

# Structural weakening of the Merapi dome identified by drone photogrammetry after the 2010 eruption

Herlan Darmawan<sup>1,2\*</sup>, Thomas R. Walter<sup>1</sup>, Valentin R. Troll<sup>3,4</sup>, Agus Budi Santoso<sup>5</sup>

5 <sup>1</sup>Dept. Physics of Earth, GFZ German Research Center for Geosciences, Telegrafenberg, 14473, Potsdam, Germany.

<sup>2</sup>Laboratory of Geophysics, Department of Physics, Faculty of Mathematics and Natural Sciences, Universitas Gadjah Mada, Yogyakarta, Indonesia.

<sup>3</sup>Dept. of Earth Science, Section for Mineralogy, Petrology and Tectonics, Uppsala University, Villavägen 16, SE-752 36 Uppsala, Sweden.

10 <sup>4</sup>Faculty of Geological Engineering, Universitas Padjajaran, Jatinangor 45363, Bandung, Indonesia.

<sup>5</sup>BPPTKG (Balai Penyelidikan dan Pengembangan Teknologi Kebencanaan Geologi), Jalan Cendana 15, Yogyakarta 55166, Indonesia.

15 *Correspondence to:* HerlanDarmawan (herlan@gfz-potsdam.de; herlan\_darmawan@mail.ugm.ac.id)

**Abstract.** Lava domes are subjected to structural weakening that can lead to gravitational collapse and produce pyroclastic flows that may travel up to several kilo meters from a volcano's summit. At Merapi volcano, Indonesia, pyroclastic flows are a major hazard, frequently causing high numbers of casualties. After the VEI 4 eruption in 2010, a new lava dome developed on Merapi volcano and was structurally destabilized by six steam-driven explosions between 2012 and 2014. Previous studies revealed that the explosions produced elongated open fissures and a delineated block at the southern dome sector. Here, we investigate the geomorphology, structures, thermal fingerprint, alteration mapping and hazard potential of the Merapi lava dome by using drone-based geomorphologic data and forward-looking thermal infrared images. The block at the southern dome of Merapi is delineated by a horseshoe-shaped structure with maximum depth of 8 m and it is located on the unbuttressed southern steep flank. We identify intense thermal, fumarole, and hydrothermal alteration activities along this horseshoe-shaped structure. We conjecture that hydrothermal alteration may weaken the horseshoe shaped structure, which then may develop into a failure plane that can lead to gravitational collapse. To test this instability hypothesis, we calculated the factor of safety, and ran a numerical model of block-and-ash flow using Titan2D. Results of the factor of safety analysis confirm that intense rainfall events may reduce the internal friction and so gradually destabilize the dome. The titan2D model suggests that a hypothetical gravitational collapse of the delineated unstable dome sector may travel southward for up to 4 km distance. This study highlights the relevance of gradual structural weakening of lava domes, which can influence the development fumaroles and hydrothermal alteration activities of cooling lava domes for years after initial emplacement.

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**Keywords:** Merapi volcano, unstable dome sector, geomorphology, thermal mapping, factor of safety, pyroclastic hazard.

## 1. Introduction

Lava domes are viscous lava **extrusions** that accumulate at volcanic vents and experience **exogenous** and **endogenous** growth (Hale, 2008). During formation of lava domes, they may start lateral flow as coulees, be subject to cooling **and** subsidence, **and can** develop concentric **fractures** on the flat-topped summit of the dome (Walter et al., 2013b; Salzer et al., 2017; Rhodes et al., 2018). **Many** details of the development, geometric organization, and actual **formation** processes of dome structures remain poorly understood. **External** factors such as intense rainfall, hydrothermal alteration, gas overpressure, mechanical weakening, and earthquake may **further** augment instability and promote a dome **collapse** (Voight and Elsworth, 2000; Reid et al., 2001; Ball et al., 2015). Once fracture arrangements are established in a lava dome, **volcanic gas and** rain water **are** able to **flow that may cause** hydrothermal alteration and gas overpressure **along the structure, which may lead** to dome **destruction** even during **quiescent** periods (Voight and Elsworth, 2000; Reid et al., 2001; Elsworth et al., 2004; Simmons et al., 2004; Taron et al., 2007; Ball et al., 2015). **Structural weakening and instability of a lava dome due to these** processes may **then** cause hazardous **rock falls or** block-and-ash flows (Calder et al., 2015).

The dome collapse at Soufrière Hills Volcano (SHV), Montserrat **in 1998-99**, is an example of rain-triggered collapse that followed a period of quiescence. The SHV dome collapse produced pyroclastic density currents (PDC) with a volume of  $22 \times 10^6 \text{ m}^3$  that **travelled up to ~3 km** distance **from the summit** (Norton et al., 2002; Elsworth et al., 2004). The rain water infiltrated through identified fractures, produced gas overpressure within the lava dome carapace, **and then triggered dome collapses** that were characterized by hydrothermal alteration and **structural instability** (Voight, 2000; Elsworth et al., 2004). This event **demonstrates that identifying a structural weakening is crucial** for volcanic hazard mitigation.

