensity of the geomagnetic

1 field at within the outer core is estimated to be of the order of 2-4 mT (rms) (Olson et al., 1999, Olson,
2 2015), while at earth surface varies between 20,000 and 60,000 x 10⁻⁹ T.
3

4 The most relevant magnetic feature in the Chilean sector is the low magnetic intensity values that
5 correspond to the influence of the South Atlantic Magnetic Anomaly (SAMA) (e.g. Cordaro et al., 2016).
6 Recently, Tarduno et al. (2015), argued that SAMA is being created by a topography structure in the CMB
7 beneath south Africa. SAMA is not only linked with global magnetic features as a geomagnetic dipole
8 moment (e.g. Heirtzler, 2002, Gubbins et al., 2006), it is also corresponding to the closer area between
9 earth's surfaces and radiation belt. This proximity allows more charged particles and more disturbances in
10 the magnetic field near the Chilean margin (e.g. Kivelson and Russell, 1995). That is why a proper magnetic
11 response to external disturbances is required before and after earthquakes occurrences.
12

13 **2.3 Magnetic perturbation during seismic events of 27/2/2010 in Maule, 1/4/2014 in Iquique and** 14 **16/9/2015 in Illapel** 15

16 The manifestation of space climate in the geomagnetic field during the periods concerned is
17 defined by the Kp magnetic activity index as shown in Figure 2 for the months previous to the three
18 earthquakes: Maule 2010 (Dec 12, 2009 to Mar 15, 2010), Iquique 2014 (Jan 1, 2014 to Apr 15, 2014) and
19 Illapel 2015 (Jul 1, 2015 to Sep 30, 2015). For Maule 2010 the magnetic activity reached a Kp index equal
20 to or greater than 4 on only three isolated occasions, it is therefore considered a calm period; for Iquique
21 2014, activity was concentrated around Feb 19, 2014 while for Illapel 2015 the maximum activity was
22 recorded between September 8 and 10. In all three cases, activity did not persist in time. In fact, according
23 to figure 2, there is no evidence of an increase in the amount of external magnetic perturbations prior each
24 earthquake.
25

26 **3 Main magnetic evolution and frequency analysis** 27

28 Magnetic measurements and analysis are carried out in this section. The main aim of this section
29 is to use different magnetic methodologies and figure out which of them seems more earthquake-related.
30 The stages correspond to the long-term magnetic evolution, the simple frequency analysis, wavelet and
31 anomaly analysis. Stations used here are Putre (PUT), Easter Island (IPM, also known as Isla de Pascua),
32 Los Cerrillos (CER), Pilar (PIL), Osorno (OSO) and Laboratorio antártico de radiación cósmica (LARC).
33 See Figure 1 for their location and information of PUT, CER and LARC in Table 1. In the case of PUT and
34 IPM, the Dobrovolsky area and the earthquake distances will be used in the following subsections (Table
35 2).
36

37 **3.1 Long-term magnetic records** 38 39

40 A high correlation between the vertical component of the earth's magnetic field and seismic
41 activity at the Putre station was found (Cordaro et al., 2018). That is why we seek to specify this behavior
42 in a shorter time window than the period studied previously (1975-2010). In addition, the B_z component in
43 Easter Island (IPM) station is also used because it has not been thoroughly investigated (Note that the IPM
44 station was closed in 1968 and subsequently reactivated in 2008 by the French INTERMAGNET Group
45 and the Meteorological Service of Chile) (Chulliat et al., 2009, Soloviev et al., 2012). The Putre observatory
46 is at 18°11'47.8S, 69°33'10.9W, 3,598 m.a.s.l. (meters above sea level); and it is located on the western
47 edge of the South American Plate. This zone includes the South Atlantic Magnetic Anomaly (SAMA), the
48 center of which is 1,700 kilometers east of this observatory. The measurements confirm low Bz values at
49 the station Putre. The instrument error of the geomagnetic measurements is of the order of 5 nT (Cordaro
50 et al 2012). IPM is located at 27.1°S, 109.2W, 82,83 m.a.s.l, on the western edge of the Nazca plate,
51 characterized as a hotspot (e.g. Vezzoli and Acocella, 2009). OSO is located in the coordinates 40°20'24"
52 S, 74°46'64" W and PIL at 31°40'00.0" S, 63°53'00.0" W (Figure 1).
53

54 In the Putre, a diminution in the values of the whole magnetic field and each of its components is
55 found. This can be attributed to the fact that Putre observatory is influenced by the South Atlantic Magnetic
56 Anomaly, while on Easter Island the influence of SAMA is weaker (Storini et al., 1999). These magnetic
57 influences are also found in Los Cerrillos observatory. The scientific and technical characteristics of the
58 Putre (PUT) and Los Cerrillos observatories, i.e. location, altitude, atmospheric depth, type of detectors,
59 geomagnetic cutoff rigidities and operating times, may be found in Refs. (Cordaro et al., 2012, Cordaro et

1 al., 2016) while for Easter Island (IPM) the information is available in SuperMag Network (Chulliat et al.,
2 2009, Gjerloev, 2012). The main characteristics for the observatories i.e., location, altitude, atmospheric
3 depth, type of detector, and operations time, are shown in Table 1.
4

5 Measurements of the B_z component are represented in Figure 3. We observe that similar gradients
6 in Iquique 2014 and Illapel 2015 to those found in Maule 2010, giving rise to a jump in each case. It is
7 known that these magnetic signals are generated by the earth's core and disseminated through the mantle,
8 implying changes in its electrical conductivity (Stewart et al., 1995).
9

10 The jump in the B_z component for Maule 2010 was recorded in the Putre station on Jan 23, 2010
11 (purple solid line in Figure 3a), a time lapse of 36 days before the earthquake (solid red line) and the moment
12 at which a change appears in the gradient or trend. It alters from a diminution of 225 nT in the period Oct
13 31, 2009 to Jan 23, 2010, to a less abrupt diminution of 30 nT between Jan 23, 2010 and Apr 3, 2010; prior
14 to the jump on Jan 16, 2010 there is a small, abrupt diminution from -5048 nT to -4927 nT. Discounting
15 this small, abrupt diminution, the delta between the gradients falls from -4960 nT to -4926 nT, delta = 34
16 nT as it is shown in Figure 3a.
17

18 For Iquique 2014 the jump recorded in Putre (Figure 3b) occurred on Dec 27, 2013 (purple solid
19 line), a time lapse of 96 days before the earthquake (red solid line). A change appears in the gradient on
20 this date from a diminution of 123 nT in the period Nov 14, 2013 to Dec 27, 2013, to a diminution of 113
21 nT between Dec 27, 2013 and Apr 15, 2014; the jump presents a change from -7355 nT to -7235 nT, delta
22 = 120 nT as it shown in Figure 3. For Iquique 2014 the jump measured at IPM occurred on Dec 31, 2013,
23 a time lapse of 91 days before the earthquake (Figure 3c). The trend shows a slight increase between Sep
24 30, 2013 and Jan 3, 2014, from -19116 nT to -19104 nT, while a further slight increase occurs in the period
25 Jan 3, 2014 to May 6, 2014, from -19101 nT to -19099 nT. Note that the size of the jump was -3 nT as it is
26 shown in Figure 4. For Illapel 2015 the jump measured at IPM occurred on Aug 31, 2015, a time lapse of
27 16 days before the earthquake. The trend shows a slight diminution between Aug 31, 2015 and Sep 20,
28 2015, from -19054 nT to -19072 nT, a jump of -11 nT, as one can observe in Figure 3d.
29
30

31 3.2 Simple Fourier analysis

32
33 Regarding the frequency analysis, the frequency spectrum values were analyzed for the Maule,
34 Iquique and Illapel earthquakes using the second derivative of the vertical component at PUT and IPM
35 stations. Fundamental frequencies before these earthquakes ranged from 5.606 to 3.481 μ Hertz or from 1
36 cycle / 48.9 hours to 1 cycle / 79.13 hours (Figure 4a). The increase in one or a group of frequencies reflects
37 the oscillations of the radial magnetic field whose oscillation period takes from ~2 to ~4 days. Specifically,
38 in the Maule event, peaks for the frequencies 4.747; 5.064 and 5.154 μ Hz were recorded (blue squares in
39 Figure 4a). In Iquique peaks of 4.611; 4.882 and 5.154 μ Hz were recorded (black dots in Figure 4a), and
40 for Illapel, 3.739; 4.630 and 5.520 μ Hz (red rhombuses in Figure 4a).
41

42 In order to identify a temporal domain where these frequencies arise, FFT is applied each 20 days
43 as a first approximation (Figure 4b, c). Before the Iquique 2014 event a jump in intensity was observed
44 associated with the frequency of 5.154 μ Hz for the period Dec 27, 2013 to Jan 11, 2014, i.e. after the jump
45 (Figure 3b, Figure 4b). Similar frequencies (3.739 μ Hz) rise during Sep 1, 2015 to Sep 8, 2015 before the
46 Illapel 2015 event (Figure 4c). These findings imply a more detailed methodology in order to study the
47 origin of these frequencies.
48

