

Authors' response to Editor's comments

Dear Editor,

On behalf of myself and all of the co-authors, I would thank you sincerely for your work. We worked on your suggestions, trying to improve the discussion on the possible outcomes deriving from our paper related to landslide hazard. We think the addition of these sentences significantly improves the paper.

Best regards,

Sandro Rossato

Comment

Also a quick note from my side: I would suggest a better description (4-5 sentences at the end of the discussion, and then 1-2 sentences in the conclusions) of the implications of your study for natural hazards assessment and risk management of such study area in order to better meet the goals of our journal NHESS (that is not a pure "geomorphologic" journal).

Reply

We reconsidered the hazard perspective of the paper, retrieving and better discussing those parts we had earlier deleted to accommodate the comments by reviewer 1. In this perspective, we propose another change to the title. We think that the new one is more comprehensive of all aspects of the article.

Text added in the manuscript

The title has been replaced with the following: "Timing, drivers and impacts of the historic Masiere di Vedana rock avalanche (Belluno Dolomites, NE Italy)"

In addition to several minor changes we made to the initial and middle parts of chapter 5.3, we redesigned the final part to gather all discussion on natural hazards assessment and risk management. The text now reads as follows:

"The Veneto region is prone to earthquake activity and the study area is categorized as level 2 seismic hazard ("possible strong earthquakes" in Ordinanza del PCM n. 3519/2006), as the historical record testifies (up to $M_w=6.5$; Vigano` et al., 2013, 2015; Rovida et al., 2016). Continuous instrumental monitoring of the Belluno area dates back only to 1977 (Sandron et al., 2014). Preceding cataloged major seismic events in the region are based on either historical chronicles, dated building damages and/or observed rockfalls (Piloni, 1607; Taramelli, 1883; Guidoboni et al., 2005, 2018). Within a radius of 30 km from Mt. Peron, eight earthquakes with M_w greater than 5.0, and one exceeding 6.0, were recorded (Fig. 1). The strongest ($M_w = 6.3$) occurred just 20 km to the east of Mt. Peron on 29th June 1873 (Rovida et al., 2016, and references therein). Severe damage to Belluno city was reported during the Asolo (25th February 1695 AD; $M_w=6.4$), Friuli (6th May 1976 AD; $M_w=6.5$) and Verona (3rd January 1117 AD; $M_w=6.5$) earthquakes whose epicenters were located, respectively, 60, 65 and 140 km away (Guidoboni et al., 2018). As for the time frame suggested by our chronology, historical records report an important seismic event in July 365 AD that resulted in damage to the city of Belluno (Piloni, 1607). These data suggest that the Belluno area is sensitive to seismic shakings originating even hundreds of km away. Galadini et al. (2005) discuss evidence that active tectonics plays a key role in the reported intensification of slope instability registered in this area during the last 1500 yr. The most important effect of the frequent seismic activity is the progressive increase in the rock fatigue, with the formation of failure surfaces and the removal of rock bridges and roughness on discontinuity planes (Friedmann et al., 2003; Brideau et al., 2009; Parker et al., 2013; Stead and Eberhardt, 2013; Preisig et al., 2015; Gischig et al., 2016). Earthquakes have been suggested as triggers for several Alpine rock avalanches (e.g., Prager et al., 2009; Grämiger et al., 2016; Ivy-Ochs et al., 2017; Köpfli et al., 2018) and are known to have caused several rockfalls in the Belluno area (e.g., Piloni, 1607; Miari, 1830; Guidoboni et al., 2005). Moreover, they are considered as a possible trigger for failure of some pillars located on the Mt. Peron southern side (Di Giusto, 2012).

The predisposition of the Mt. Peron southwest face to failure is attributable to the structural setting and the regional-scale framework. The entire Belluno Dolomites experienced a long deformation history since the Miocene (Doglioni, 1990), related to regional-scale stress connected to the counter-clockwise rotation of the Adria plate, indented with the Alpine orogeny (Márton et al., 2003; D'Agostino et al., 2008). The thrusts and backthrusts (WSW-ENE oriented) that encase Mt. Peron relate to this phase of activity, also created the two conjugate fracture sets (NW-SE oriented) and induced the reactivation of the Jurassic faults (N-S oriented) and the overturning of the beds. The area between the Belluno thrust to the south, the Val

Carpenada - Val di Vido - Val Madonuta backthrust to the north, the Caorame Valley to the west and the Piave Valley to the east (Fig. 1) is structurally homogeneous and characterized by the same fracture sets identified at Mt. Peron (Bosellini et al., 1981; Masetti and Bianchin, 1987; Bigi et al., 1990; Costa et al., 1996). In limestones, where these intersect, rock dissolution and caves can form (Filipponi et al., 2009; Sauro et al., 2013), further weakening the rock (Pánek et al., 2009; Gutierrez et al., 2014). The Mt. Peron southern wall is known locally as the ‘weeping rock’ due to the numerous caves and karst springs along the steep rock face (Fig. 2b).

Our dating of the Masiere di Vedana landslide to late Roman times casts considerable doubt on the previous Lateglacial chronological assessment (e.g., Pellegrini et al., 2006). Consequently, stress release following downwasting of the LGM Piave glacier, that had been previously considered as the main driver of the rock avalanche, actually played no direct role in the process. Climatic and tectonic factors were much more important. In addition, the classification of the deposits as Lateglacial led to underplaying of the possible hazard at Mt. Peron. The situation at Mt. Peron, with steeply dipping-to-overturned bedding in a limestone massif crisscrossed by numerous faults and riddled with karst fissures, is presently the same that produced the massive failure of the Masiere di Vedana. Several hundred-meter-high rock prisms along the top of the headscarp are partially detached along the discussed fracture systems and loom perilously over the inhabited valley below (Di Giusto, 2012). Our results suggest the need for a reconsideration of the hazard related to not only the Mt. Peron southern face, as the rock slopes still present evident structural weakness, but also the whole area lying between the Belluno thrust and its backthrusts, as intense rainfall and earthquakes can occur at any time."

Structural Timing, drivers and geomorphological characterization and dating impacts of the historic Masiere di Vedana rock avalanche (Belluno Dolomites, NE Italy)

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Abstract. The “Masiere di Vedana” rock avalanche, located in the Belluno Dolomites (NE Italy) at the foot of the Mt. Peron, is re-interpreted as Historic on the base of archeological information and cosmogenic ³⁶Cl exposure dates. The deposit is 9 km² wide, has a volume of ~170 Mm³ corresponding to a pre-detachment rock mass of ~130 Mm³, and a maximum runout distance of 6 km and an H/L ratio of ~0.2. Differential velocities of the rock avalanche moving radially over different topography and path-material lead to the formation of specific landforms (tomas and compressional ridges). In the Mt. Peron crown the bedding is subvertical and includes carbonate lithologies from lower Jurassic (Calcari Grigi Group) to Cretaceous (Maiolica) in age. The stratigraphic sequence is preserved in the deposit with the formations represented in the boulders becoming younger with distance from the source area. In the release area the bedding, the SSE-vergent frontal thrust planes, the NW-vergent backthrust planes, the NW-SE fracture planes, and the N-S Jurassic fault planes controlled the failure and enhanced the rock mass fragmentation. The present Mt. Peron crown still shows ~~hundreds of meters high~~ hundreds of meters high rock prisms bounded by backwall trenches. Cosmogenic ³⁶Cl exposure ages, mean 1.90 ± 0.45 ka, indicate failure occurred between 340 BC and 560 AD. Although abundant Roman remains were found in sites surrounding the rock avalanche deposit, none was found within the deposit, and this is consistent with a Late Roman or early Middle Age failure. Seismic and climatic conditions as landslide predisposing factors are discussed. Over the last few hundred years, earthquakes up to Mw 6.3 including that at 365 AD, affected the Belluno area. Early in the first millennium, periods of climate worsening with increasing rainfall are known in the NE Alps. The combination of climate and earthquakes induced progressive long-term damage to the rock until a critical threshold was reached and the Masiere di Vedana rock avalanche occurred.

1 Introduction

Landslides have an enormous impact on landscapes and can be a serious threat to human lives and buildings. Assessment of the potential for future events is distinctly dependent on knowledge of the conditions under which past failures occurred in the

immediate vicinity (Samia et al., 2017). This entails detailed analysis and interpretation of driving factors as well as possible
85 triggers of past events (Eisbacher and Clague, 1984; Nicoletti and Sorriso-Valvo, 1991; Hungr, 2006; Strom, 2006; Hermanns
and Longva, 2012). Bedrock bedding, faults, fractures and other discontinuities predispose a rock mass to fail (Stead and
Wolter, 2015). Damage accumulation in rock (fatigue) contributes to the location of failure (Friedmann et al., 2003; Brideau et
al., 2009; Parker et al., 2013; Stead and Eberhardt, 2013; Preisig et al., 2015), while seismic shakings (Keefer, 1993; Friedmann
et al., 2003; Dunning et al., 2007; Cui et al., 2011; Stead and Eberhardt, 2013) and periods of extreme rainfall (Guzzetti et al.,
90 2008; Tsai and Wang, 2011; Loew et al., 2017; Preisig et al., 2015) can trigger landslides.