**However, identifying** the potential hazard of lava domes is often difficult and requires high quality observational **datasets** complemented by modelling analyses (Voight, 2000). Dome building volcanoes are often steep sided hazardous edifices, where direct access is very limited and acquisition of high quality field data **is** challenging. **In contrast, remote sensing** techniques, such as satellite imageries, aerial **photogrammetric** and thermal imaging can provide detailed information on the structure, deformation, **geomorphology**, and thermal signature of active lava domes (James and Varley, 2012; Walter et al., 2013a; Salzer et al., 2014; Thiele et al., 2017; Darmawan et al., 2018), which allows to study important parameters for assessment of dome instability and potential hazards (Voight and Elsworth, 2000; Elsworth et al., 2004; Simmons et al., 2004; Taron et al., 2007). **In this respect, the degree of instability of lava domes can be assessed** by using a factor of safety **equation** (Voight and Elsworth, 2000; Simmons et al., 2004; Taron et al., 2007). **Factor of safety (FS) is widely used to calculate slope stability** (Bishop, 1955) and it is calculated by dividing resisting forces to driving forces that act on a failure plane ( $\alpha$ ) (Voight, 2000). A result of  $FS \leq 1$  indicates a failure condition. However, in a lava dome some additional forces may act on a failure plane due to e.g. degassing and rainfall activities (Simmons et al., 2004) and here, we test the first factor of safety model at Merapi lava dome to assess its stability under rainfall conditions.

**In a case of structural dome instability, the hazard arising from dome collapses can be simulated by geophysical mass flow software, such as Titan2D** (Patra et al., 2005; Sheridan et al., 2005). Titan2D is a software to model 2D geophysical mass

flow based on a depth averaged model for an incompressible continuum granular flow and was validated through laboratory experiments (Patra et al., 2005). It is publicly available and has been used to map dynamics and distribution of block-and-ash flows at Merapi during the 2006 and 2010 eruptions (Charbonnier and Gertisser, 2009, 2012; Charbonnier et al., 2013).

In this study, we employed drone photogrammetry and Terrestrial Laser Scanning (TLS), thermal mapping, factor of safety calculation, and Titan2D simulation to assess structural instability and hazards potential of the current Merapi lava dome. Combination of TLS and drone photogrammetry is able to generate a high resolution Digital Elevation Model (DEM) of the Merapi summit, which compares favourably to the satellite-based DEM. For the first time, we are now able to generate realistic model of the morphology and structure at Merapi summit. Thermal mapping by using a FLIR camera also provides detailed locations of hydrothermal fluid activity. The information of geomorphology, structure, and thermal activity allows us to analyze the factor of safety, and to set up a forward simulation of the Titan2D model. The combined results help to better understand the relevance of dome fracturing, structural weakening, and to outline the potential hazard zone affected in case of a dome sector collapse.

### 1.1. Merapi volcano

Merapi volcano is a basaltic to andesitic volcano that formed due to subduction of the Indo-Australian oceanic plate beneath the Eurasian continental plate (Hamilton, 1979). Merapi volcano is one of the most active and dangerous volcanoes in Indonesia, with more than 1 million people living on the volcano's flanks. Moreover, the city of Yogyakarta with 3 million inhabitants is located only ~30 km distant from the volcano's summit (Fig. 1) (Lavigne et al., 2015). The volcanic activity of Merapi is well documented since the 1800's and its typical eruption style is dome extrusion and block-and-ash flows (Voight et al., 2000). The extrusion rate of a lava dome at Merapi may strongly vary, ranging from ~0.04 m<sup>3</sup>/s (Siswowardoyo et al., 1995), up to 35 m<sup>3</sup>/s or more during e.g. the 2010 volcanic crisis (Pallister et al., 2013).

Merapi shows signs of interactions with surrounding environmental influences, and for instance rainfall appears to correlate with fumarole activity and seismic intensity (Richter et al., 2004), and tectonic earthquakes can influence eruptive activity (Walter et al., 2007; Walter et al., 2015; Carr et al., 2018). The volcano erupted several times during the last decades, on average once every 3-5 years, with the largest explosive event recorded in 2010. The 2010 eruption removed parts of the summit area (Surono et al., 2012), excavated a ~200 m deep crater and was followed by re-growth of a new dome (Kubanek et al., 2015). The new lava dome was intermittently destroyed by several explosive events again between 2012 and 2014, which also caused elongated open fissures (Fig. 1b and c) (Walter et al., 2015), and a horseshoe-shaped structure that highly altered and delineated the southern part of the dome (Fig. 1c) (Darmawan et al., 2018). The horseshoe-shaped structure is posing a safety risk due to weakening from hydrothermal alteration and may possibly collapse in the foreseeable future.

## 2. Data and methods

### 2.1. Observational data

We conducted Terrestrial Laser Scanning (TLS), drone photogrammetry, and thermal infrared field campaigns to investigate geomorphology, structure, hydrothermal alteration, and thermal distribution of the Merapi lava dome. The TLS data was acquired on 18 September 2014 by using a Riegl 6000 instrument from the eastern rim of the summit crater (latitude =  $7^{\circ} 32' 25.0161''$  S, longitude =  $110^{\circ} 26' 51.2110''$  E), looking down westward onto the dome. The TLS instrument was set by using a Pulse Repetition Rate (PRR) of 30 kHz, an observation range of 0.129–4393.75 m, a theta range (vertical) of  $73^{\circ}$ – $120^{\circ}$  and a sampling angle of  $0.041^{\circ}$ , and a phi range (horizontal) of  $33^{\circ}$ – $233^{\circ}$  with a sampling angle of  $0.05^{\circ}$ . We used 12 local reflectors to correct rotation errors. The TLS instrument extracted a 3D point cloud model of the Merapi summit with 2.8 Millions data points. A major benefit of the TLS methodology is the high resolution and precision in the field of view, however shadowing effects are significant.