49 3.3 Wavelet analysis

50
51 We have used the wavelet transformation to analyze localized versions of power within a
52 geomagnetic time series. In this way, it can break down a time series into the time-frequency space,
53 determine the dominant modes of variability and how they vary over time (Torrence and Compo, 1998).
54 Here, the goal is to look for the rare variations that could not be attributed to space weather in the daily
55 average measurements. According to Cordaro et al. (2018), the magnetic field's vertical component showed
56 variations related to the Maule 2010 earthquake. That is why values of the vertical component of the
57 geomagnetic measurements at the OSO station were considered. Note that the OSO station is the closest
58 station to the main earthquake. In order to avoid the space weather influence, the highest variations were
59 not considered. One way to consider these two restrictions is by using statistical analysis. For example, a
60 lower and upper threshold could be defined by using the standard deviation. Consider higher magnetic

1 peaks, but not too meaningful because they could be related to space weather conditions. An example of
 2 this statistical analysis when an upper threshold of 2 standard deviations is used can be found in Figure 5,
 3 in which panel (a), (b), and (c) stand for Maule 2010, Iquique 2014, and Illapel 2015 earthquakes,
 4 respectively. For Maule 2010, the spectral analysis shows a dramatic increase 30 days before Maule
 5 earthquake and a decrease ten days after the earthquake occurrence (yellow and green arrows in Figure 5a).
 6 The frequencies that rise comprise a range close to 3-5 μ Hz. Note that no other significant increase is seen
 7 during the two years of measurements between days -365 up to -30 and between 10 to 365 is clear that there
 8 is no significant rise of frequencies (blue shades in Figure 5a). Panel (b) of Figure 5 shows the results for
 9 Iquique 2014, which is characterized by two peaks. The first one rises 89 days before the earthquake (yellow
 10 arrow in Figure 5b), while the second one occurs after the Iquique earthquake (after the red line, which
 11 indicates the earthquake day). The Illapel case is also characterized by two peaks, as it is shown in panel
 12 (c) of Figure 5. The first rise occurs \sim 159 days before the earthquake (yellow arrow in Figure 5c), while
 13 the second rise is 52 days before the main earthquake (grey arrow in Figure 5c). Despite this promising
 14 methodology that considers the daily average and the upper threshold, an improved implementation of
 15 physical (i.e., space weather conditions) and statistical (adequate definition of anomalies and frequency
 16 considerations) analysis is required.

17
 18 Finally, let us point out that more profoundly and sophisticated multiresolution wavelet analysis
 19 on time series related to earthquakes has been performed by Telesca et al. (Telesca et al. 2004 & Telesca et
 20 al. 2007). This kind of study will be considered in future works.

21 22 3.4 Anomaly analysis

23
 24 In order to identify and discriminate external variations from those that could be considered as
 25 lithospheric (variations with lithospheric origin), this subsection handles the definitions of anomalous
 26 variations. This definition will be obtained considering the external perturbation by using the Dst index
 27 (<http://wdc.kugi.kyoto-u.ac.jp>). Then, spectral analysis will be performed. Additionally, the data used in
 28 this subsection is standard and comes from the supermag network (<http://supermag.jhuapl.edu/>). The data
 29 has a sampling frequency of one data per minute, and a period of one year before and one year after each
 30 earthquake was chosen.

31 32 3.4.1 Magnetic threshold definition

33
 34 In the method of cumulative magnetic anomaly on the surface of earth, we used statistically an
 35 atypical or anomalous value, that is, data that is quite far from the average values of the sample. So, we
 36 compare real values of B_i with a more representative value of the sample, its average \bar{B}_i . We will call the
 37 difference between the two as the magnetic residual ΔB_i . By using the distribution of data, we can define
 38 when a value is atypical or anomalous in a normal distribution by statistical definitions of quartiles and
 39 outliers.

40
 41 On the one hand, we create a filter that eliminates the frequencies averaged near Nyquist and
 42 establishes a filter that eliminates high frequencies (stochastic filter). The option was to consider a weighted
 43 moving average of five points: $\bar{B}_i = aB_{i-2} + bB_{i-1} + cB_i + bB_{i+1} + aB_{i+2}$. Here, other researchers have
 44 used cubic splines instead of moving average (e.g. De Santis et al., 2017). In our case we use $a = 0.07$, $b =$
 45 0.25 and $c = 0.5 - 2a$. The uncertainty of the Flux-gate magnetometers (supermag) from OSO and PIL
 46 station is $\delta B_i = \pm 0.1 nT$, while the error of the moving average implementation is $\delta \bar{B}_i = \pm 0.1 nT$. As
 47 residual values are defined as the difference between real and smoothed data ($\Delta B_i = B_i - \bar{B}_i$), the total
 48 error propagation of residual is $\delta B_i + \delta \bar{B}_i = \pm 0.2 nT$.

49
 50 Let us comment that the error of propagation is used to define a threshold that determines when a
 51 residual is considered anomalous or not. For instance, in statistics 0.6745σ represents 50% of the data that
 52 is closer to the average (where σ is the standard deviation). This means that residuals less than $0.6745\sigma nT$
 53 are closer to the average and therefore are more common. As residuals are considered anomalous when
 54 they are unlikely, anomalous data should be defined as those residuals larger than $0.6745\sigma nT$. By adding
 55 the error propagation as condition ($0.6745\sigma + 0.2 nT$), and considering that the standard deviation is
 56 similar to $\sigma \sim 0.1 nT$, the percentage of residuals that meets this condition are considerably smaller than
 57 the 50% of the data (less common). Thus, residuals ΔB_i are consider anomalous (ΔB_{ai}) when

$$58 \quad | \Delta B_i | \geq 0.6745\sigma + 0.2 nT \quad (1).$$

1
2 The vertical magnetic thresholds found are 0.2246995 nT at OSO (Feb 27, 2009 - Feb 27, 2011), 0.2362868
3 nT at PIL (Apr 01, 2013 – Apr 01, 2015) and 0.2352825 nT at PIL (Sep 16, 2014 – Sep 16, 2016). These
4 threshold are ~ 6 , ~ 4.5 and ~ 4.5 times larger than each σ respectively. This means that each anomaly above
5 this threshold meets the 3σ criteria (a valid observation). Furthermore, the thresholds are close to the 5σ
6 criterion which corresponds to the standard discovery criteria in the physical sciences, for example this
7 criterion was used in the discovery of the Higgs boson ([https://home.cern/news/news/physics/higgs-within-](https://home.cern/news/news/physics/higgs-within-reach)
8 reach)
9

10 Regarding the external contribution, the data considered are for quiet periods $Dst < 10$ nT, and only
11 quiet magnetic data between 16:00 to 05:00 local time (Hitchmn et al., 1998). Some researchers who have
12 used satellites consider only the time periods in which the DST index is less than or equal to 20 nT (e.g.
13 Marchetti and Akhoondzadeh, 2018), or equal to 10 nT (e.g. De Santis et al., 2017). That means that the
14 space weather conditions could invalidate the anomaly condition defined in equation (1) if $DST > 10$ nT.
15 Then, the proper application of equation (1) is linked to those times where space weather activity is low.
16 That is when $Dst < 10$ nT and quiet magnetic data (16:00 to 05:00 local time).
17

18 Finally, the fourth filter process considered corresponds to the study of the recurrence of
19 anomalous residuals. For instance, if the anomalous threshold is three times the standard deviation, it
20 implies that one of every 740 data is anomalous. It is the same as two every 1480. As this work uses 1440
21 data per day, it is expected that \sim two will be anomalous per day. In terms of probabilities, it is the same to
22 $P_{day} \approx 0.0014$. Contrariety, if one single day shows, let's say, 30 anomalies, the probability is $P_{30} =$
23 $(P_{day})^{30} \sim 10^{-86}$. This means that the occurrence of 30 anomalies is virtually zero, and it could imply that
24 the previous filters (anomaly definition, DST, and quiet time) failed during that day. Then, it is not possible
25 to consider those days where the number of anomalies per day is considerably larger than 2. That is why
26 days with larger than ten anomalies are not considered valid.
27
28
29

30 3.4.2 Spectrogram

31
32 The filtered data correspond to a strong candidacy of lithospheric magnetic origin. This means that
33 any spectral analysis could reveal lithospheric variations. That is why simple spectrograms analysis is
34 performed. The spectrogram corresponds to the application of the moving Fourier transform. Here, the
35 temporal window size is 1 month with a 50% of overlap, which holds a reasonable spectral and time
36 resolution (see Rabiner and Schafer (1980) and Oppenheim et al. (1999) for spectrogram theory and
37 application). The OSO and PIL spectrograms for Maule 2010, Iquique 2014 and Illapel are shown in Figure
38 6.
39

40 In the Maule 2010 event, the spectrogram of the vertical magnetic component at the OSO station
41 is shown in Figure 6a. There is no significant rise of frequencies during the period before the Maule event
42 (before \sim January 10, 2010). Nevertheless, a dramatic increase during the period January 10, 2010 - May
43 02, 2010, occurs. Specifically, the rise of frequencies lies in the range $\sim 1 - 2.2$ mHz. The onset of this rise
44 (January 10) occurs more than one month before the Maule earthquake (February 27) and lasts almost four
45 months. That means that the spectral density reduces their activity after ~ 2 months of the earthquake.
46