In this perspective, the Italian landslide inventory project (IFFI; Trigila et al., 2007; [http://www.progettoiffi.isprambiente.it/
cartografia-on-line/](http://www.progettoiffi.isprambiente.it/cartografia-on-line/)) is the ideal starting point for hazard maps. These are tools for landscape management and civil protection
plans, they thus require continuous updating and accurate input data. In the study area (Fig. 1), the southern side of Mt. Peron
(Belluno Dolomites, NE Italy) is classified in the most recent landslide hazard map (<http://www.geoviewer.isprambiente.it>) as
95 “Attention Area” (i.e., a failure is possible, but an evaluation is needed). In 2011 a rockfall (volume ~1000 m³, blocks up to 2.5
m³) detached from the upper part of the Mt. Peron which led the Municipality of Peron to commission the evaluation of the
hazard along its southern cliff (Di Giusto, 2012). Numerous partially detached rock prisms were recognized, up to 18,000 m³
and with trenches up to 50 cm wide. According to Di Giusto (2012), nine out of 16 pillars are at risk for failure in the case of
an earthquake.

100 The IFFI catalog, the landslide hazard map and the evaluation by Di Giusto focused on the Mt. Peron southern wall and the
scree slope, neglecting the well-known in literature deposit of Masiere di Vedana that lies on the southern plain (Abele, 1974;
Eisbacher and Clague, 1984; Pellegrini et al., 2006). This is also known as “Rovine di Vedana” (Mazzuoli, 1875; Squinabol,
1902; Montandon, 1933) or “Marocche di Vedana” (Dal Piaz, 1912; Venzo, 1939). The deposit covers an area of 8-9 km², has
a maximum thickness of 40 m and an estimated volume of 100-120 Mm³ (Abele, 1974; Genevois et al., 2006; Pellegrini et al.,
105 2006). The Masiere di Vedana is one of the largest catastrophic events in the Alps (Heim, 1932; Abele, 1974; Eisbacher and
Clague, 1984), comparable with large events in the Himalaya (e.g., Hewitt, 2006; Hewitt et al., 2008; Mitchell et al., 2007),
Rocky Mountains (e.g., Blais-Stevens et al., 2011; Charrière et al., 2016) and Andes (e.g., Hermanns et al., 2004; Welkner et
al., 2010).

As there is no consensus on the age and dynamics of the Masiere di Vedana, a re-evaluation is needed in the light also of
110 hazard assessment. The Masiere di Vedana deposit was interpreted as: a glacial deposit (Hoernes, 1892), a landslide transported
by a glacier during the Lateglacial (Mazzuoli, 1875; Squinabol, 1902; Dal Piaz, 1912; Venzo, 1939) and as the result of a
catastrophic flood due to collapse of a natural dam (Taramelli, 1883). Some other authors proposed as origin for the deposit: (1)
the combined effect of a landslide over a glacier, followed by a second landslide that evolved into a rock avalanche (Pellegrini,
2000; Pellegrini et al., 2006), and (2) a small rock slide followed by a larger rock avalanche (Genevois et al., 2006). The age
115 attribution for the deposit ranges from the Lateglacial (Pellegrini et al., 2006) to historical times (Piloni, 1607; Miari, 1830).

Amongst all types of landslides, rock avalanches are particularly relevant, being both difficult to predict (Hungr, 2006) and
representing a very high risk for the population living in mountain areas (Guzzetti, 2000; Hungr, 2004; Geertsema et al., 2006;
Evans et al., 2007; Sosio et al., 2008; Cui et al., 2011; Hermanns and Longva, 2012; Mitchell et al., 2020). The moving masses

are composed of dry debris, that in subaerial settings range from about 0.5 to more than 10,000 Mm³ (Crosta et al., 2007).

120 The initial phase, rockfall or rockslide, evolves into a flow-like movement of crumbling rock debris, sized from clay/silt up to decametric boulders, which can travel for several kilometres, even uphill, and overcome obstacles up to some hundred meters high (e.g., Hungr et al., 2001; Mangeney et al., 2010; Bowman et al., 2012). Rock avalanche deposits are characterized by inverse grading of the sediment, with large blocks dominating the carapace, the inclusion of path material and, in some cases, preservation of the stratigraphic sequence (Hewitt, 2002; Strom, 2006).

125 The aim of this study is to provide dating and to evaluate driving factors, potential triggers, and process dynamics of the Masiere di Vedana rock avalanche, in the light of a better assessment of the hazard, potential extent and runout also in similar settings.

2 Geological setting

The Masiere di Vedana lies at the mouth of the Cordevole Valley in a broad plain at the confluence of the Cordevole and Piave Rivers (Fig. 1). Mt. Peron (1486 m a.s.l.) is the south-western peak of the Schiara Group (highest peak: Mt. Schiara: 2565 m a.s.l.). The Mt. Peron is composed, from the west to the east, by Calcarei Grigi Group (Lower-Middle Jurassic), Vajont Limestone (Middle Jurassic), Fonzaso Formation, Rosso Ammonitico (Upper Jurassic) and Maiolica (Cretaceous) limestones. Their distinctive characteristics, useful for tracking the source of the Masiere di Vedana deposit, are: thick-bedded, fossiliferous blue-gray limestones (Calcarei Grigi Group); thick-bedded, locally oolitic limestones and calcarenites (Vajont Limestone); 135 thick-bedded siliceous limestones with clay interbeds (Fonzaso Fm.); pink to red nodular limestones rich in ammonites (Rosso Ammonitico); thick-bedded, white limestones with chert nodules (usually gray to black) containing nannofossils, calpionellids and radiolaria (Maiolica). Scaglia Rossa and the Cenozoic formations (Belluno flysch, Belluno glauconitic sandstone, Bastia siltstone, Libano sandstone and Bolago marl) crop out at the base of the Piz Vedana (Fig. 3), form the Castel Cuch ridge and underlie the fluvial plain between the Mis and Cordevole rivers. Outcrops of cemented Pleistocene fluvial gravels (“Roe” or 140 “Sass Muss conglomerate”; Costa et al., 1996) are located just to the west of the present course of the Cordevole River (near the town of Vignole in Fig. 3).

The investigated area is bounded by Alpine tectonic lineaments (Fig. 1): the Valsugana thrust fault to the north, the S-verging Alpine folds and thrusts to the south (Doglioni, 1990). The Mt. Peron belongs to the hanging wall of the WSW-ENE oriented Belluno thrust, one of the main tectonic lineaments of the eastern Southern Alps (Doglioni, 1990; Galadini et al., 2005), that 145 crops out at the northern limb of the Belluno syncline. The sedimentary strata of the forelimb are sub-vertical to slightly overturned (Doglioni, 1990; Costa et al., 1996) and they converge into the Belluno thrust. The Val Carpenada - Val di Vido - Val Madonuta thrust is the backthrust of the Belluno thrust (Costa et al., 1996). To the east, reactivated Jurassic faults (Masetti and Bianchin, 1987) displaced the Val Carpenada - Val di Vido - Val Madonuta thrust (Fig. 1) and induced a wealth of fractures in the Mt. Peron rock wall (Fig. 2). The Mt. Peron is at the nucleus of an ENE-WSW oriented anticline with a very steep 150 forelimb, followed southwards by the Belluno syncline (Fig. 2a) that hosts Cenozoic sedimentary units (Costa et al., 1996).

3 Methods

3.1 Field survey, structural analysis and remote sensing

Detailed geomorphological maps (Caneve, 1985; De Zorzi, 2013), aerial and satellite images (Google Earth and Bing databases) and DTM analysis (cell size: 5 m, vertical accuracy: 30 cm; <http://idt.regione.veneto.it/app/metacatalog/>) were used to obtain topographic profiles and to estimate the area of the Masiere di Vedana deposit. The areal distribution of lithologies in the deposit was gauged by observation of boulders in the field and verified by thin section analysis (Supplementary Material SM1). Six boreholes, up to 3 m deep, were taken in the fine-grained sediments of the Torbe area (Supplementary Material Figure SM2a) with a hand-auger (Edelman combination-type, EjikelkampTM), which allows the extraction of 10 cm-wide cylindrical cores. Orientations of bedrock discontinuities, such as bedding, foliation, joints, fractures and faults, were measured in the southern wall of Mt. Peron.

3.2 Cosmogenic ³⁶Cl exposure dating

Twelve different boulders located in topographically high positions with respect to the surroundings within the deposits were sampled for dating with cosmogenic ³⁶Cl. For boulders VB13 (VB13a, VB13b) and VB14 (VB5 same boulder as VB14) two samples were taken. Samples were taken to cover the full extent of the deposit, from right near the source area to the distal sector.