In order to solve the shadow effects, we applied a Structure from Motion (SfM) technique (Szeliski, 2011) to generate a 3D model based on 2D drone images that were acquired on 6 October 2015. We used a DJI Phantom drone that flew loops at a height of  $\sim 140$  m over the dome and took nadir photographs with 2 seconds regular interval and 12 megapixel resolution. These photographs were processed by using agisoft photoscan professional software to generate a 3D point cloud model of the Merapi summit. We then combined the 3D point clouds of TLS and SfM data by using point pair-picking registration method in Cloud Compare software. More details about the data acquisition and the processing of TLS and SfM data are described in Darmawan et al. (2018). The combined 3D TLS-SfM point cloud was interpolated in ArcMap to generate a Digital Elevation Model (DEM) with resolution of 0.5 m. The DEM was used for geomorphological, topography, and slope analysis.

To further investigate any changes related to structural instability, we conducted drone photogrammetry on 2 September 2017 by using a DJI Mavic pro drone. The drone flew  $\sim 50$  meters above the dome, carried a camera with resolution of 12 Megapixels, and captured around 408 aerial images. However, as strong degassing at the fumaroles limited visibility, 3D point cloud reconstruction by using the SfM-MVS technique was very noisy. The aerial images acquired in 2017 were used to generate photomosaic image and were qualitatively compared to the 2015 aerial images for structural analysis and for alteration mapping.

As we mapped the structural architecture of the lava dome, we are also interested in fluid transport and fumarole activities. Fractures and lithology contrasts may lead to permeability differences that control the pathways of thermal fluids (Ball and Pinkerton, 2006). We recorded apparent temperature distribution of the Merapi lava dome by using a forward looking infrared (FLIR) P660 thermal camera in September 2014. Images were taken from the eastern crater rim close to the TLS station (Fig. 1b). The FLIR camera operates on a spectral band of  $7.5 - 13 \mu\text{m}$  which allows us to identify an apparent temperature range which was calibrated in a range of  $0 - 500^{\circ}\text{C}$ . The resolution of the FLIR cameras is  $640 \times 480$  pixels. The FLIR camera is equipped with a  $7^{\circ}$  ( $f = 131$ ) zoom lens with a 0.38 mrad instantaneous field of view (Walter et al., 2013a), allowing generation of very detailed and high resolution thermal images, with estimated pixel dimensions of 1 px = 0.05 m on the dome center.

Thermal infrared data is dependent on a number of environmental parameters, such as the distance and emissivity of the target (the dome), the solar reflection, the viewing angle, the atmospheric effect, and the presence of particles/gases in the electromagnetic radiation path (Spampinato et al., 2011). We recorded the thermal images during night time (5 am local time), so that background temperature was low, and insulation artefacts and solar reflection were minimized. Other factors were solved in data processing by setting the emissivity and transmissivity values to 0.98 and 0.7, respectively, following Carr et al. (2016) and Ball and Pinkerton (2006). Relative humidity was set to 45% according to weather observation. The relative distance to the dome was on average 300 m and the background temperature was assumed to be 10 °C. After defining the parameters, the thermal images were set to constant color scales for all images, and then were mosaicked to obtain a high resolution panorama image of the apparent thermal distribution of the Merapi lava dome.

## 2.2. Factor of safety (FS)

Factor of safety is widely used to assess slope stability by estimating the load carrying capacity of a flank. The factor of safety describes if a system is stronger or weaker for the given load. It is affected, in our case, by rainfall, and has been applied for numerous engineering problems (Aleotti and Chowdhury, 1999). At dome building volcanoes, the factor of safety calculation allows estimating slope instability during precipitation events, as dome collapse events are favoured by heavy rainfall (Yamasato et al., 1998; Elsworth et al., 2004). Here, we follow the work of Simmons et al. (2004) and test the instability of the southern sector of the Merapi lava dome during intense rainfall by first estimating how deep rain water is able to percolate ( $d$ ) through identified fractures:

$$d = \frac{is^2}{4K_R} \frac{\rho_w c_w}{\rho_R c_R} \frac{\Delta T_w}{\Delta T_R} \times 1.13 \sqrt{t_D} \quad (1)$$

Where  $i$  is the rain intensity as measured by a proximal weather station, and  $\Delta T_w$  is the required thermal energy to vaporize water,  $\Delta T_R$  is the required thermal energy to cool the fracture surface,  $\rho_R$  and  $\rho_w$  are the density of lava dome rock and water, respectively,  $K_R$  is thermal diffusivity,  $t_D$  is a non-dimensional time which is described as  $t_D = K_R t / l^2$ ,  $t$  is the rainfall duration,  $l = s/2$ ,  $c_w$  and  $c_R$  are heat capacity of water and rock, respectively. The estimated of water percolation ( $d$ ) is then used to calculate the factor of safety:

$$FS = \frac{C.s + (W \times \cos(\alpha) - (Fu)) \times \tan(\theta)}{W \times \sin(\alpha) + F_w + F_v} \quad (2)$$

Where  $C$  is the cohesive strength,  $W$  (unstable dome sector weight) =  $s \times h \times \rho_R \times g$ , and  $s$  is the fracture spacing,  $h$  is the unstable dome sector thickness,  $\alpha$  is the inclination of failure plane,  $\theta$  is friction angle, and  $g$  is the gravitational force.