47 The spectrogram for Iquique 2014 is characterized by two peaks (Figure 6b). The first one
48 corresponds to September 22, 2013, and the second one in \sim March 08, 2014. Here, the frequency range
49 comprises between ~ 1.2 to 2.7 mHz, which is similar to that found in the Maule spectrogram. However,
50 the main peak occurred during March, which is characterized by a dominant frequency close to 2.5 mHz,
51 while the dominant frequencies in Maule 2010 are close to 1.2 and 2.2 mHz, respectively. There is an
52 additional difference compared to the Maule event: The rise of frequencies in Iquique 2014 comprises a
53 significant decrease (or "valley") in the spectral density, which last almost two months (\sim November 09,
54 2013-Dec 27, 2013). It means that the second rise of frequencies lasts more than four months (\sim December
55 27, 2013-May 03, 2014), which is a similar duration compared to the frequency rise of Maule 2010.
56

57 The final spectrogram is found in Figure 6c, which corresponds to the Illapel 2015 event. Here, it
58 can be seen that almost the entire period was characterized by close to zero frequency variations.
59 Nevertheless, the frequency rise is similar to that obtained in Maule 2010. That is, significant frequencies

1 rise only on dates close to the earthquake event. The rise lasts almost three months (~ August 06, 2015 -
2 October 27, 2015). It is important to note that the gap between September 2015 is due to the strong spatial
3 weather activity. Despite this, it is clear that the earthquake occurrence lies during periods of high-frequency
4 activity, which is a similar feature compared to Maule 2010 and Iquique 2014.
5

6 Panels (a)-(c) of Figures 6 show that three strong earthquakes (Maule 2010, Iquique 2015, and Illapel 2015,
7 respectively) occurred during the rise of ultra-low frequencies of the vertical magnetic component. It is
8 important to highlight that these frequencies (mainly 1-2.5 mHz) vanish or reduce their intensity values
9 during other time periods. This is in agreement with other authors who have claimed that ultra-low
10 frequencies accompanied and the increases of the number of magnetic anomalies are related to the
11 earthquake preparation process (e.g., De Santis et al., 2017). It means that anomalous peaks produce the
12 magnetic oscillations in the magnetic records. By following the Venegas-Aravena et al. (2010) findings,
13 the number of these peaks also should increase (decrease) in time before (after) earthquake occurrences.
14
15
16

17 **3.4.3 Cumulative daily anomalies**

18
19 By following the anomaly definition (subsection 3.4.1), it is possible to find out the daily number of
20 anomalies- For example, in Figure 7 is shown the case of OSO station. Black dots follow a stable linear
21 increase in the number of cumulative anomalies (red line). Nevertheless, this tendency breaks close to Jan
22 11-12, 2010. From that day up to the first week of April, the numbers of anomalies experience a dramatic
23 increase. In the middle of this increase, the Maule earthquake hit (Feb 27, 2010). By subtracting the initial
24 linear tendency and comparing it to PIL station (Iquique 2014 and Illapel 2015), the sigmoidal feature is
25 clearer (Figure 8). The anomalies start to increase prior to each earthquake. For example, this increase
26 started ~47 days before the Maule 2010 earthquake, ~90 days before the Iquique 2014 earthquake and ~60
27 days before the Illapel 2015 earthquake (Figure 8).
28

29 Other researchers have used very different implementations, definitions, methodologies, and data
30 in order to find out these anomalies. For example, Marchetti and Akhoondzadeh, (2018) have also found a
31 sigmoidal signature in the anomalies of the Y components recorded by different satellites for the Mexico
32 2017 earthquake. In order to compare Mexico 2017 with Maule 2010, Iquique 2014 and Illapel 2015, the
33 initial linear trend has been removed (Figure 9). The initial onset of anomalies increased close to 60 day
34 prior to the Mexico 2017 earthquake. Note that in the four cases, the sigmoidal features are almost the same:
35 a linear stable number of anomalies characterize the initial period. Then, a dramatic increase in the number
36 of daily anomalies is followed by the main earthquake. This time is different in each earthquake but it lies
37 between 50 – 90 days after the initial anomalies increase. After the seismic events happen, the cumulative
38 numbers behave not similarly. For example, in the Mexico 2017 earthquake, the anomalies remain stable,
39 while in the Maule 2010 still is increasing in a less dramatic manner. At the end of the OSO measurements,
40 several anomalies appear, but it is not clear that these events could be related to other seismic events. In
41 order to understand the physics that lies in these events, a theoretical mechanism is required.
42
43

44 **4.-Magnetic anomalies and fracture mechanics by considering the seismo-electromagnetic theory.**

45
46 The frequency analysis (Figure 5, 6) and cumulative number of magnetic anomalies (Figures 7, 8,
47 9) shows an increase (spectral intensity and anomalies number) before each earthquake occurs. In the
48 anomalies case, a clear sigmoidal feature rises in Maule, Iquique and Illapel, which is a similar behavior
49 recorded in the Mexico earthquake (Figure 9, Marchetti and Akhoondzadeh, (2018)). This indicates that
50 anomalies behavior could correspond to a lithospheric origin. Currently, it has been shown that the origin
51 of these anomalies is associated with the cracking (or micro-cracking) of the semi-fragile-ductile part of
52 the lithosphere (crust) due to changes in stress (Venegas-Aravena et al., 2019). Typically, strain appears
53 when solids undergo loads or stress accumulation. However, micro-cracks rise specifically when solids do
54 not hold more deformation and prior to the main failure (e.g., Stavarakas et al., 2019, Li et al. 2020).
55 Experimentally, it has been shown that these conditions break the electrical neutrality within materials and
56 generate an electrical flux through rocks in a process known as pressure stimulated currents or PSC (e.g.,
57 Anastasiadis et al., 2004). Furthermore, it has been shown that PSCs can explain that the fractal nature of
58 cracks is sufficient to generate the frequency spectrum, co-seismic variations, the generation and behavior
59 of anomalies, and variation in the ionosphere in a theory known as seismo-electromagnetic theory
60 (Venegas-Aravena et al., 2019). Regarding the time evolution of magnetic anomalies, De Santis et al.

1 (2011) have shown that the sigmoidal shape is due to a manifestation of the stress changes when it is
2 reaching a critical point. Nowadays, theoretical development, geodynamical measurements, and
3 experimental studies have shown that the sigmoidal shape appears as a consequence of the dramatic increase
4 in the number of micro cracks (at a depth of few tens of kilometers) prior the main earthquake ruptures (De
5 Santis et al., 2015, Stavrakas et al., 2019, Venegas-Aravena, et al., 2019).

6
7 A schematic representation of the crack generation in the geodynamical context can be seen in
8 figure 10. At the initial time $t = t_0$ the intact lithosphere undergoes a uniaxial non-constant stress σ (Figure
9 10a). Then the first signs of micro cracks appear at $t = t_1$ due the increase of the stress (Figure 10b). When
10 the lithosphere can not hold more deformation, a dramatic increase in the crack generation appear
11 throughout the lithosphere ($t = t_2$ in Figure 10c). At this point ($t = t_3$ in Figure 10d), the crack generation
12 is not sufficient to release the excess of uniaxial stress. Then the lithosphere cannot release energy by neither
13 deformation nor crack generation mechanism. That is why the rupture (earthquake) occurs (green area in
14 Figure 10d) at $t = t_4$. After the main rupture, another aftershock occurs (green smaller patches within the
15 fault in Figure 10e). Nevertheless, the number of anomalies start to decrease. Finally, the micro crack
16 generation stops because the deformation is sufficient to handle the lithospheric response to non-constant
17 uniaxial stress (Figure 10f).

18
19 Additionally, Venegas-Aravena et al. (2019) found that the increase in the number of anomalies
20 are controlled by the same fractal nature that drives the micro crack generation. This means that the
21 frequency of the electrical flux could cover several magnitude orders. For example, figures 4, 5, and 6 are
22 characterized for the rise of different frequencies (micro to mili-hertz), which are known as ultra-low-
23 frequency (ULF), prior main earthquakes. These frequencies ranges were also found and described by
24 others researches as Fenoglio et al. (1995), Vallianatos and Tzanis (2003), Fraser-Smith (2008), De Santis
25 et al. (2017), Cordaro et al. (2018), among others.

26
27 Finally, it has been concluded that there must be precursory magnetic anomalies of the order of
28 0.1 nT related to earthquakes on the earth's surface (e.g. De Santis et al, 2017, Chernogor, 2019, Venegas-
29 Aravena et al., 2019). In the previous section, it was found that the minimum value to define an anomaly
30 was close to 0.2 nT. Therefore, this experimental result is in agreement with the theoretical value obtained.
31 Consequently, as the seismo-electromagnetic theory indicates, those magnetic anomalies may have a
32 lithospheric origin. Furthermore, the behavior of all these anomalies has a preceding increase similar to that
33 of other seismic events that use different data and methods (e.g., De Santis et al., 2019; De Santis et al.,
34 2019b and references therein).

35 36 **5.- Discussions and Conclusions**

37
38 The most significant characteristics of the magnetic field and its variations are found in the z-component,
39 which we have observed and recorded at the Putre and IPM observatories. The previous measurements
40 show that there is evidence of a progressive increase in the phenomenon known as the South Atlantic
41 Magnetic Anomaly (SAMA) (Cordaro et al., 2019). As expected, it generates a significant deviation in the
42 intensities present in the OP station as it is shown in the magnetic iso-values (Figure 1). Combining this
43 information with data from the IPM station, the behavior of the radial component of the geomagnetic field
44 for the three most significant seismic events in the Chilean Pacific sector during the period 2010-2015 was
45 recorded, and it corroborates the magnetic relation with seismology shown in Potirakis et al. (2016),
46 Contoyiannis et al. (2016), De Santis et al. (2017), which have been used other methods.