For ³⁶Cl sample preparation we used the method of isotope dilution as described by Ivy-Ochs et al. (2004). Total Cl and ³⁶Cl were determined at the ETH AMS facility of the Laboratory for Ion Beam Physics (LIP) with the 6 MV tandem accelerator. The ³⁶Cl/Cl ratios of the samples were normalized to the ETH internal standard K382/4N with a value of $^{36}\text{Cl}/\text{Cl} = 17.36 \times 10^{-12}$ which is calibrated against the primary ³⁶Cl standard KNSTD5000 (Christl et al., 2013; Vockenhuber et al., 2019). Full process chemistry blanks (3.4×10^{-15}) were subtracted from measured sample ratios. All fourteen rock samples were processed. Only seven were measured successfully due to too high ³⁶S, also in relation to the very low ³⁶Cl concentrations in these samples. All measured data are presented here. Major and trace element concentrations were determined with XRF (Supplementary Material SM3) and ICP-MS (Supplementary Material SM4), respectively. We calculated surface exposure ages with the LIP ETH in-house MATLAB code based on the parameters presented in detail in Alfimov and Ivy-Ochs (2009, and references therein). A production rate of 54.0 ± 3.5 ³⁶Cl atoms (g Ca)⁻¹ yr⁻¹, which encompasses a muon contribution at the rock surface of 9.6%; and a value of 760 ± 150 neutrons (g air)⁻¹ yr⁻¹. These values are in excellent agreement with production rates recently published by Marrero et al. (2016). Production from all major elements and through low energy neutron capture in light of the trace elements (Supplementary Material Table SM4a) were fully considered. Production rates were scaled to the latitude, longitude, and altitude of the sites based on Stone (2000). No correction was made for karst weathering of the boulder surfaces (cf. Styllas et al., 2018). The extent of karst dissolution on the boulder surfaces varies significantly from boulder to boulder. Implementing a rate of 5 mm/ka would change the ages by less than 4%, which does not affect any of the conclusions drawn here. Stated errors of the exposure ages (Table 1) include both analytical uncertainties and those of the production rates

(Alfimov and Ivy-Ochs, 2009). Two different surfaces of boulder 13 were analyzed (VB13a, 1.45 ± 0.12 ; VB13b, 1.45 ± 0.12 ka); the weighted mean of 1.45 ± 0.08 ka is used for further discussion.

185 4 Results

4.1 Mt. Peron release area

The Mt. Peron scarp is 700 m wide and 600 m high, S-to-SW facing and partially circular. No secondary scarps are present. Numerous faults and fractures are well visible on the wall (Fig. 2b) and are grouped into five main discontinuity sets (Fig. 2c). These comprise: (1) bedding, (2) WSW-ENE directed frontal thrust planes, (3) NW vergent backthrust-related planes, (4)
190 NW-SE aligned local conjugate fracture planes sets, and (5) persistent N-S oriented planes interpreted as reactivated Jurassic faults (Masetti and Bianchin, 1987). Bedding is nearly vertical, its orientation ranges from 146/78 to 170/80 (dip direction/dip angle). The Belluno thrust, average orientation 337/64 (Costa et al., 1996), crops out at the base of the steep wall, whilst other Belluno thrust planes (295/53 to 340/67) were measured higher up along the wall (Fig. 2b). The NW-verging planes related to the backthrust are 111/15 to 175/54 and steepen to 80° at higher elevations along the wall. The NW-SE aligned fractures are
195 209/16 to 245/32 with an associated conjugate set, from 20/44 to 62/30, and nearly-vertical fractures with dip direction of 219 to 255. The N-S striking fracture planes dip both to the east (75/40 to 83/75) and to the west (240/71 to 299/18). Today myriad large and small individual rock prisms bounded by these discontinuities are present in the upper part of the release area.

4.2 Masiere di Vedana rock avalanche deposit

The deposit covers an area of ~ 9 km² from the base of Mt. Peron southwards to Roe Basse (~ 5 km) and westwards to Mis
200 River (~ 3 km; Fig. 3). By means of open sections, the thickness of the deposit is estimated ~ 10 m in the proximal area (near boulder VB2, Fig. 3), ~ 15 m in the central sector near Torbe, >30 m near the boundary between the Vedana and Masiere sectors (Ponte Mas section, Fig. 4), ~ 5 m in the Masiere central sector and ~ 15 m in the southern (distal) sector (Suppiei section, Fig. 4). Using a mean thickness of 20 m, a rough estimation of the total debris volume of ~ 170 Mm³ is obtained. Such volume corresponds to a released rock mass of about 130 Mm³ (bulking coefficient of 25% cf. Genevois et al., 2006). With a vertical
205 drop (H) of about 1150 m (from the top of Mt. Peron: 1486 m a.s.l., to Roe Basse area: ~ 340 m a.s.l) and a travel distance (L) of about 5900 m (Fig. 3), we calculate an H/L ratio of ~ 0.2 (Fahrböschung angle of 11°). Based on spatial pattern, boulder lithology and morphological character, we distinguish five sectors of the deposit: Peron, Vedana, Torbe, Masiere and Roe (Roe Alte and Roe Basse).

The **Peron** sector includes the talus apron deposits at the foot of Mt. Peron, the rock avalanche deposits on the east side of the
210 river and the terrace of the town of Peron (at about 380 m a.s.l.). Boulders at the foot of the slope range up to 20 m in diameter (Fig. 5a). Three boulders in the Peron sector were dated with cosmogenic ³⁶Cl (Fig. 3; Table 1): VB3a (Rosso Ammonitico; 1.83 ± 0.28 ka), VB3c (Fonzaso Fm.; 3.62 ± 0.41 ka) and VB14a (Calcarei Grigi Group; 2.39 ± 0.39 ka). Based on the trend of all obtained ages, the age of VB3c is interpreted as an outlier, its age possibly reflecting the presence of inherited ³⁶Cl due

to pre-exposure. The town of Peron lies on a terrace made of rounded gravel layers with rare sand lenses that are interfingering
215 with talus deposits (Caneve, 1985).

In the **Vedana** sector, the rock avalanche deposit displays an irregular forested topography with relief on the order of tens of meters (Fig. 5b). Huge blocks hundreds of cubic meters in size, mostly made of Calcari Grigi, dominate the carapace. This covers the main body of the deposit made of shattered rocks, which is comprised of very angular clasts (up to tens of cm in diameter) in a matrix of silty sand. VB2 boulder (Calcari Grigi) gave an age of 1.49 ± 0.26 ka; it lies on top of ~10 m thick
220 sequence of rock avalanche deposits. In the Ponte Mas quarry (Figs. 3, 5g), an open section showed glacial till (up to 3 m thick) incorporated into the base of the rock avalanche deposit, its original bedding is completely obliterated. This sediment is >30 m thick and is composed of sub-rounded clasts (up to 20 cm in length), some of them striated, supported by a silty clay matrix. Clasts are sedimentary and volcanic, reflecting the catchment of the Cordevole paleoglacier (cf. Pellegrini et al., 2006). Several ENE-WSW trending incisions cut through the Vedana and Torbe sectors (main ones highlighted on Fig. 3). Irregular patches
225 of sandy-silty and fine gravel sediments are found in the Vedana low-lying areas between the blocky reliefs.

The **Torbe** sector encompasses the distal northern lobe of the rock avalanche, characterized by 10 to 20-m high isolated hills and hummocks (Fig. 5d) that emerge from a flat topography. They are roughly aligned ENE-WSW, circular at the base and have slope angles of 35°-40°. They are made of very angular Calcari Grigi boulders and clasts, with many jig-saw puzzle structures in a sandy, gravelly matrix (Fig. 5e). Such morphological structures are “toma” (Turnau, 1906; Abele, 1974; More
230 and Wolkersdorfer, 2019). They are found in association with some large rockslides and are mainly made of landslide material, in many cases showing a gradation from very comminuted fragments in the outer part to less fractured material at the core (cf. von Poschinger and Rugg, 2012; More and Wolkersdorfer, 2019). Six cores taken in the flat area between the hills (see Fig. 3 and Supplementary Material SM2) show up to two meters of fining upward silty sand above the rock avalanche. Torbe is crossed by the largest incision of the whole Masiere di Vedana (Fig. 3), ENE-WSW trending, ~50 m wide and up to 20 m
235 deep in respect to the mean topographic surface. Cenozoic lithologies crop out at the base of this incision, the rock avalanche deposit being ~15 m thick. A shallower incision, few m deep, is located at the base of Piz Vedana slope, still conveying a small amount of water coming from the Vedana Lake.

The **Masiere** sector is strikingly different from Vedana and Torbe. It is a bleak, vegetation-free, desert-like sea of limestone blocks, mainly of dolomitized Vajont Limestone (Fig. 5c). Boulders up to 3 m in diameter and abundant angular and sub-
240 angular clasts, with almost no matrix, are present in the surficial part. In the southern part of Masiere, numerous 2-3 m high and up to 150 m long ridges, aligned roughly E-W, are present (Fig. 6). The contact of the rock avalanche with the underlying Pleistocene conglomerate crops out near the southern boundary of the Masiere (white asterisks in Fig. 6), where the deposit is only ~5 m thick. Next to the northern boundary of Masiere, the Cordevole River flows upon the Cenozoic rocks of the Castel Cuch ridge, covered elsewhere by the rock avalanche. An ENE-WSW trending shallow incision (Fig. 3) with associated well
245 sorted, medium-to-coarse grained sandy deposits, is present, almost parallel to Castel Cuch. A roughly N-S trending incision, few meters deep, ~10 m wide and about 500 m long, is located just to the west of the Cordevole River. The town of Mas is built upon a flat terrace, mainly made of sorted rounded gravels with a sandy matrix, that bounds the Masiere to the east.