During intense rainfall, rainwater is able to percolate through identified fractures, interacts with the hot interior of lava dome, increases degassing activity and then generates water forces ( $F_w = 0.5 \times d^2 \times \cos(\alpha) \times \rho_w \times g$ ), uplift force from the volcanic gas ( $F_u = 0.5 \times d \times \cos(\alpha) \times \rho_g \times g \times s$ ), and vaporized water force ( $F_v = d \times \cos(\alpha) \times \rho_g \times g \times (h-d)$ ) (Fig. 7a), where  $\rho_g$  is the density of gas. A result of  $FS \leq 1$  indicates a potential failure, whilst for a FS larger than 1 describes a stable condition.

Factor of safety calculation requires careful parameter justification. For the parameters, we consider the rain gauge data that recorded by hydro-meteorological stations around Merapi volcano, and set the rainfall intensity ( $i$ ) to 10-100 mm/h. The fracture spacing ( $s$ ) is 100 m and mimics a translational fault with hanging wall thickness ( $h$ ) of ~40 m (Fig. 7a). The temperature gradient from the surface to the dome interior ( $\Delta T_w$ ) is 200 - 800°C, which is based on our thermal data and thermodynamic models of lava dome interior (Matthews and Barclay, 2004). Friction angle is from 25° to 45°, which is on the range of friction for rock on rock material (Husain et al., 2014; Simmons et al., 2004). Density of Merapi rock is 2242 kg/m<sup>3</sup> (Tiede et al., 2005). As the rock is progressively altered, we assume that the dome rock is homogenous and has cohesion strength of 10 MPa, following studies of rock strength of altered rock from Mayer et al. (2014); Pola et al. (2014). Details of the parameters used to calculate the factor of safety and water percolation are listed in table 1 and a critical discussion of the parameters can be found in section 4.1.

### 2.3. Scenario modeling of block-and-ash flow

Based on analysis of geomorphology, structure, and factor of safety, we are able to identify potential hazards. We then simulated a hazard scenario of gravity driven avalanches by using Titan2D software. The Titan2D software has been used by previous studies to simulate block-and-ash flow due to lava dome collapses (Widiwijayanti et al., 2007; Charbonnier and Gertisser, 2009; Procter et al., 2009; Charbonnier and Gertisser, 2012; Charbonnier et al., 2013). The input parameters of Titan2D should be defined carefully to reduce uncertainty during simulation and to obtain the most realistic result. For parameterization, the volume of the collapse is based on the structurally delineated southern dome sector. The collapse volume is set to  $0.3 \times 10^6$  m<sup>3</sup>, which represents a deep water percolation and gentle slope failure plane ( $\alpha$ ) scenario. The bed coulomb friction parameter, the most sensitive parameter that controls the flow and material distribution (Sheridan et al., 2005; Charbonnier and Gertisser, 2009), is set to between 28° and 16° from the top of the dome to the lowest slope, respectively (Table. 2), following a study of single dome collapses after the June 14, 2006 eruption (Charbonnier and Gertisser, 2009). This range of bed friction parameter will consider the topography effect during simulation and produce realistic mass flow model. The initial velocity is set to 0 m/s as we assume that the failure mechanism is not involving large magmatic pressure. We used the updated Digital Elevation Model of the Merapi summit from our TLS and drone photogrammetry data and extended it in the far field by merging it with the published 2005 digital elevation model (Gerstenecker et al., 2005). A full set of parameters used for Titan2D simulation are listed in table 2 and the limitations of Titan2D are discussed in section 4.1.

## 3. Results

### 3.1. Geomorphology and structure of the Merapi summit

The high resolution slope map and the photomosaic show the geomorphology and structure of the Merapi summit (Fig. 2). The 2010 explosive eruption formed a deep crater that opened to the south east direction and is surrounded by old domes with slopes of 45°. Shortly after the 2010 eruption, a lava dome formed at the middle of the crater. The deep crater is steeply

inclined with a slope of  $\sim 80^\circ$  and has a diameter of 356 m, a maximum depth of 118 m at the north west crater wall, 146 m at the north east crater wall, and 73 m at the south west crater wall as shown in cross section of lines p-q and r-s (Fig. 2a, d, and e). The high resolution drone photomosaic clearly shows that the summit is highly fractured with an azimuth of N150°E and is highly altered, especially around the crater wall (Fig. 2 b and c). A remnant of altered rocks after the 2010 eruption is exposed at the north crater wall and south east basal surface. Degassing activity is identified at the fissure area, southern dome, west crater wall, and north east crater (green points in Fig. 2c). This degassing activity causes progressive hydrothermal alteration up to now that may weaken and destabilize the dome rock. Some of the altered rocks at the crater wall fall and produce gravity driven rock falls that are deposited inside the crater. Some materials of the 2012-14 explosions are also deposited inside the crater and on the top of the lava dome (Fig. 2c).

Further analysis of slope and structure of the lava dome shows that the top of the dome is relatively flat, while the open fissure is steeply inclined with slope of  $\sim 80^\circ$ . A horseshoe-shaped fault-like structure is identified and it delineates a block at the southern dome sector (Fig. 3a). The structure can be traced for a length of over 165 meters and the block has dimension of 100 m x 80 m. Cross section profiles of line k-l and m-n show that the maximum depth of the horseshoe-shaped fault structure in the northwest, northeast, and southwest is 6 m, 8 m, and 3 m, respectively (Fig. 3b and c). The delineated block is steeply inclined at  $\sim 50^\circ$ , hosts abundant fractures, has blocky appearance, and consists of two or three steep regions, which are separated by gently inclined terraces that may indicate different flow unit as also observed from drone aerial image (Fig 3a and 4a). As the unstable dome sector is located on a steep slope (Fig. 3a), it is critical to monitor changes at the southern part of the Merapi dome.