47
48 The normal magnetic trend showed some long-term variations. For example, there were breaks in the trend
49 or jump of B_z , followed by a time-lapse, and seismic movement as one can observe Figure 3. These jumps
50 occur in different forms: in Putre they are significant, reaching values of tens of nT, while in IPM the jump
51 is barely 10nT. The time lapse between each jump and the seismic event differs in each event. For Maule
52 2010 it was 36 days, for Iquique 2014 it was 96 days, and for Illapel 16 days. This time difference may be
53 due to an important factor: it appears that the jump is not equally strong in the three events, since the jump
54 before the Iquique 2014 event was considerably. weaker than the one before Illapel 2015, and preceded the
55 event by a longer time lapse (96 days). The more abrupt jump recorded in Illapel was followed by a shorter
56 time lapse (16 days). These changes are notorious, that is why a first approach by using frequency analysis
57 was done.

58
59 Specifically, significant frequencies data obtained for Maule earthquake Chile, 2010, range from 4.747 to
60 5.154 μ Hz, for Tohoku earthquake. Japan 2011 from 4.747 to 5.606 μ Hz, for Sumatra earthquake

1 Indonesia 2004 from 3.481 to 5.425 μ Hz (Cordaro, et al 2018). These fundamental frequencies were
2 detected before the earthquake in the areas of Pacific Ocean in the Southern Hemisphere, in the Eurasian
3 (2011) and Philippine (2004) areas in the Northern Hemisphere. Now these significant frequencies are
4 obtained again in different places and time on earth: in Iquique 2014, peaks of 4.611, 4.882, and 5.154 μ
5 Hz and for Illapel 2015, 3.739, 4.630, and 5.520 μ Hz (Figure 4). Up to this point, the rise of these
6 frequencies could be thought as a normal magnetic behavior with a high degree of coincidence. That is
7 why, other methodologies were performed in order to clarify the origin of these frequencies.
8

9 From now on, in order to avoid bias or technical malfunction, we decided to use different stations that
10 belong to an international network with open source data (supermag). These stations (OSO and PIL) were
11 the closest to the three earthquakes that had continuous measurements. The time period was 1 year before
12 each earthquake and 1 year after, giving 6 years of combined measurements where the frequency sample
13 was 1 data per minute.
14

15 The first approach was performed by using wavelet analysis at OSO station. Here, in order to avoid normal
16 variations and external perturbation, daily average values were performed by imposing a lower and upper
17 restriction before applying wavelet analysis. The Figure 5a shows the increase of the frequency range (> 2
18 μ Hz) \sim 30 days before the Maule 2010 earthquake. These frequency activities last up to \sim 10 days after the
19 earthquake. Similar frequency results were obtained for Iquique 2014 (Figure 5b) and Illapel 2015 (Figure
20 5c). Despite the abovementioned results, the previous restrictions might be seen as arbitrary. That is why
21 we moved to a stricter, stronger and bias-free methodology. Besides, three facts should be taken into
22 account. The first one is considering the physics-based filter processes, which remove most of the noise
23 and external disturbances. Thus, it allows performing more simple frequency analysis as the moving Fourier
24 transform (spectrogram). The second one is the tridimensional representation. This is important because it
25 is possible to observe the relative frequency intensity differences in a proper way. Furthermore, the third
26 one is that this spectrogram, and its tridimensional representation, allow us to compare our results to
27 previous works in the field (for example, Cordaro et al., 2018).
28

29 The definition of magnetic anomalies was performed in subsection 3.4.1. There, the anomalous magnetic
30 variations were defined by using statistical analysis. That is, one variation or peaks will be considered
31 anomalous if it reaches values beyond a certain threshold, threshold defined by the same data. In order to
32 avoid the external perturbations, Dst index and quiet time were considered. This gives six years of combined
33 data with variations that could be associated to lithosphere or internal. An increase in the frequency range
34 (\sim 1 mHz) before each earthquake was obtained after applying spectrogram analysis (Figure 6). If we look
35 at those periods not close to the earthquake's occurrences, almost no frequency activity was recorded. In
36 addition, these frequencies cannot be considered as part of tidal effects because the last one belongs to a
37 different frequency range (\sim 0.01-0.06 mHz) (Casotto and Biscani, 2004, Park et al., 2005). Prior studies
38 have shown that the frequencies in the range of \sim mHz are also related to the earthquake preparation stage
39 (Zlotnicki et al., 2001). This implies that Maule 2010, Iquique 2014, and Illapel 2015 occurred during very
40 high-frequency activity comparable to those found by Zlotnicki et al. (2001). As the considered data is free
41 of external perturbations, and earthquakes occurrences are within these frequencies' activations, the idea of
42 the existence of lithospheric frequencies related to earthquakes are reinforced.
43

44 In order to compare these data with other results, we performed the count of the daily anomalies. Here, the
45 anomalies behave as a sigmoidal function (Figure 7, 8, 9). In all of the earthquakes there was found a
46 dramatic increase in the number of anomalies between 50 to 90 days before each earthquake. This long-
47 term behavior is similar to those found in Nepal 2015 (De Santis et al., 2017), Mexico 2017 (Marchetti and
48 Akhoondzadeh, 2018), Central Italy (Marchetti et al., 2019a), Indonesia 2018 (Marchetti et al., 2019b) or
49 other big earthquakes worldwide (De Santis et al., 2019b). Note that the above-mentioned studies used
50 satellite data in contrast with this study which employs ground-based magnetometers and different
51 methodology. Additionally, the station selection followed the preparatory phase described by Dobrovolsky
52 et al. (1979). This means that any magnetic station, close to impending earthquake (\sim 1000 km), should
53 detect anomalies or the lithospheric microcracking beneath Earth's surface. The horizontal distance of the
54 preparation phase also agrees to geodetic findings. For example, Bedford et al. (2020) found a preparation
55 phase characterized by a high increase in the strain close to \sim 1000 km in subduction margin. Note that the
56 dates when anomalies rise dramatically are Feb 6, 2010 for Maule and Jan 8, 2014 for Iquique as it is shown
57 in Figure 8. These dates match well compared to the onset of critical seismicity throughout the concept of
58 characteristic precursory minima (β) in the framework of the Natural Time Analysis. For instance, Feb 1,
59 2010 and Dec 28, 2013 respectively (Sarlis et al., 2015). In addition, Figure 6b shows that the second (main)
60 rise of frequencies begins on Dec 27, 2013. On the other hand, it is possible to observe that the onset of a

1 second rise in the cumulative number of anomalies for the Mexico 2017 earthquake is close to the day 118
2 (Figure 9) that corresponds to Jul 28, 2017. Our date is almost the same date (27 July, 2017) obtained by
3 Sarlis et al. (2019) by applying the critical seismicity methods. It is then possible to claim that these
4 similarities among frequencies, cumulative anomalies, and seismicity features should be considered
5 different manifestations of a lithospheric phase transition.

6
7 By considering the above and the four filters applied (DST, quiet daily time, stochastic and
8 recurrences) which led the definition of residual anomalies, it is possible to explain our results in terms of
9 the physical mechanism described in the seismo-electromagnetic theory. This scheme explains different
10 empirical observations that indicate a direct relation between magnetic fields and earthquakes in which one
11 the essential group of measurement corresponds to the recording of ultra-low-frequency magnetic signal,
12 mainly close to millihertz and microhertz (Venegas-Aravena et al., 2019, 2020). It is important to note that
13 this theory considers the microcracking process (due to stress changes in the semi-brittle-ductile regime)
14 as a fundamental physical unit because it has been widely studied their electrical properties (e.g., Triantis
15 et al., 2020) or their influence in the propagation of the main failure (e.g., McBeck et al., 2021). Under this
16 concept, the anomalies correspond to the manifestation of the crack or micro crack within the lithosphere
17 which allows the flux of electrical current. Due to the Zener-Stroh mechanism (e.g., Stroh, 1955, Ma et al.,
18 2011), while changes in the seismicity rate are due to changes in the b-value (Venegas-Aravena et al., 2020)
19 which could generate seismic foreshock or slow slips as in Iquique 2014 (e.g., Herman et al., 2016). Also,
20 this theory considers the main earthquake as a crack that releases seismic energy as coseismic magnetic
21 signals (Kanamori, 1977, Utada et al., 2011). The temporal evolution of these cracks and its relation to the
22 sigmoidal magnetic anomalies' behavior can be seen in Figure 10. The framework of this theory states that
23 the micro cracks appear as a consequence of the excess of shear stress that cannot be released by the
24 lithospheric deformation (Venegas-Aravena et al., 2019). Then, frequencies rise and anomalies behavior
25 should be considered as a manifestation of the internal lithospheric collapse at the last stage (preparation
26 stage) of the seismic cycle, when solids cannot hold more strain. Since the electrical currents are intense
27 after the linear regime (Triantis et al., 2020), the phase transition described by Sarlis et al. (2019) could be
28 considered a (physical and statistical) manifestation of the changes in the semi-brittle-ductile regime. Then,
29 they can generate the above-mentioned magnetic anomalies found in this work.