Roe Alte and **Roe Basse** comprise the distal sector of the Masiere di Vedana to the south (Fig. 3), where angular clasts up to 20 cm in diameter and very few boulders immersed in a sandy matrix are scattered in the meadows. Boulders belong to the Fonzaso Formation, Rosso Ammonitico and Maiolica, very few of them made of Vajont Limestone. Roe Basse is made of silty/sandy alluvial sediments deposited by the Gresal and other minor streams coming from the east, mantling the rock avalanche deposit on the northwestern side. The Suppiei section, 100-m-long and 25-m-high, on the left flank of the Cordevole River (Figs. 3, 5h; Supplementary Material SM5), shows the Bolago Marl unconformably overlain by 0.5-2 m of glacial till. This is in turn covered, with a sharp and undulated contact, by up to 20 m of rock avalanche debris decimetric in size, with boulders (~1 m diameter) on top. On the Roe Alte rocky upland, the rock avalanche is at most 2 meters thick, with rare boulders (1-2 m diameter). Two boulders made of Vajont Limestone (Fig. 3) have been dated with ^{36}Cl (VB12, 2.35 ± 0.21 ka; VB13, 1.45 ± 0.08 ka; Table 1). South of the town of Mas, the Cordevole River flows into narrow meanders entrenched ~20 m into rock avalanche deposits, alluvial material, glacial sediments and bedrock. The terrace of the Vignole village is almost totally made of rock avalanche debris, despite being remarkably flat.

5 Interpretation and discussion

5.1 Age of the Mt. Peron rock avalanche

^{36}Cl surface exposure ages from boulders all across the deposit range from 1.45 ± 0.08 ka to 2.39 ± 0.38 ka (Table 1). All ages show a good overlapping within uncertainties. A single sample (VB3c) gave a result markedly different from the others: 3.62 ± 0.41 ka. Although this age may point to pre-exposure of the sampled boulder surface (cf. Sewell et al., 2006; Merchel et al., 2013), as for example seen at Lavini di Marco (Martin et al., 2014), the possibility exists that this boulder is part of a partially buried older deposit located right at the foot of Mt. Peron. The poorly developed karst dissolution features (0.5-1 cm deep karren) on the tops of many boulders suggest as well that the deposit is relatively young. The average of ^{36}Cl ages, excluding VB3c as an outlier, is 1.90 ± 0.45 ka. The uncertainty of the mean is based on the cumulative probability of uncertainties for all samples based on a Gaussian probability distribution (one sigma level). Such a value indicates that the rock avalanche, considering the error range, occurred during historical times, between 340 BC and 560 AD. These results are in stark contrast to previous reconstructions, which pointed to a Lateglacial age (Mazzuoli, 1875; Hoernes, 1892; Squinabol, 1902; Dal Piaz, 1912; Venzo, 1939; Genevois et al., 2006; Pellegrini et al., 2006). The date 1113, 1114, 1117 AD proposed for the main landslide event suggested by some authors (Piloni, 1607; Miari, 1830) may be associated to the Verona earthquake at 1117 AD. That event was clearly felt in the Belluno area, inducing numerous landslides (Guidoboni et al., 2005), but its age is not consistent with the cosmogenic dates on the Masiere di Vedana. To search for independent constraints for the age of the main landslide event, a detailed research in numerous archives and chronicles was undertaken. This area during Roman time was largely and uniformly inhabited by “incolae” for agricultural aims, being the area located next to the Claudia Augusta Altinate road connecting Feltre with Belluno (Fig. 1) (Alpago-Novello, 1957, 1988). The presence of a Roman bridge crossing the Cordevole River north of the Mt. Peron indicates there was a connection to the main Claudia Augusta Altinate road. While numerous archeological Neolithic and Roman sites are reported around the Masiere di Vedana (Capuis et al., 1988; Frassine et

al., 2016), no Roman or pre-Roman archaeological remains have been found within the rock avalanche deposits (Fig. 3). If the Masiere di Vedana deposit was settled after Roman times, previous settlements eventually located in the area would have been buried by the event. The oldest record for the post-event human presence is a hospice built in the 12th century AD (1155 AD; Magoga and Marin, 1998) on the fluvial terrace near the village of S. Gottardo (Fig. 3). Therefore, historical data indicate a
285 time frame between late Roman times and the early Middle Ages.

The age uncertainties do not allow to directly determine if the Masiere di Vedana deposit was due to a single failure or multiple events. The distribution of the available dates has no spatial pattern across the deposit, no physical boundaries occur, no buried soil layers have been found within the deposits. Moreover, the volume and H/L ratio of the landslide (see Section 5.2 for further discussion), together with the morphology of the scarp and the absence of secondary scarps, indicate a single
290 huge catastrophic event. Therefore, a single rock avalanche occurred in historical time contradicting previous interpretations (e.g., Genevois et al., 2006; Pellegrini et al., 2006). Such a young age implies that glacial unloading was not responsible for the destabilization of the southern side of the Mt. Peron and asks for a re-evaluation of the involved driving factors and potential triggers.

5.2 Release, emplacement and post-event modification of the deposit

295 A schematic reconstruction of the reach of the Cordevole Valley involved in the rock avalanche can be depicted before, during and after the event. Before the rock avalanche, the Cordevole River flowed through a gentle rolling landscape along the foot of Piz Vedana and through a breach cutting the Castel Cuch bedrock ridge just north of the village of Mas (Figs. 3, 7a). Topographic highs, like Castel Cuch and Roe Alte, were at that time mantled with glacial sediments attributable to the last glaciation.

300 The rock avalanche involved the detachment of about 130 Mm³ from the southern face of Mt. Peron. Initial movement was sliding along the NW vergent backthrust-related planes (Fig. 2). En bloc movement may have been only briefly sustained as the pervasive network of fractures favored a massive collapse. The rock mass immediately evolved into a rock avalanche whose volume increased by fragmentation up to 170 Mm³ and spread out onto the flat plain below. The H/L ratio of ~0.2 (apparent friction angle of 11°) and comparison with empirical and modeling plots of H/L vs. volume (e.g., Spreafico et al., 2018; Aaron
305 and McDougall, 2019) mark the Mt. Peron event as extremely mobile. As a basis for comparison, the Fernpass rock avalanche has a H/L of 0.9, a volume 1 km³ and a significantly longer runout distance of 15.5 km (Prager et al., 2009). It may be possible to glean information about the failure style from the distribution of boulder lithologies, which follows the stratigraphic order of the bedrock exposed in the source area. This has been as well noted at the Tschirgant rock avalanche deposits in Austria (Dufresne et al., 2016) and the Frank slide in Canada (Charrière et al., 2016). In the Mt. Peron bedrock, the lithologic sequence
310 from west to east is: Calcari Grigi Group, Vajont Limestone, Fonzaso Fm., Rosso Ammonitico and Maiolica. This pattern is mirrored in the deposits (Figs. 3, 4): Vedana and Torbe are dominated by Calcari Grigi Group, Masiere by Vajont Limestone and Roe by Fonzaso Fm., Rosso Ammonitico and Maiolica. Experiments and modelling suggest that this kind of zonation is likely to occur when the sliding mass propagates as a flexible sheet, with laminar flow (Friedmann et al., 2006).

Several landforms within the Masiere di Vedana provide further clues on the processes of propagation and emplacement.

315 The tomas in the Torbe sector suggest differential velocities in the moving mass propagating on a water-saturated substrate (Strom, 2006; Prager et al., 2009; Dufresne, 2012; Dufresne et al., 2016; Aaron et al., 2017). Tomas, with likely similar origin, are present in the distal deposits at Fernpass in Tyrol (Prager et al., 2009) and at Flims (von Poschinger and Ruegg, 2012). Recently, More and Wolkersdorfer (2019) proposed for the Toma Hills at Fernpass an alternative origin from internal erosion by suffusion. However, at Masiere di Vedana the fluvial deposition above the rock avalanche suggests that the suffusion process

320 can be ruled out. In contrast to the increased mobility seen in the Torbe sector, in the central Masiere area, landforms indicative of stalling are present (Fig. 6). The stacked sub-parallel transverse ridges, much like those noted at Tschirgant (Patzelt, 2012; Dufresne et al., 2016; Ostermann et al., 2017), with slight overrunning of the ridges in front by those behind, indicate slowing down of the moving mass due to longitudinal compression (Nicoletti and Sorriso-Valvo, 1991; Dunning et al., 2005; Strom, 2006; Dufresne et al., 2016). Outcrop relationships (Fig. 6) suggest that the Pleistocene conglomerate inhibited the rock mass

325 flow, in combination with the slight uphill gradient. The ridges at Masiere were previously interpreted as neotectonic lineaments by Baggio and Marcolongo (1984).