Close range aerial images show more details of the horseshoe-shaped structures and the southern block (Fig. 4). We find five fractures in three different areas (c, d, and e). A closer view of those fractures (1<sup>st</sup>, 2<sup>nd</sup>, 4<sup>th</sup>, and 5<sup>th</sup>) reveals that they have a width of 0.3 - 1.3 meters (Fig. 4c, d, and e). Comparison of drone aerial images between 2015 and 2017 shows a progressive hydrothermal alteration processes around those fractures within just two years. The yellow color surrounding the active fractures indicates sulfur deposit around the fumaroles, which are stronger expressed in the 2017 images, especially around the fracture number 5 (area e-e'). It may indicate a structural weakening due to hydrothermal alteration. The hydrothermal activities at the horseshoe-shaped structure are also observed by our thermal camera, which is described below.

### 3.2. Thermal variation of the Merapi dome

Forward looking infrared thermal mapping allows identification of the apparent temperature of the dome surface and main regions of hydrothermal fluids flow at the horseshoe-shaped structure. We find that the mean apparent temperature at the dome surface is about 6 - 14°C (Fig. 5a). The low apparent temperature of the dome surface is related to data acquisition that performed at night and the insulating ash deposits that covered the dome during six distinct phreatic explosions that occurred between 2012 and 2014 (Darmawan et al., 2018). Highest apparent temperatures are found at the northern margins of the dome with a maximum temperature of 201.7°C. The high resolution of 1 px= 0.05 m in the 2014 thermal data allowed further investigation of the horseshoe-shaped structure in more detail. We show the thermal fingerprint of the fractures in



three area, c, d, and e, with a maximum apparent temperature of 161°C, 150°C, and 31°C, respectively (Fig. 5b). Cross section temperature profile of the horseshoe-shaped structure (Fig. 5b) shows a strong thermal signal in the horseshoe-shaped structure, which indicates a prominent pathway for hydrothermal fluids.

We repeated the thermal mapping campaign of the lava dome three years later (September 2017). The apparent temperature of fracture number 5 (area e) increased from 31 up to ~70°C, which may indicate increase of hydrothermal fluid activity in fracture number 5 (area e). The increasing of thermal activity in fracture 5 (area e) is highly correlated with the increase of hydrothermal alteration activity as observed by drone images in 2017 (Fig. 4 e-e'). However, as the thermal cameras used in 2014 and 2017 are different, the results cannot be directly compared. More details on this repeat thermal mapping can be found in the supplementary material.

### 3.3. Factor of Safety results

Assessment of factor of safety during intense rainfall first requires to quantify the effect of rain water. Based on a typical rainfall event (intensity of 10 – 35 mm/h) and assuming a rain duration of ~3 hours, we calculate the rain water percolation between 10 and 60 meters by using eq. 1 (Fig. 6). The 10 and 60 meters depth water percolation are then used to calculate the factor of safety as a function of failure plane inclination (Fig. 7). Results show that failure ( $FS \leq 1$ ) may occur when the failure plane is 25° and 45° ( $\alpha \geq \theta$ ) during shallow water percolation (10 m) scenario (black lines in Fig. 7b). It indicates that friction ( $\theta$ ) remains control the stability during shallow water percolation. For the deep water percolation scenario (60 m), the plane inclinations at failure mode ( $FS \leq 1$ ) are 15° and 39° for friction angles ( $\theta$ ) of 25° and 45°, respectively (red lines in Fig. 7b). This indicates that friction cannot resist the total driving forces when rain water percolates deeply and therefore destabilizes the lava dome sector. Calculation of the factor of safety reveals that the delineated dome sector is particularly unstable during deep water percolation ( $d \sim 60$  m). Using a basal inclination of 15°, the estimated unstable rock volume during intense rainfall events is  $0.3 \times 10^6 \text{ m}^3$ .

### 3.4. Scenario numerical model of block-and-ash flow

The estimated volume is now used as an input for the Titan2D simulation. Titan2D simulation results show that the debris materials mobilize down into the south eastern valley and reach 1.9 km from the summit at the first minute (Fig. 8). After 10 minutes, the debris materials are deflected by the Kendil hills (yellow triangle) and the main flow travels further to Gendol river valley with distance of 2.6 km from the summit. Within 30 minutes, the main flow reaches a distance of 3.1 km and it continues to travel along Gendol river valley. The flow finally stops with a maximum run out distance of 3.6 km from the summit. Most of the material is deposited at the upstream of Gendol river with a maximum thickness of ~10 meters. The potential hazard area (red polygon) due to the small volume single dome collapse is 1.5 km<sup>2</sup>.

## 4. Discussion



#### 4.1. Limitations

We find some limitations during drone, TLS, and thermal data acquisition due to complexity and hazardous access at Merapi summit after the 2010 explosive eruption. The drone was caught by turbulences due to fumarole activity and strong winds and the TLS data could only be obtained from the eastern crater wall since different scan position was too hazardous at Merapi summit. Therefore, the TLS data have significant shadowing effects. However, the advantageous of the TLS data is highly accurate and the drone is able to cover the shadow area. Combination of TLS and drone photogrammetry is therefore able to generate a Digital Elevation Model with resolution of 0.5 m and a photomosaic with resolution up to 0.03 m. We find that the combination of TLS and drone photogrammetry is robust and can be applied for geomorphology and structural mapping at steep sided dome building volcanoes.