30
31 Regarding the mechanism that generates the micro-cracks, we found that the minimum value to define an
32 anomaly was close to 0.2 nT, and this experimental result is in agreement with the theoretical value obtained
33 in Venegas-Aravena et al. (2019), where a ~ 0.2 nT rise when cracks are created at the semi-brittle ductile
34 regime (depth of 10-20 km) (Scholz, 2001, Sun, 2011).

35
36 Let us mention that the frequencies obtained by the Fourier analysis and anomalies are inherent to the
37 lithosphere. The variation of the low frequencies before the earthquake in the magnetic field is part of the
38 ionosphere-atmosphere-lithosphere coupling. Previously, we have shown that the frequencies in μHz are
39 related to the Maule earthquake in 2010 (Mw 8.8) (Cordaro et al. 2018). According to Vallianatos and
40 Tzanis (2003), the magnetic field frequencies, which are possibly related to earthquakes, could span a range
41 of at least three orders of magnitude. Specifically, they have detected a range of frequencies between 5 -
42 100 $m\text{Hz}$ before one month before earthquakes based on the Ionosphere-Lithosphere-atmosphere coupling.

43
44 We also remark that the possibility to predict the future occurrence of these seismic events is not yet
45 possible because the seismological mechanism of seismic movements is not yet precise. This means that
46 the role played by the fault's rupture parameters are not well understood. Specifically, the heterogeneities,
47 frictional properties, asperities and fault roughness are relevant to increase the complexity of the nucleation
48 process that determine the released energy, earthquake size, fracture energy and ground motion (e.g., Saltiel
49 et al., 2017, Selvadurai, 2019, Heimisson, 2020). However, a correlation does appear to exist between a
50 cumulative number of magnetic anomalies, time-lapse, frequency arise, and the Maule 2010, Iquique 2014,
51 and Illapel 2015 earthquakes. This methodology could be used as a tool to show the behavior of some
52 geophysical variables to indicate plate movements in the future. This condition, based on the increase of
53 low frequencies in the range of ($\mu\text{Hz} - m\text{Hz}$) suggests that these magnetic variations in the radial
54 component are probably a necessary but not sufficient condition on the Chilean margin. Further
55 investigations on this subject are required.

56
57 The next experimental step in this analysis is to gather the measuring instruments of the network
58 (magnetometers, Neutrons, others) and their variables recorded in the lithosphere, ionosphere,
59 magnetosphere, cosmic rays particles,) as neutrons, making a synapse or communication between them, in
60 real time (Machine Method Learn and others), in order to detect, directions, intensities, start and end of

1 frequencies, magnetic clusters, anomalies, or others, that could allow us to generate a warning prior to a
2 seismic movement.
3

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2
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12

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Figure Captions

Figure 1: Left side: Latitudinal effect of the Geomagnetic cutoff rigidity projected over the Chilean convergent margin close to the 70° W meridian. The pink solid lines indicate the edges of tectonic plates. Nazca Plate from 18° North to 45° degrees latitudinal, The South American Continent on the South American Plate. The 45 ° to 79° of the Antarctic Plate. The black lines indicate the coast line. In blue the iso values of magnetic intensity due SAMA proximity. The symbols indicate the station's location. Right: History of Chilean earthquakes.

Figure 2: The Kp magnetic activity index for the periods prior to the Maule 2010 (top), Iquique 2014 (middle) and Illapel 2015 (bottom) earthquakes. [spidr NOAA] [WDCFG Kyoto University] .

Figure 3: Vertical Component B_z as a function of time at Putre and IPM stations. a) Maule 2010 at the Putre station, b) Iquique 2014 at the Putre station, c) Iquique 2014 at the Easter Island station and d) Illapel 2015 at Easter island station. Trends changes have been observed in the four cases.

Figure 4: a) Fast Fourier Transformation (FFT) of the second derivative B_z component at Putre station for different events: Maule 2010 and Iquique 2014. The rise of frequencies in the range of micro Hz are compared to the FFT of the second derivative at IPM station for Illapel 2015. b) FFT every 15 days for Iquique 2015 at the Putre magnetometer. c) FFT every 8 days for Illapel 2015 at the Easter Island magnetometer.

Figure 5: Wavelet for B_z at OSO station is shown. This graph is obtained by restricting the peaks considered in a band and daily average values during 2 years of measurements. Wavelet spectrum shows an increase prior and after the Maule Earthquake. Unlike the spectrogram method where it is enough to consider the anomalous peaks on a threshold, wavelet analysis is more complex to calibrate than spectrogram analysis (upper limit).

Figure 6: Spectrograms analysis of vertical magnetic components after the external influence is filtered. a) The rise of a range of frequencies (1-2.5 mHz) appear prior and after the Maule 2010 earthquake (OSO station). The active frequencies last less than 3 months. b) The rise of similar frequencies appears prior to the Iquique 2014 earthquake in the vertical component of PIL station. This frequency activity lasts more than five months. c) The solar events were intense during September 2015. Nevertheless, it can be seen as an increase in the spectrum since August 2015. This frequency activity lasts close to 3 months. Three earthquakes hit when exist the rise of ultra-low frequencies (mHz)

Figure 7: Accumulated magnetic anomalies of B_z and a linear interpolation in the period during two years starting on 29 February 2009. The data were taken at OSO Station. Close to the Main earthquake, the linear trend breaks and the number of anomalies increase. Other important seismic events hit near the stations during the last period. Nevertheless, it is not clear that the anomaly increases are due these specific events.

Figure 8: Variation of the accumulated diary of magnetic anomalies of B_z during two years close to the three earthquakes: (a) Maule, (b) Iquique and (c) Illapel. The data were taken at OSO station (a) and PIL station (b & c), respectively. Is clear that the sigmoidal shape is similar in all of the earthquakes. This means that these stations recorded a dramatic increase in the number of magnetic anomalies between 50 to 90 days prior each earthquake.

Figure 9: (Upper panel) Accumulated Diary of magnetic anomalies during two years, in component Y from Apr 1 to Oct 15, 2017 in Mexico Earthquake Sep 8, 2017 Mw8.2. (Lower panel) Residual behavior of Mexico Earthquake. The data is open source and was taken from the swarm project (<ftp://swarm-diss.eo.esa.int/>). Methodology was developed by Marchetti and Akhoondzadeh (2018).

Figure 10: Schematic representation of the seimo-electromagnetic theory. The anomalies generation are owing the creation of several microcracks. The number of cracks increase because the internal collapse of the lithosphere when a non-constant uniaxial stress is applied.

1 **Table Captions**

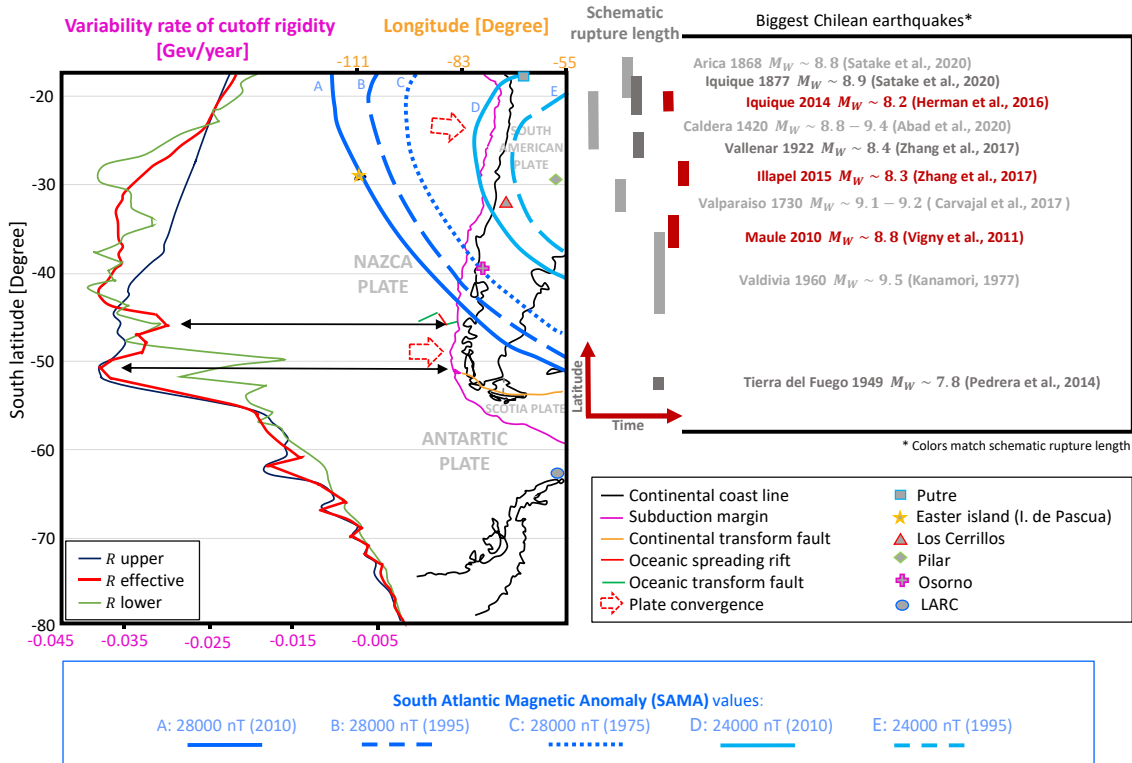
2
3 **Table 1:** The main characteristics for the detector of Chilean network Cosmic Rays and Geomagnetic
4 Observatories as location, altitude, and atmospheric depth, type of detectors (Cordaro et al., 2012).