After the event, the Cordevole River changed its channel several times. The rock avalanche blocked the river, creating accommodation space to the north, where possibly a temporary lake formed. The river was then forced to flow westward across the deposit as indicated by the paleochannels in the Vedana and Torbe sectors (black arrows, Fig. 3), taking different

330 paths at different times. Low-lying areas were progressively filled, as shown by the fining upwards sequence recorded in core TB1 (Supplementary Material Figure SM2a). The Torbe, Vedana and Peron terraces are flat surfaces at ~380 m a.s.l. (Fig. 7b). Afterwards, a further sedimentation was hindered by the trenching in Torbe. The Cordevole river finally breached the landslide deposit to the southeast, through the Castel Cuch ridge made of Cenozoic rocks (Fig. 7a). The river initially flowed from Mas (about 375 m a.s.l.; green line in Fig. 7a) to the southern flank of Castel Cuch as suggested by the still recognizable

335 paleochannel (Fig. 3) filled with well sorted, medium- to coarse-grained sand (Caneve, 1985). Subsequently, the Cordevole moved to the eastern side of the Pleistocene conglomerate cliff (Fig. 3). The final diversion of the river formed the Peron, Mas and Vignole terraces and currently flows some meters below with upstream migration of the knickpoint (Fig. 7).

5.3 Driving factors and potential hazard

In the Cordevole and Piave Valleys many landslides have been recorded (Fig. 8) and have caused a great deal of damage

340 and casualties (Rossato et al., 2018). ~~Moreover, rock~~ Rock avalanches such as Masiere di Vedana are difficult to predict (Hung, 2006) and may be very destructive due to their huge volume and extreme runout (Guzzetti, 2000; Hung, 2004; Geertsema et al., 2006; Evans et al., 2007; Sosio et al., 2008; Cui et al., 2011; Hermanns and Longva, 2012). In ~~the~~ light of the results we obtained, the search for the drivers of the Masiere di Vedana rock avalanche is both timely and imperative. Even if what determines the moment of failure may be difficult to pinpoint, increased pore pressure and seismic ground shaking

345 are the primary candidates in such cases (Wieczorek, 1996; Schuster and Wieczorek, 2002; Takahashi, 2001). However, rock avalanches may start without a definite external trigger, the progressive accumulation of rock fatigue being enough to overcome the resisting forces of the rock mass. This is the case, for example, of the Tsatichhu landslide (10th September 2003) in Bhutan

(Dunning et al., 2006) and the several Randa events (total of 30 Mm³) in 1991 in Switzerland (Loew et al., 2012; Stead and Eberhardt, 2013).

350 The Belluno area has a high mean annual rainfall (1643 mm in the time interval 1994-2018 at <https://www.arpa.veneto.it/dati-ambientali/open-data>) and is prone to extreme rainfall events (e.g., >300 mm of rain at Sospirolo during a single event: 27th October – 1st November 2018; ARPAV, 2018). Moreover, at the time of the Masiere di Vedana rock avalanche, soon after the beginning of the Christian Era, the ~~eastern European Alps and NE~~ south-eastern Alps and north-eastern Italy were affected by various periods of climate degradation during which several extreme meteorological events occurred (Wirth et al., 2013; Rossato
355 et al., 2015). One of these extreme events had an impact all over Europe between 50 and 250 AD, with marked intensity and widespread flooding recognizable in the stratigraphic records (~~(Macklin et al., 2006; Benito et al., 2015; Rossato et al., 2015)~~
(Macklin et al., 2006; Benito et al., 2015; Rossato et al., 2015; Cremonini et al., 2013)). This period of severe rainfall could possibly have been the trigger for the Masiere di Vedana rock avalanche or, at least may have acted as a driving and destabilizing factor. Likewise, the 1987 Val Pola rock avalanche in the central Alps was triggered by a period of exceptional rainfall (Crosta
360 et al., 2004).

The Veneto region is prone to earthquake activity and the study area is categorized as level 2 seismic hazard (“possible strong earthquakes” in Ordinanza del PCM n. 3519/2006), as the historical record testifies (up to ~Mw=6.5; Vigano` et al., 2013, 2015; Rovida et al., 2016). Continuous instrumental monitoring of the Belluno area dates back only to 1977 (Sandron et al., 2014), ~~preceding~~ Preceding cataloged major seismic events in the region are based on either historical chronicles, dated building damages and/or observed rockfalls (Piloni, 1607; Taramelli, 1883; Guidoboni et al., 2005, 2018). Within
365 a radius of 30 km from Mt. Peron, eight earthquakes with Mw greater than 5.0, and one exceeding 6.0, ~~are listed were~~ recorded (Fig. 1). The strongest (Mw = 6.3) ~~in the nearby area~~, occurred just 20 km to the east of Mt. Peron on 29th June 1873 (Rovida et al., 2016, and references therein). Severe ~~damages damage~~ to Belluno city ~~were was~~ reported during the Asolo (25th February 1695 AD; Mw=6.4), Friuli (6th May 1976 AD; Mw=6.5) and Verona (3rd January 1117 AD; Mw=6.5)
370 earthquakes whose epicenters were located, respectively, 60, 65 and 140 km away (Guidoboni et al., 2018). As for the time frame suggested by our chronology, historical records report an important seismic event at in July 365 AD ~~with damages~~ that resulted in damage to the city of Belluno (Piloni, 1607). These data suggest that the Belluno area is sensitive to seismic shakings originating even hundreds of km away. ~~In the Alps, earthquakes have been suggested as triggers for several rock avalanches (e.g., Grämiger et al., 2016; Ivy-Ochs et al., 2017; Köpfl et al., 2018)~~ Galadini et al. (2005) discuss evidence
375 that active tectonics plays a key role in the reported intensification of slope instability registered in this area during the last 1500 yr. The most important effect of the frequent seismic activity ~~, even of markedly different magnitude,~~ is the progressive increase in the rock fatigue, with the formation of failure surfaces and the removal of rock bridges and roughness on discontinuity planes (Friedmann et al., 2003; Brideau et al., 2009; Parker et al., 2013; Stead and Eberhardt, 2013; Preisig et al., 2015; Gischig et al., 2016). ~~Where these intersect, rock dissolution and the formation of caves is favored (Filipponi et al., 2009; Sauro et al., 2013)~~
380 ~~, further weakening the mechanical properties of rocks (Pánek et al., 2009; Gutierrez et al., 2014). The Earthquakes have been suggested as triggers for several Alpine rock avalanches (e.g., Prager et al., 2009; Grämiger et al., 2016; Ivy-Ochs et al., 2017; Köpfl et al. and are known to have caused several rockfalls in the Belluno area (e.g., Piloni, 1607; Miari, 1830; Guidoboni et al., 2005).~~

Moreover, they are considered as a possible trigger for failure of some pillars located on the Mt. Peron southern wall is known locally as the "weeping rock" due to the numerous caves and karst springs along the steep rock face (Fig. 2b). side (Di Giusto, 2012).

Where a structural setting similar to that at The predisposition of the Mt. Peron is present, the occurrence of huge landslide events deserves evaluation. The southwest face to failure is attributable to the structural setting and the regional-scale framework. The entire Belluno Dolomites experienced a long deformation history since the Miocene (Doglioni, 1990), related to regional-scale stress connected to the counter-clockwise rotation of the Adria plate, indented with the Alpine orogeny (Márton et al., 2003; D'Agostino et al., 2008). Such forces overturned the bedding, formed the The thrusts and backthrusts (WSW-ENE oriented) that encase Mt. Peron relate to this phase of activity, also created the two conjugate fracture sets (NW-SE oriented) and led to re-activation induced the reactivation of the Jurassic faults (N-S oriented). The belt characterized by these deformations, where the Mt. Peron is located, lies and the overturning of the beds. The area between the Belluno thrust and the to the south, the Val Carpenada - Val di Vido - Val Madonuta backthrust (Fig. 1), and extends from the Piave to the north, the Caorame Valley to the east to the Caorame west and the Piave Valley to the west (Bosellini et al., 1981; Masetti and Bianchin, 1987; Bigi et al., 1990; Costa et al., 2003). The Belluno Dolomites are also seismically prone and active tectonics has been suggested to contribute to intensification of slope instability registered in this sector during the last 1500-yr (Galadini et al., 2005). Moreover, the area is densely inhabited east (Fig. 1) is structurally homogeneous and characterized by the same fracture sets identified at Mt. Peron (Bosellini et al., 1981; Masetti et al., 2003). In limestones, where these intersects, rock dissolution and caves can form (Filipponi et al., 2009; Sauro et al., 2013), further weakening the rock (Pánek et al., 2009; Gutierrez et al., 2014). The Mt. Peron southern wall is known locally as the 'weeping rock' due to the numerous caves and karst springs along the steep rock face (Fig. 9), and some artificial lakes are present (e.g., Lake Mis; Fig.3). A massive rock failure that would hit such lakes or damage the dams may pose a serious threat, possibly triggering a tsunami, as happened for instance at Vajont (e.g., ?)2b).