The thermal variation was investigated by using a FLIR camera. Parameters such as emissivity, surface roughness, viewing angle, atmospheric effects, volcanic gas, instrumental errors, solar radiation, and solar heating may affect the pixel value of the FLIR thermal images (Spampinato et al., 2011). The effect of solar radiation and solar heating was largely reduced by acquiring the FLIR data before sunrise. However, parameters of emissivity, transmissivity, relative humidity, distance, and temperature background may influence during data processing. We tested the sensitivity of these parameters and we found that emissivity is the most sensitive parameter. Increasing emissivity by 0.01 will reduce the apparent temperature by  $\sim 1^\circ\text{C}$ . By assuming a range of emissivity between 0.95 and 0.98, which is common at dome building volcanoes (Merapi and Colima, Mexico) (Walter et al., 2013a; Carr et al., 2016), we infer that our apparent temperature has an uncertainty of  $\sim 3^\circ\text{C}$ . For the structural analysis performed, this is an acceptable effect.

The degree of dome instability is estimated by using the factor of safety calculation, assuming an intense rainfall event similar to the study of Simmons et al. (2004), where the parameters of dome sector geometry (thickness and fracture spacing), temperature, the friction angle, the rock strength, the intensity and duration of the rainfall may influence the result. Our factor of safety analysis is constrained for the southern Merapi dome sector. For this we hypothesize a fracture spacing (s) of 100 m, thickness (h) of 40 m, cohesive strength of 10 MPa following the studies of rock strength of altered rock from Mayer et al. (2016) and Pola et al. (2014), dome temperature of 200 – 800  $^\circ\text{C}$  during typical rainfall at Merapi (intensity of 10-35 mm/h and duration of  $\sim 3$  hours), and friction angles of  $25^\circ$  and  $45^\circ$  (Simmons et al., 2004; Husein et al., 2014).

Our morphological analysis, thermal images, and rainfall gauges provide realistic information of the fracture spacing (s), temperature to cool the dome ( $\Delta T_k$ ), and rainfall intensity, however, the parameters of dome thickness, rainfall duration, and temperature to vaporize rainwater ( $\Delta T_w$ ) have some uncertainty. Here, we tested those parameters and found that the rainfall duration is the most sensitive parameter as it influences the depth of water percolation. Doubling the rainfall duration from 3 to 6 hours with intensity of 35 mm/h will increase the water percolation by up to 10 meters, which will decrease the factor of safety by  $\sim 0.09$  and reduce the failure plane inclination ( $\alpha$ ) by  $1^\circ$ , while the dome thickness and temperature to vaporize the rain water ( $\Delta T_w$ ) are not significantly affected. Doubling the block thickness reduces the factor of safety by  $\sim 0.01$  and reducing the temperature to vaporize the water ( $\Delta T_w$ ) from 100 to  $90^\circ\text{C}$  only increases the factor of safety by  $\sim 0.006$ . By assuming rainfall duration of 12 hours during the rainy season, we estimate the failure plane inclination have an uncertainty

of  $\pm 3^\circ$  that may affect the uncertainty of the volume of the collapsing block by  $\pm 65.000 \text{ m}^3$ . We also assume that the rock cohesion is homogenous, while our drone photomosaic data shows that the degree of alteration that may influence the rock cohesion is vary spatially. We then further analyzed the factor of safety with heterogeneous rock cohesion in section 4.3.

In a case of dome sector failure, the potential hazard zone is estimated by using Titan2D software. Our Titan2D model represents an approximation of run out distance, deposit and potential hazard area due to single small volume dome collapses. However, Titan2D is not able to model pyroclastic surges. The pyroclastic surges that occurred and jumped over Kendil hills during the 2010 eruption could not be modelled by Titan2D as surges are dilute, mixed with gas and the propagation is not controlled by topography (Charbonnier et al., 2013). In order to solve the propagation of pyroclastic surge, a two-layer model has been proposed by assuming that pyroclastic density current (PDCs) consist of two distinct layers, a concentrated layer (block-and-ash flow) and a dilute layer (ash-cloud surge) (Kelfoun et al., 2017). The mobility of each layer is solved by using a depth-averaged algorithm. Results of this model were successfully to simulate the mobility of pyroclastic density currents of Merapi eruption in 26 October and 5 November 2010.

Other limitation of Titan2D is the grains interaction which is controlled and simply solved by coulomb frictions (bed and internal). While in reality, the interaction of grains in pyroclastic density currents is complex as the grains size is vary and the momentum produced by this interaction is able to transport large lithics in a great distance ( $\geq 10 \text{ km}$ ) (Dufek et al., 2009). A study of grains size of pyroclastic flow also suggests that finer grain size may produce a higher mobility of the center of the mass flow (Cagnoli and Piersanti, 2017; Cagnoli and Piersanti, 2015). As the grains interaction is only controlled by coulomb frictions in Titan2D, adjustment of coulomb frictions should be taken carefully and we used validated coulomb frictions from Charbonnier et al. (2012) in this study to obtain a good result.