5
6 **Table 2:** The maximum radius where the ionosphere-lithosphere-atmosphere coupling may affect magnetic
7 measurements to each earthquake studied at the station of Putre and IPM. (Dobrovolsky et al., 1979,
8 Pulinets and Boyarchuk, 2004). The Preparation area or Dobrovolsky area is defined by the radius
9 $r = 10^{0.43M}$, where M is the earthquake magnitude. This table shows that Putre and IPM stations are within
10 the earthquake preparation stage for Maule, Iquique and Illapel.

11
12 **Table 3:** Days before and after each frequency or anomalies rises for each considered earthquakes (Maule
13 2010, Iquique 2014, Illapel 2015 and Mexico 2017 for anomalies).

1

FIGURES



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Figure 1

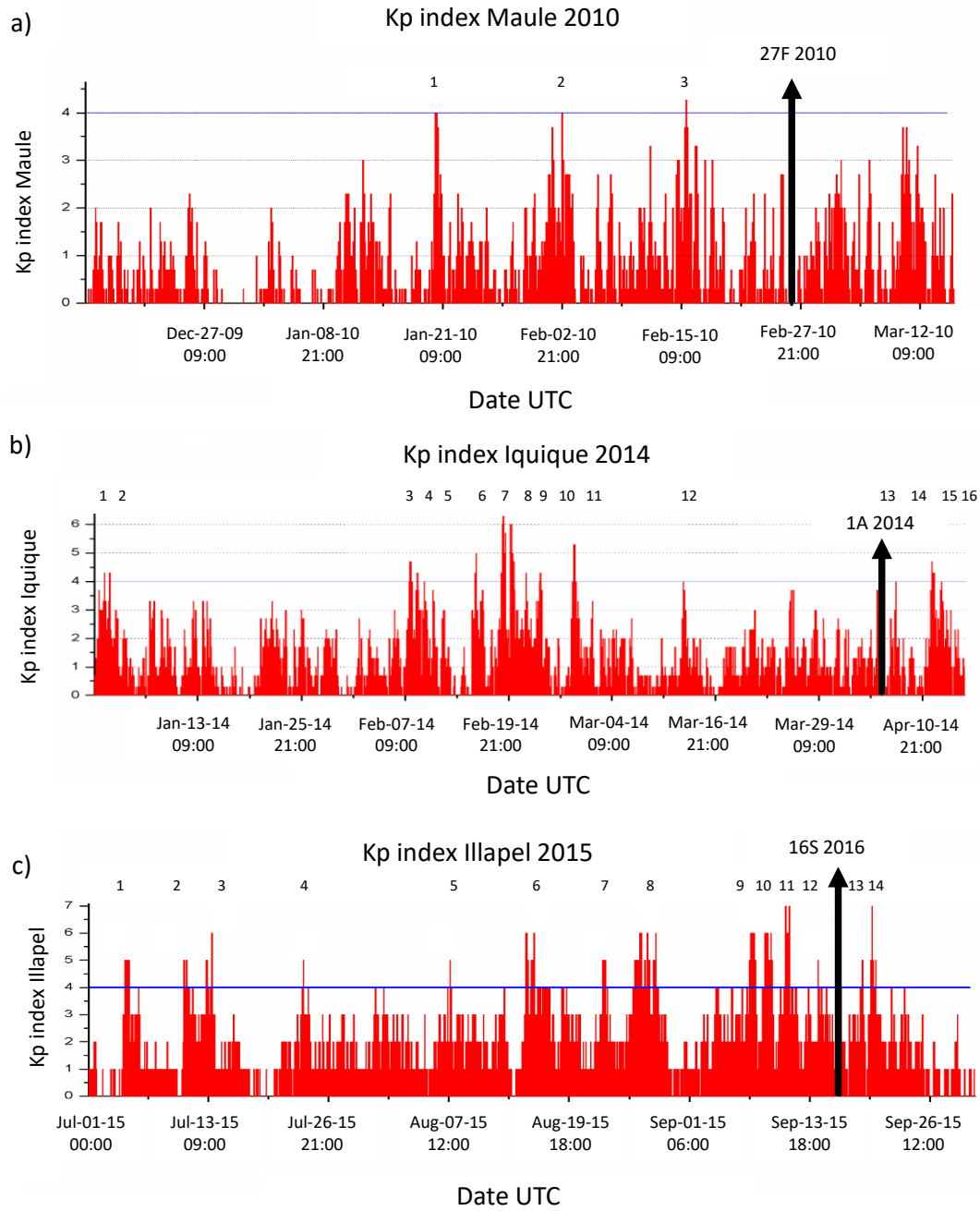


Figure 2 a,b,c

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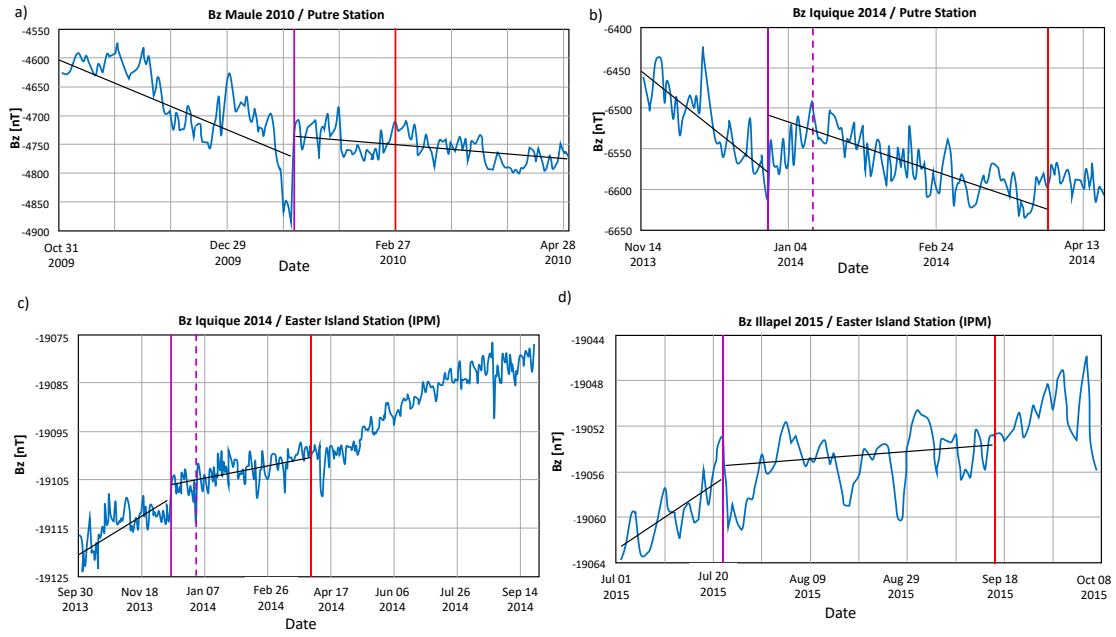


Figure 3 a, b, c, d

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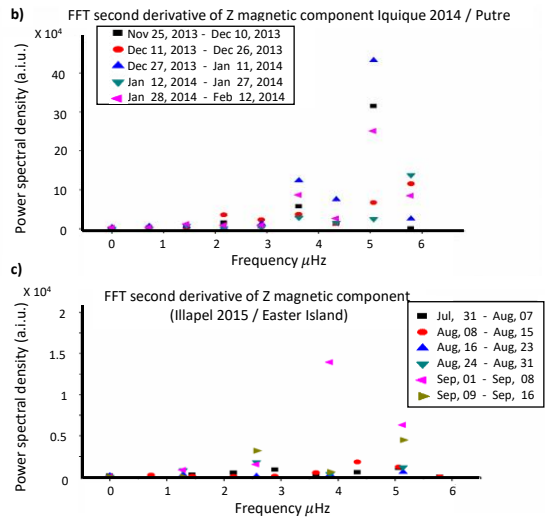
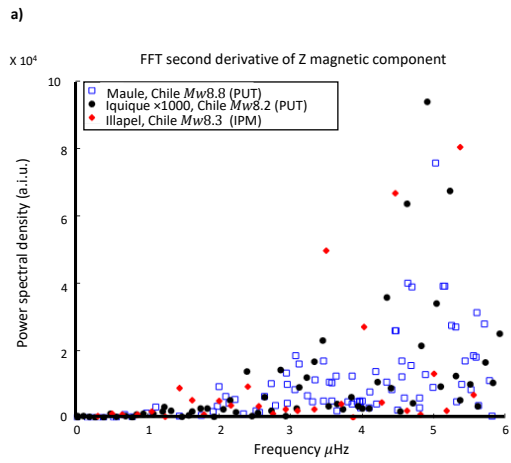