Our dating of the Masiere di Vedana landslide to late Roman times casts considerable doubt on the previous Lateglacial chronological assessment (e.g., Pellegrini et al., 2006). Consequently, stress release following downwasting of the LGM Piave glacier, that had been previously considered as the main driver of the rock avalanche, actually played no direct role in the process. Climatic and tectonic factors were much more important. In addition, the classification of the deposits as Lateglacial led to underplaying of the possible hazard at Mt. Peron. The situation at Mt. Peron, with steeply dipping-to-overturned bedding in a limestone massif crisscrossed by numerous faults and riddled with karst fissures, is presently the same that produced the massive failure of the Masiere di Vedana. Several hundred-meter-high rock prisms along the top of the headscarp are partially detached along the discussed fracture systems and loom perilously over the inhabited valley below (Di Giusto, 2012). Our results suggest the need for a reconsideration of the hazard related to not only the Mt. Peron southern face, as the rock slopes still present evident structural weakness, but also the whole area lying between the Belluno thrust and its backthrusts, as intense rainfall and earthquakes can occur at any time.

Data acquired in this study provide a new contribution to the knowledge of the timing, failure and propagation of the Masiere di Vedana rock avalanche. The rock avalanche ($\sim 130 \text{ Mm}^3$) detached from the southern slope of the Mt. Peron. The deposit extends over an area of 9 km^2 , with a total volume of $\sim 170 \text{ Mm}^3$. A H/L ratio ~ 0.2 is calculated, marking it as extremely mobile, which is also shown by the maximum runout of 6 km. Geomorphological, stratigraphic and historical evidence when
 420 combined with cosmogenic ^{36}Cl exposure ages, mean age $1.90 \pm 0.45 \text{ ka}$, point to a single event that occurred in or after late Roman times but before the early Middle Ages.

The steep rock wall on the south face of Mt. Peron shows a pervasive deformation; numerous fractures and faults cross-cut the sub-vertical to slightly overturned carbonate Mesozoic bedrock. The WSW-ENE directed backthrust planes, which are the most continuous ones, constituted the planes along which the rock mass initially slid, rapidly breaking-up and evolving into a
 425 rock avalanche.

The stratigraphic sequence is preserved in the rock avalanche deposit. Lithologies that presently constitute the western part of the source area, were deposited in the proximal sectors (Vedana, Torbe), while the more easterly outcropping ones reached the distal areas (Masiere, Roe Alte). Landforms of the deposit suggest differential velocities during emplacement. In the NW sector (Torbe) enhanced mobility likely due to interaction with water-saturated path material is evidenced by the numerous
 430 ENE-WSW aligned tomas. In contrast, in the middle sector (Masiere) stacked transverse ridges point to stalling, perhaps related to the gentle uphill gradient and impeded propagation over Pleistocene conglomerates. Post-event evolution comprises formation of backwater alluvial terraces and the wandering of the Cordevole River in the rock avalanche deposits, with incision and aggradation phases.

Identified pivotal drivers are the overall structural setting, exceptional rainfall events and seismic shakings. Their combination produced a pervasive fracturation and weathering of the rock mass, with progressive increase of rock fatigue. No
 435 exceptional event may actually be required for such rock avalanches to occur, as accumulation of damage markedly lowers the energy needed to trigger failure.

~~In the light of the~~ Considering the new data provided, we ~~suggest considering a re-evaluation of the landslide hazard in the highlight that the same driving factors responsible for the Masiere di Vedana rock avalanche are still present today at Mt.~~
 440 Peron. The area between the Belluno thrust and its backthrusts, and from the Caorame to the Piave Valleys -is structurally analogous to Mt. Peron, and therefore re-evaluation of the landslide hazard may be warranted. The possible occurrence of huge rock avalanches and minor rockfalls is a scenario that is necessary to take into consideration in future hazard evaluation and mapping.

Data availability. All data are in the paper or in the supplemental material.

445 *Author contributions.* All authors contributed to discussion, field survey, data collection and improving the text, that has been written mostly by S. Rossato, S. Ivy-Ochs, S. Martin and G. Monegato. Each author contributed to different parts, here listed: geomorphology: S. Rossato, S. Ivy-Ochs, G. Monegato, M. De Zorzi, N. Surian, P. Mozzi; geological and structural analysis: S. Martin, A. Viganò, P. Campedel; remote sensing and GIS elaborations: S. Rossato; dating: S. Ivy-Ochs, C. Vockenhuber, S. Martin; thin sections analysis: M. Rigo.

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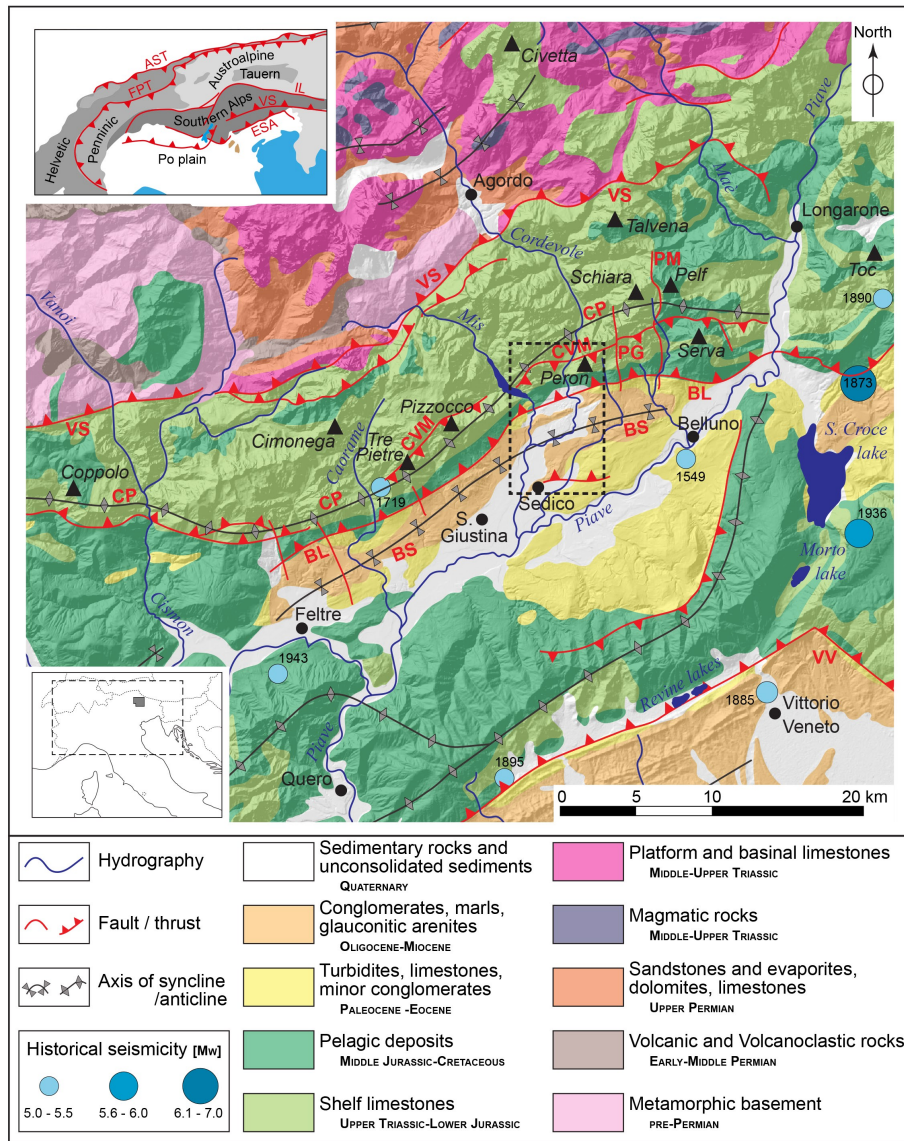


Figure 1. Simplified regional geological map (black dashed-line box indicates location of Fig. 3). The map is based on Bosellini et al. (1981); Masetti and Bianchin (1987); Bigi et al. (1990); Costa et al. (1996). Epicenters of earthquakes from the last two millennia are shown with color coding and year of occurrence (source: Parametric Catalogue of Italian Earthquakes, 2015 version Rovida et al., 2016). Base map is the SRTM derived Digital Elevation Model (30-m cells) (source: <http://viewfinderpanoramas.org/>). Structural setting is shown in inset upper left (based on Doglioni, 1990). AST: Alpine Sole thrust; BL: Belluno thrust; BS: Belluno syncline; CP: Coppolo-Pelf anticline; CVM: Val Carpenada-Val di Vido-Val Madonuta thrust; ESA: eastern Southern Alps thrust system; FPT: Frontal Penninic thrust; IL: Insubric line; PG: Pala Alta-Gresal; PM: Pala Bassa-Val Medone; VS: Valsugana thrust; VV: Vittorio Veneto thrust.