#### 4.2. Geomorphology and structural instability at Merapi summit

The current morphology and structure at Merapi dome show progressive hydrothermal alteration that may cause structural weakening. Previous studies show that hydrothermal alteration is able to weaken the dome rock up to 0.2 - 10 MPa (Pola et al., 2014; Wyering et al., 2014) and promotes a failure even during quiescence periods (Lopez and Williams, 1993; Reid et al., 2001). Our alteration, thermal, and structural mapping datasets show that the southern Merapi dome and south west Merapi flank area are subjected to structural mechanical weakening. The southern dome sector is delineated by a curved horseshoe-shaped structure which was already identified even before the 2012-14 explosions (Darmawan et al., 2018). The structure then became strongly expressed and gradually deepened during the 2012-14 explosions. The horseshoe-shaped structure now has 8 m deep, highly fractured and provides pathways for fumaroles as identified by thermal camera. The presence of progressive hydrothermal alteration in fracture 5 (area E) probably points to a mechanical weakening and future structural instability due to hydrothermal alteration.

Whether the altered fractures are deep reaching or not, however, is difficult to quantify. Our data only identify alteration at the surface and our model assumes that the horseshoe shaped fracture is deeply altered and may transform to a translational fault due to lateral progressive hydrothermal alteration processes. Imaging the failure plane is challenging and hazardous at

Merapi summit. Resistivity tomography could only be realized at elevation of 2400 m (400 below the Merapi summit) at the south flank of Merapi and found a hydrothermal system at depth of 200 m (Byrdina et al., 2017). If alteration progressively occurs at depth of 200 m and gradually forms a failure plane, the southern dome and south west flank may pose to serious structural weakening and instability due to hydrothermal alteration.

- 5 Progressive hydrothermal alteration also intensively occurs at the open fissure area. The open fissure is highly fractured, actively degassing, and intensively altered as shown from our drone photomosaic image (Fig. 2b). The latest eruption in May 2018 occurred at the fissure area. Although no seismic or deformation precursors were observed, the thermal signal dramatically increased 15 minutes before the eruption along the fissure area (BPPTKG, 2018b). Further analysis of eruption material suggests that the May 2018 eruption contained an abundance of altered materials, which indicates that the open
- 10 fissure area is structurally already weakened due to hydrothermal alteration. The weakened structure thus provides a pathway to release gas overpressure and controls the location of steam explosion in May 2018.

#### 4.3. Implications for future dome failure

- Geomorphology and structural mapping imply a structural weakening at the southern dome, at the open fissure, and at
- 15 western crater wall. Results of factor of safety calculations show that deep water percolation may reduce the friction and may increase hydrothermal alteration that further weakens the dome structure. However, our results of factor of safety assume that the rock cohesion strength is homogenous, while in fact, the rock cohesion strength is probably heterogeneous as the magnitude of alteration and associated cohesion strength is vary spatially. Therefore, we further analyzed the dome stability by using Fellenius (ordinary slice) factor of safety and varying the rock cohesion strength. Fellenius factor of safety is widely
- 20 used to analyze slope stability and the method assumes that the mass above failure plane is divided into  $n$  slices and the external forces (vertical shear and horizontal forces,  $X_n$  and  $E_n$ , respectively) are zero (Fig. 9). The acting forces on each slice are the weight, pore pressure ( $u$ ), and rock cohesion ( $c$ ), respectively (Fig. 9a inset). We assume that the rock cohesion on basal failure is heterogeneous. The rock cohesion which located close to the altered fracture is 10 MPa, while the rock cohesion of fresh rock is 100 MPa (Pola et al., 2014; Wyering et al., 2014) and the water deeply percolates ( $\geq 60$  m) (Fig. 9).
- 25 We find that the factor of safety at the southern dome is 2.6 (Fig. 9a) which indicates a stable condition. We infer that fresh rock is strong enough to resist and to stabilize the dome. The factor of safety of the south western flank is 1.3 which may indicate a critical condition. We therefore also recommend to monitor the stability of the south western flank as historically the south western flank has frequently collapses over the past decades.

- Other factors such as a new magma extrusion and gas overpressure may also destabilize and trigger a dome failure. Gas
- 30 pressurization may promote deep-seated failure and explosive eruption, while slow rate magma extrusion can gradually oversteepen the dome and trigger gravitational collapses (Voight and Elsworth, 2000).

Currently, a new dome is growing at the middle of open fissure with volume of 135.000 m<sup>3</sup> and extrusion rate of 0.01 m<sup>3</sup>/s (BPPTKG, 2018a). The extrusion of new dome involves degassing activity that increases hydrothermal alteration at the

southern dome sector. Further investigation of the interaction of the new dome extrusion and structural weakening is now required.

#### 4.4. Block-and-ash flow hazard along the Gendol valley

5 Our simulation of block failure and mobility along the Gendol valley shows the potential hazard due to structural weakening at the southern dome. The southern dome with volume of  $\sim 0.3 \times 10^6 \text{ m}^3$  may fail and produce block-and-ash flow with a maximum run out distance of 3.6 km and an affected hazard area of 1.5 km<sup>2</sup>. This run out distance is typical for single dome collapse with volume of  $\leq 10^6 \text{ m}^3$  (VEI = 1). The single dome collapse in 2006 with a volume of  $1 \times 10^6 \text{ m}^3$  travelled along Gendol valley and destroyed the village of Kaliadem which was located 4.5 km from the summit (Charbonnier and Gertisser, 10 2009; Ratdomopurbo et al., 2013). Therefore, we infer that our potential hazard model is relevant and realistic for single dome collapse with VEI 1. However, we do not consider the potential collapses of the new lava dome that currently formed at the open fissure and grows above the frozen lava dome. As the current morphology of the Merapi summit that opened to the Gendol valley (south – southeast), we infer that the new lava dome potentially collapse to the Gendol valley due to magma intrusion, gravitational instability, gas overpressure, structural weakening, intense rainfall, and earthquake. We 15 recommend to further monitor and investigate the potential hazard of the new lava dome in the near future.