Figure 4 a,b,c

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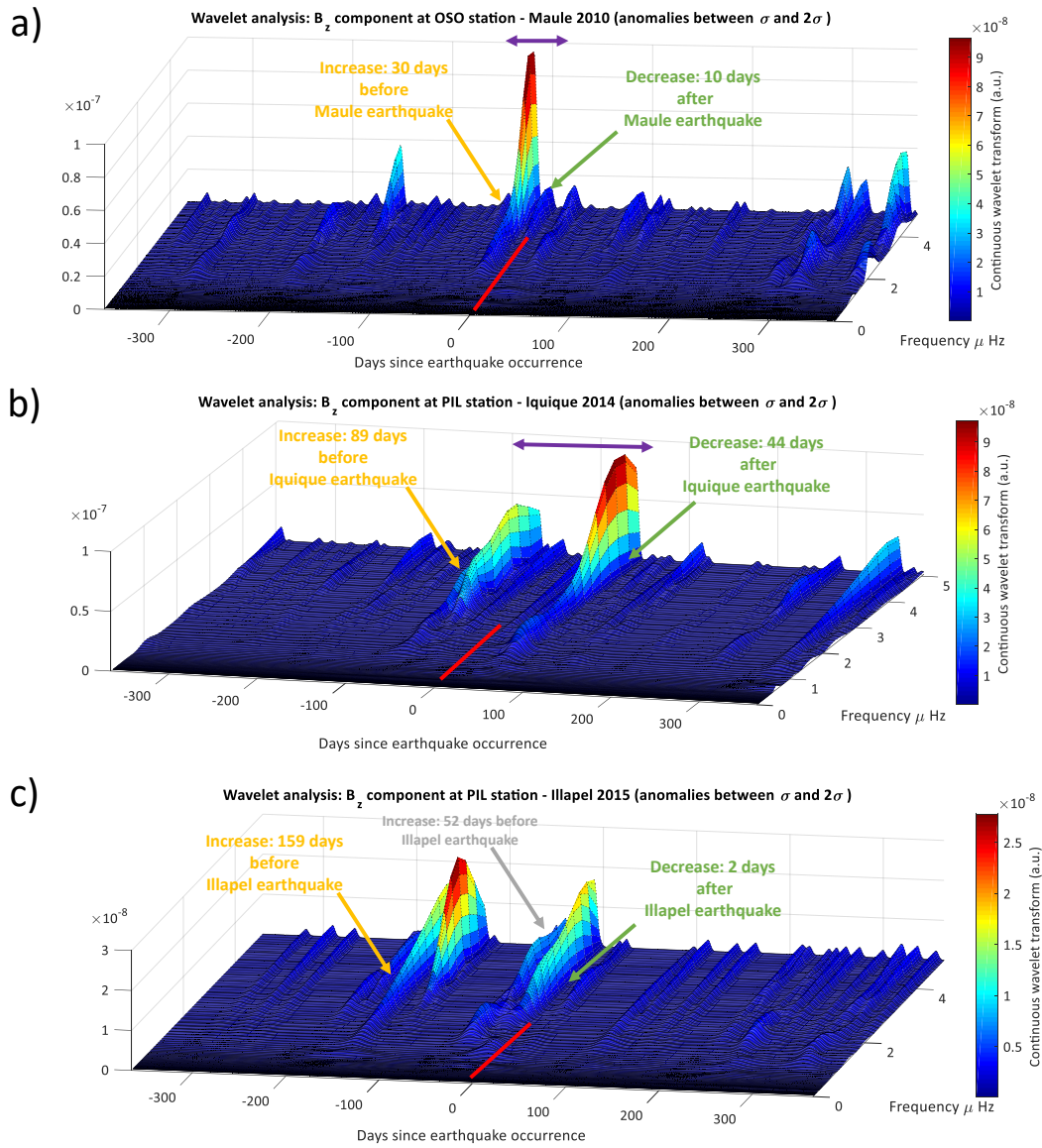


Figure 5

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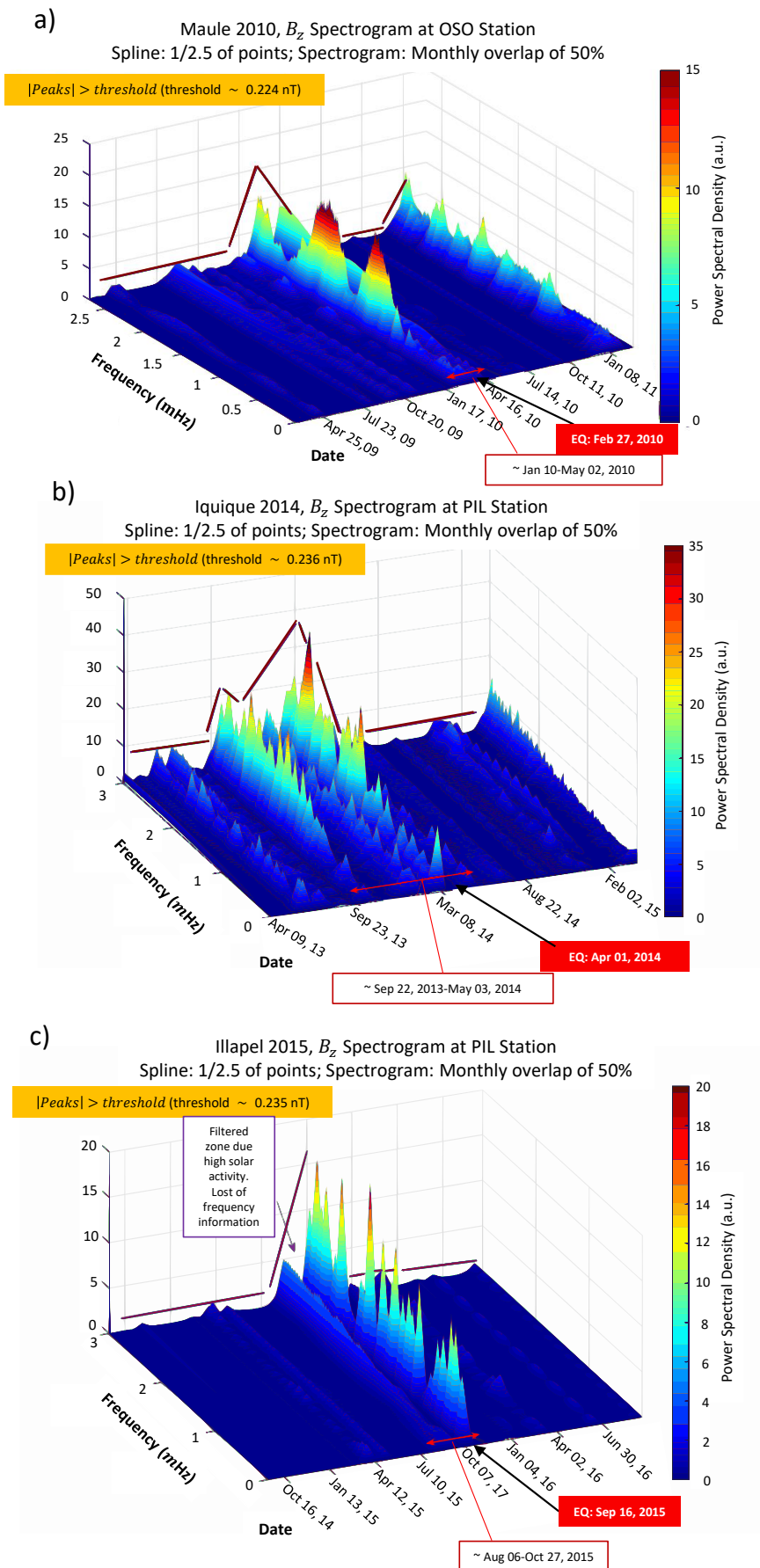
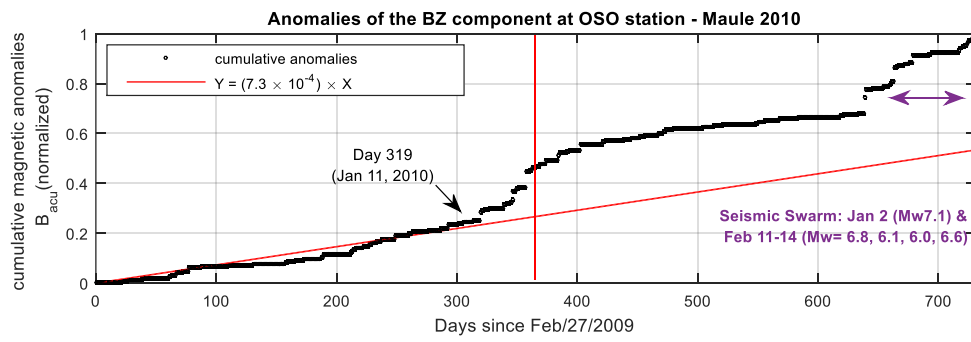