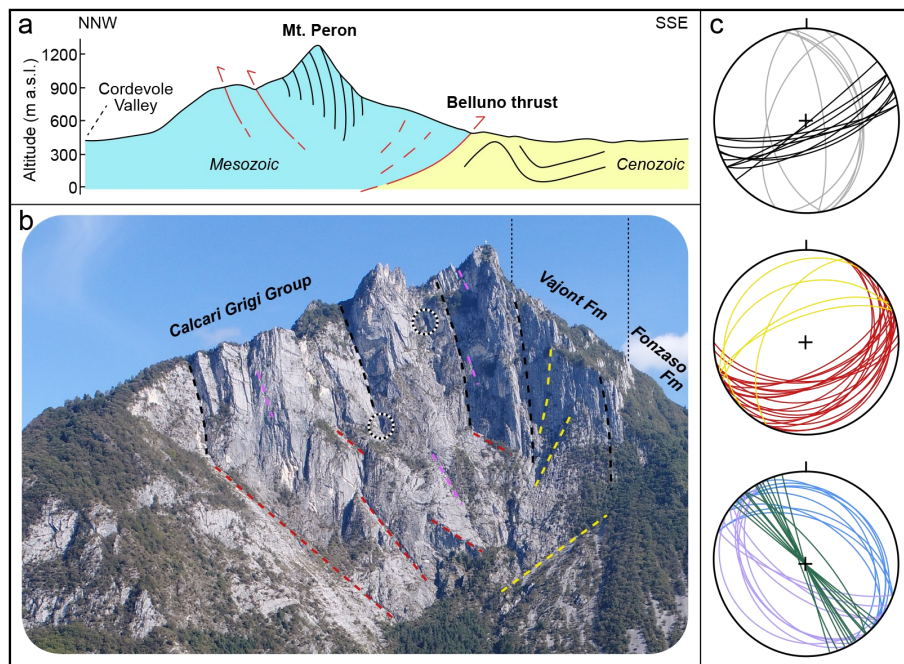


Figure 2. Structural scheme of the Mt. Peron release scarp. **a:** Schematic geological cross-section. **b:** Photograph with major structural elements and larger karst caves (black-and-white circles) highlighted. **c:** Lower hemisphere stereographic projection of principal structural elements. Colours in b and c correspond to the following: bedding (black), N-S reactivated Jurassic fault-related planes (grey), basal trenches and other high-angle fractures connected to the Belluno thrust (yellow) and backthrust (red), fractures related to the NNW-SSE fault system (light blue), their conjugates (pink) and high-angle fractures with the same orientation (green).

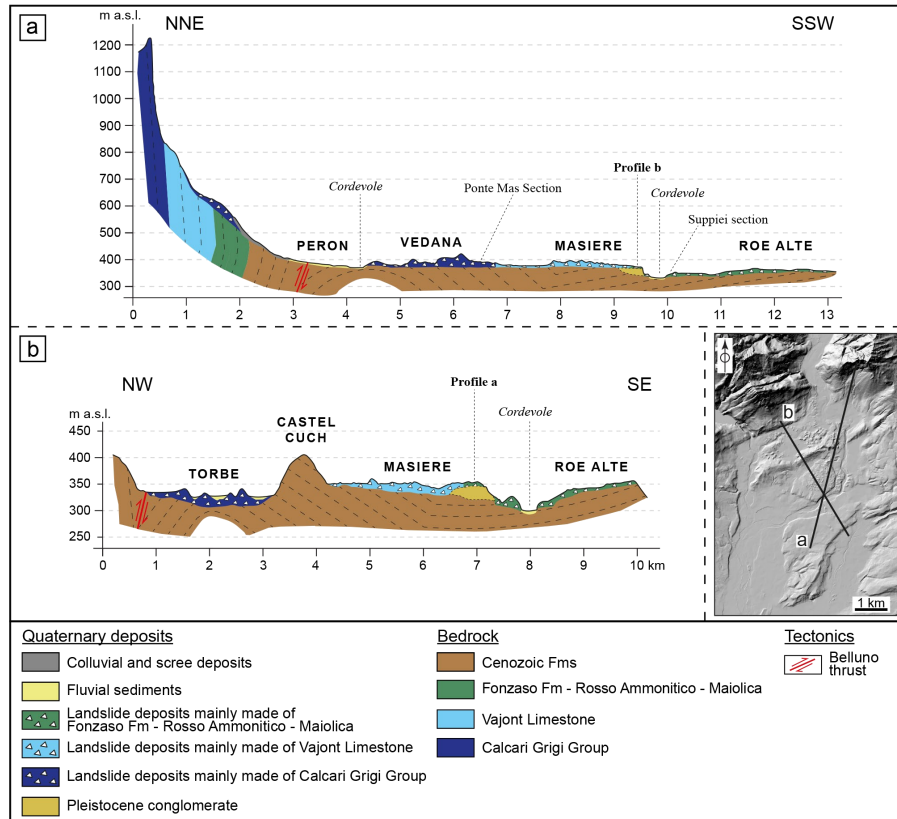


Figure 4. Geologic profiles of the Mt. Peron rock avalanche deposits (modified from Costa et al., 1996). Traces of the profiles are shown in the small DTM and in Fig. 3a, their extent being equal. The discussed sectors of the Mt. Peron rock avalanche, Vedana, Torbe, Masiere and Roe Alte, are indicated, along with the location of Ponte Mas and Suppieri stratigraphic sections (Fig. 5g, h). The vertical bedding and interpreted Belluno thrust are shown schematically. Note the preservation of bedrock stratigraphic order in the rock avalanche deposits from older to younger, i.e. from proximal to distal: Calcari Grigi Group, Vajont Limestone, Fonzaso Formation-Rosso Ammonitico-Maiolica.

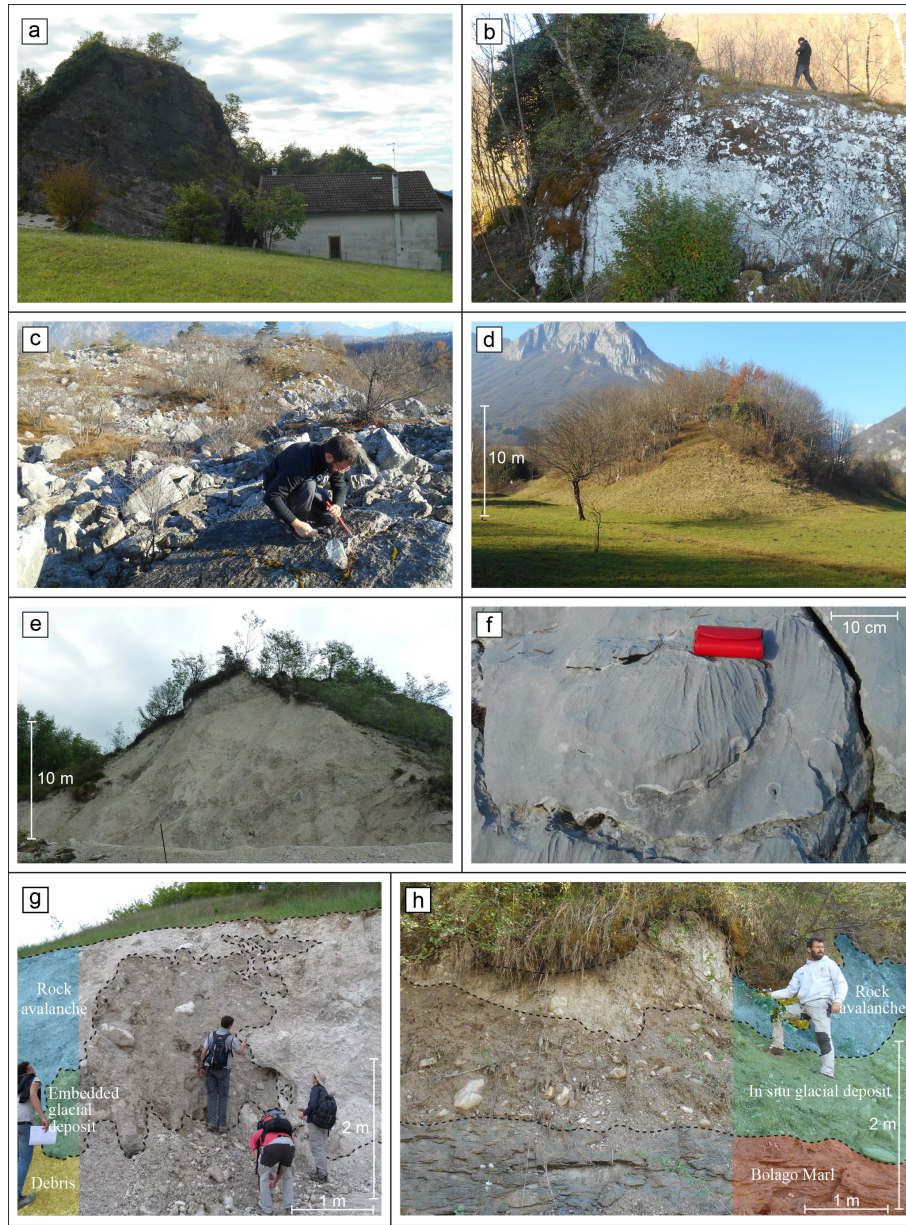


Figure 5. Photos of the deposits (**a**: decametric boulder, Peron alluvial terrace; **b**: plurimetric boulder, Vedana sector; **c**: metric boulders, Masiere sector; **d**: toma relief in the Torbe sector; **e**: open section of a toma, Torbe sector; **f**: karst evidence on a boulder, Masiere sector) and stratigraphic sections described in the text (**g**: Ponte Mas section: incorporation of the underlying glacial deposit (green) into the rock avalanche deposit (light blue) is shown; **h**: Suppei section: glacial sediments (green), covered by rock avalanche deposit (light blue), rest directly on bedrock (red); locations in Figs. 3, 4).