Figure 6

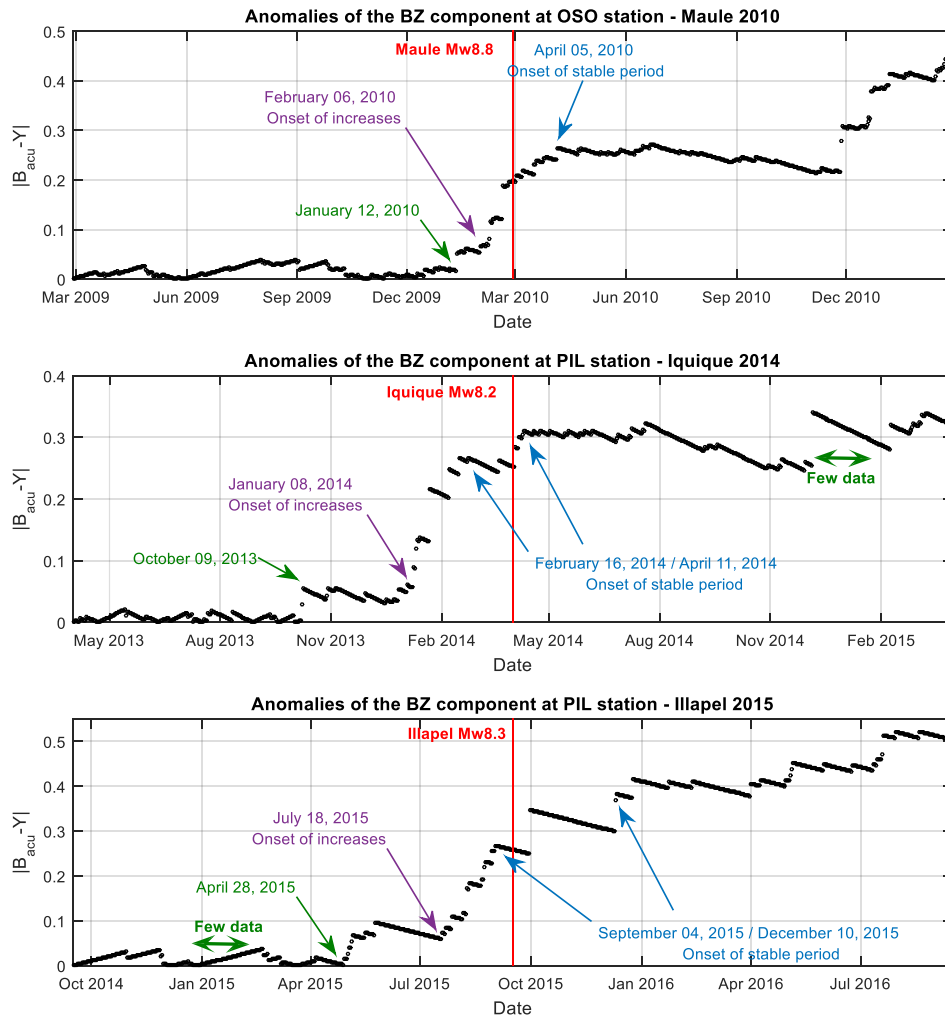
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Figure 7

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Figure 8

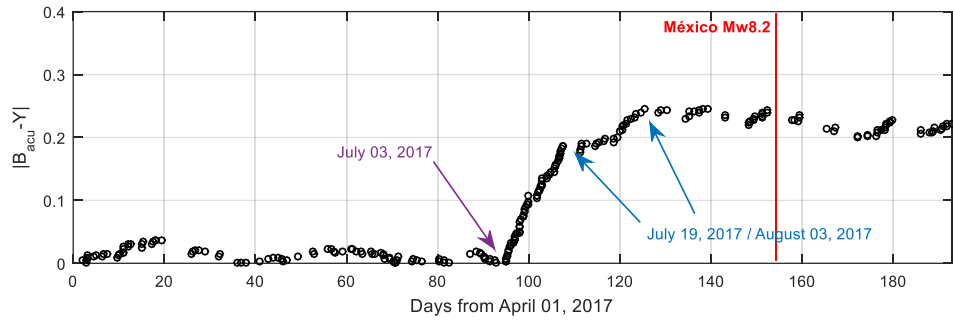
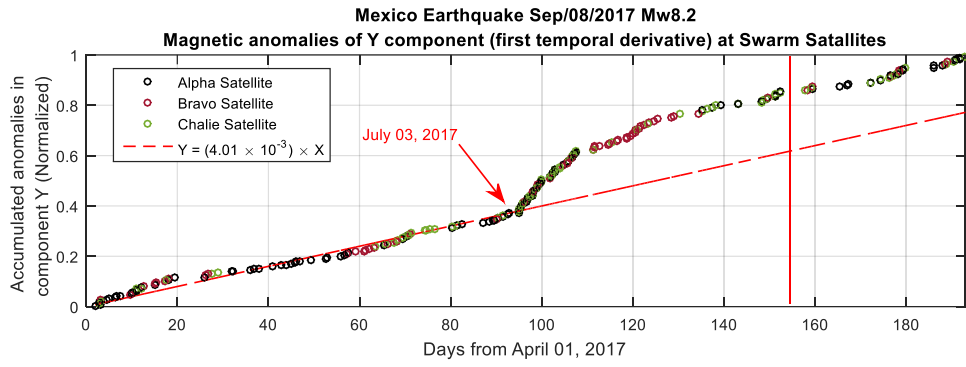


Figure 9

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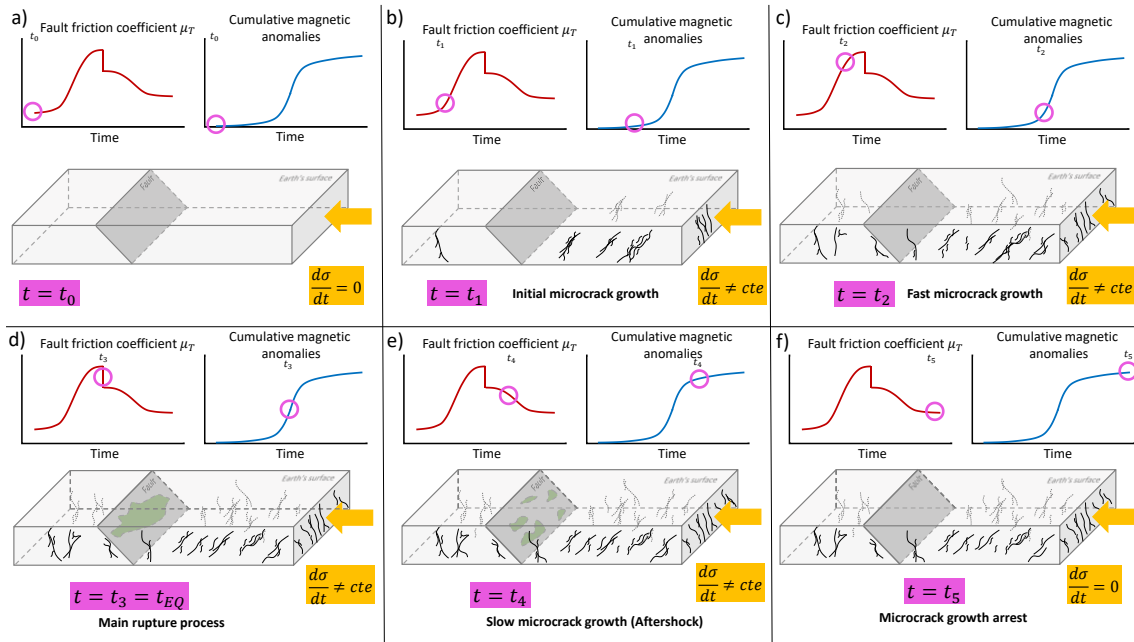


Figure 10

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Observatory	Location	Geographical coordinate	Altitude [m.a.s.l]	Atmospheric Depth [g/cm ²]	Instruments (Cordaro et al., 2012)	Time
PUTRE (PUT)	Andes Mountain, Chile	18°11'47.8" S. 69°33'10.9" W	3.600	666	Magnetometer, UCLA-Vectorial-Flux Gate. Muon telescope, 3 channels. Neutron monitor IGY, 3 channels, He-3. UTC by GPS receiver.	2003-2017
Los Cerrillos (OLC)	Santiago de Chile, Chile	33°29'42.2" S. 70°42'59.81 W	570	955	Magnetometer, UCLA-Vectorial-Flux Gate. Multi-directional muon telescope, 7 channels. Neutron monitor 6NM64, 3 channels, BF-3. UTC by GPS receiver.	1958-2017
LARC	King George Island, Antarctic	62°12'9"S. 58°57'42" W	40	980	Magnetometer, UCLA-Vectorial-Flux Gate. Neutron monitor 6NM64 - BF-3BF-3. 6 channels. Neutron monitor 3NM64 - He-3. 3 channels, Neutron monitor 3NM64 - He-3.[Flux meter] 3 channels. UTC by GPS receiver.	1990-2017

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6**Table 1**

1

Event	Magnitude [Mw]	Radius r [km]	Station Distance from earthquake [km]
Maule 2010	8.8	~6100	Putre ~ 2030
Iquique 2014	8.2	~3360	Putre ~ 300
Illapel 2015	8.3	~3700	IPM ~ 3700

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Table 2

1

Event	Rise of frequencies				Rise of cumulative anomalies, days before earthquake (secondary rise)
	Wavelet		Spectrogram		
	Days before earthquake (secondary rise)	Days after earthquake	Days before earthquake (secondary rise)	Days after earthquake	
Maule 2010	~30	~10	~48	~64	~46 (~21)
Iquique 2014	~89	~44	~191 (~95)	~32	~83
Illapel 2015	~159 (~52)	~2	~41	~41	~60
Mexico 2017	-	-	-	-	~67 (~42)

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Table 3