Figure 6. Transverse ridges in the south-western part of the Masiere sector (basemap source: Google Earth). Location of image shown in Fig. 3. The rock avalanche moved (from north to south) over a slight topographic high made up of Pleistocene conglomerate (see Fig. 4), the ridges are interpreted as compressional. White stars indicate locations where the contact between the rock avalanche and the conglomerate is exposed.

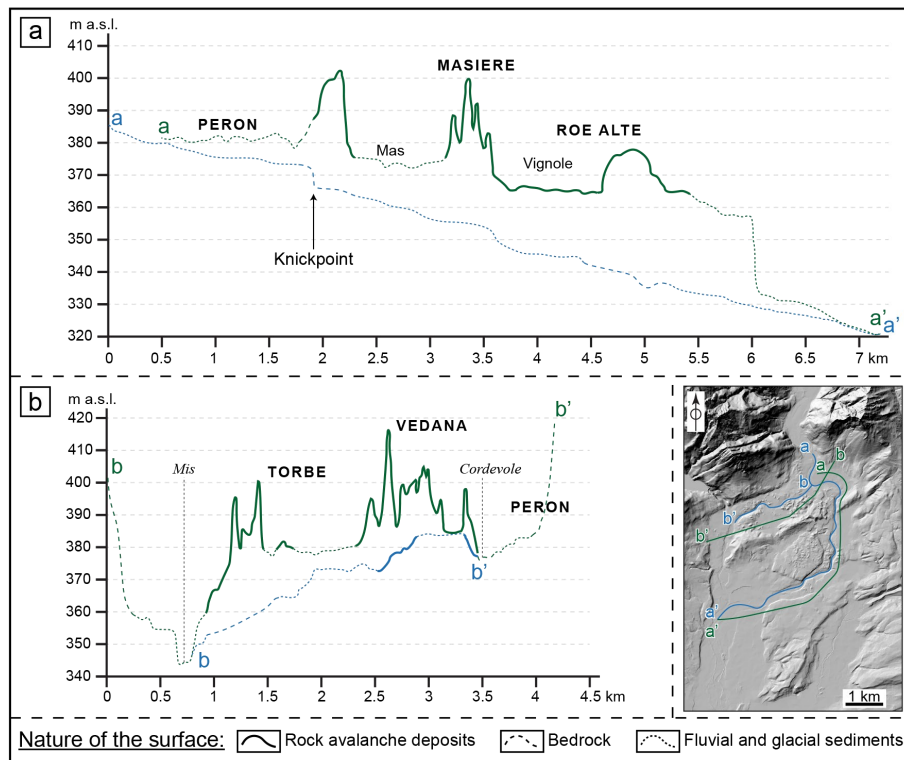


Figure 7. Curvilinear topographic profiles (10X vertical exaggeration) showing the relationship between the bigger incisions and the external main surfaces related to the rock-avalanche and the post-event evolution. Profiles correspond to the present course of Cordevole creek (**blue line in a**) paired with the main topographic surfaces on its left bank (**green line in a**) and the main incision passing through the Torbe and Vedana sectors (**blue line in b**) paired with the main topographic surfaces on its left side (**green line in b**). Traces of the profiles are shown in the small DTM, that corresponds to Fig. 3.

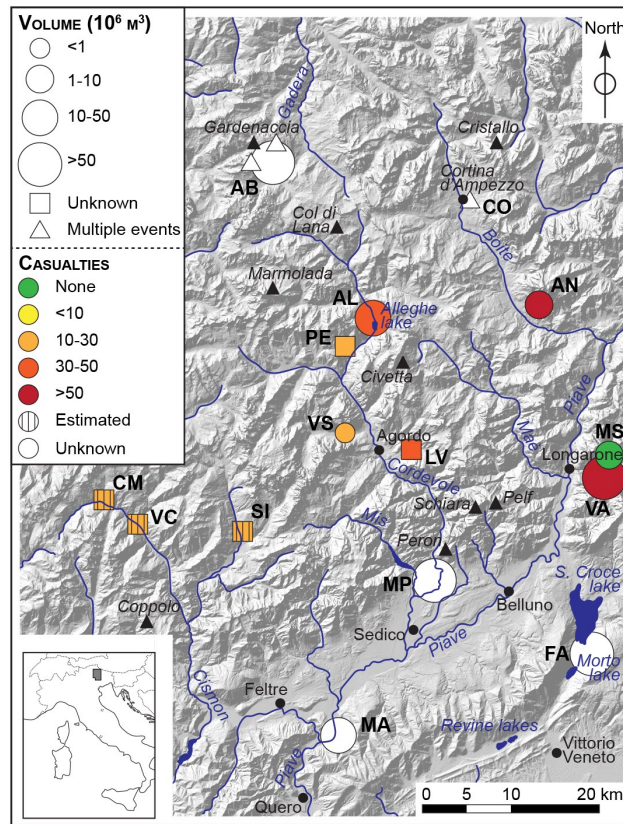


Figure 8. Largest and most damaging (in terms of human lives) landslides located near Mt. Peron (base map is a SRTM derived Digital Elevation Model with 30-m cells; source: <http://viewfinderpanoramas.org/>). Volumes of deposits correspond to the size of the symbols, casualties are shown with color coding. AB: Alta Badia, several events between 11,500 BP-present (Borgatti et al., 2004), AL: Alleghe, 1771 AD (Ermini and Casagli, 2003), AN: Antelao, 1814 AD (Montandon, 1933), CM: Col Mandro, 1825 AD (Montandon, 1933), CO: Cortina d'Ampezzo, several events between 10,700-2,000 BP (Borgatti and Soldati, 2010), FA: Fadalto, Lateglacial-to-present (Pellegrini and Surian, 1996), LV: La Valle. 1701 AD (Montandon, 1933), MA: Marziai, 17,500-15,000 BP (Pellegrini et al., 2006), MS: Mt. Salta, 1674 AD (Montandon, 1933), PE: Pecol, 1841 AD (Montandon, 1933), MP: Mt. Peron, 1.90 ± 0.45 ka - late Roman times - early Middle Ages (this work), SI: Siror, 1348 AD (Montandon, 1933), VA: Vajont, 1963 AD (Borgatti et al., 2004), VC: Val Cia, 1882 AD (Montandon, 1933), VS: Valle San Lucano, 1908 AD (Aldighieri et al., 2016).

Panoramic view of the northern side of the Piave Valley, with major peaks belonging to the studied deformation belt.

Table 1. Sample site information, AMS data and ^{36}Cl exposure ages. Samples VB3a and VB3c are not from the same boulder, whilst VB13a and VB13b are (weighted mean given). No erosion correction was made.

Sample name	Lithology	Boulder height (m)	Latitude	Longitude	Elevation (m a.s.l.)	Shielding*	Thick. (cm)	Cl (ppm)	^{36}Cl (104 atoms/g)	Age (ka)
VB2	Calcari Grigi	6	46.1593	12.1202	395	0.986	2	32.7 ± 1.1	4.37 ± 0.75	1.49 ± 0.26
VB3a	Rosso Ammonitico	2	46.1686	12.1253	520	0.526	3	20.5 ± 0.3	3.22 ± 0.48	1.83 ± 0.28
VB3c	Fonzaso	20	46.1686	12.1253	530	0.518	2	20.3 ± 0.1	6.29 ± 0.69	3.62 ± 0.41
VB12	Vajont	1.5	46.1261	12.1167	392	0.993	7	29.6 ± 0.2	7.14 ± 0.56	2.34 ± 0.21
VB13a#	Vajont	5	46.1261	12.1167	396	0.935	2	18.9 ± 0.1	3.87 ± 0.28	1.45 ± 0.12
VB13b#	Vajont	5	46.1261	12.1167	396	0.955	2.2	20.4 ± 0.2	4.01 ± 0.30	1.45 ± 0.12
VB14	Calcari Grigi	15	46.1683	12.1234	470	0.555	5.5	8.0 ± 0.1	5.15 ± 0.83	2.39 ± 0.38

*=Shielding includes surround topography and dip of sampled surface. #=Sample from same surface on same boulder, mean age is 1.45 ± 0.08 ka.