#### 5. Conclusion

Detail morphological and structural studies of the active Merapi volcano reveal a structural weakening due to hydrothermal alteration at the southern dome. We identify a 165 m long horseshoe-shaped structure with depth of 6 m that encircling the 20 southern dome sector which has volume of  $\sim 0.3 \times 10^6 \text{ m}^3$ . The structure is highly fractured and provides pathways for hydrothermal fluids which can lead to structural instability.

Our results of factor of safety calculations indicate that intense rainfall events at Merapi summit are able to reduce the failure plane inclination. The southern dome may fail with minor changes in failure plane. By using Titan2D flow simulation we estimate that the collapse of the unstable dome sector may produce block-and-ash flow that travel southward with maximum 25 run out distance of  $\sim 4 \text{ km}$  from the summit.

#### Acknowledgments

This is a contribution to VOLCAPSE, which is a research project funded by the European Research Council under the European Union's H2020 Programme/ERC consolidator grant ERC-CoG 646858 and the authors also acknowledge a 30 scholarship grant from Deutscher Akademischer Austauschdienst (DAAD), Germany reference number 91525854, a research grant by the Swedish Research Council (VOLTAGE project), and financial support by the Swedish Center for Natural Hazard and Disaster Sciences (CNDS). We would like to also thank Mehdi Nikkhoo and Nicole Richter for the data acquisition of Terrestrial Laser Scan, Michele Pantaleo for the data acquisition of thermal images in 2014, and Edgar Zon for supporting the 2014 fieldwork. We also thank François Beauducel for supporting FLUKE camera recording during field

work in 2017 and special thank goes to BPPTKG (Merapi Volcano Observatory) for all supports during field works in 2014, 2015, and 2017. We also thank Dr. S. Charbonnier and anonymous referee for their constructive review. The high resolution DEM and photomosaic in this study are available in GFZ data publication with DoI: <http://doi.org/10.5880/GFZ.2.1.2017.003>

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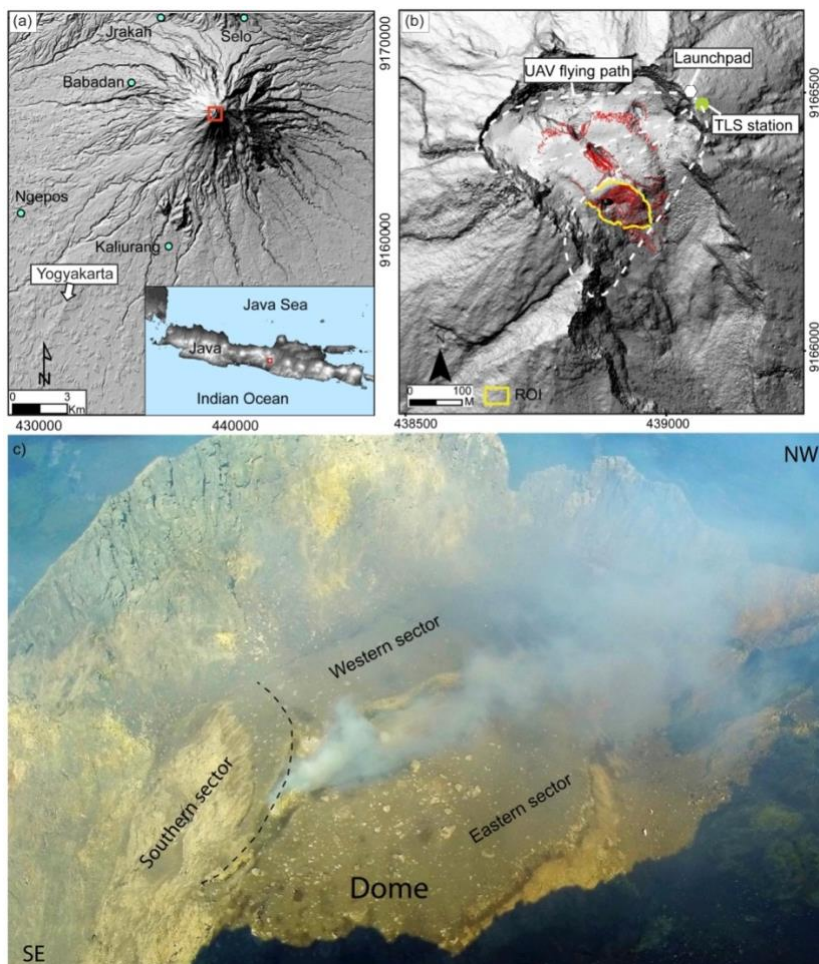


Figure 1. (a) Shaded relief of DEM from Gerstenecker et al. (2005) shows the morphology of Merapi volcano, the most active volcano in Indonesia. Merapi is located ~30 km from the densely populated city of Yogyakarta and therefore the activity of Merapi is intensively monitored by five observatories (blue dots). (b) TLS and drone photogrammetry field campaigns have been conducted in September 2014 and October 2015, respectively to investigate the detailed structure and morphology of the Merapi lava dome. Coordinates are in UTM meters. (c) The aerial image of Merapi dome in 2014 shows the delineated unstable dome sector at the southern flank that is the focus of the present investigation.

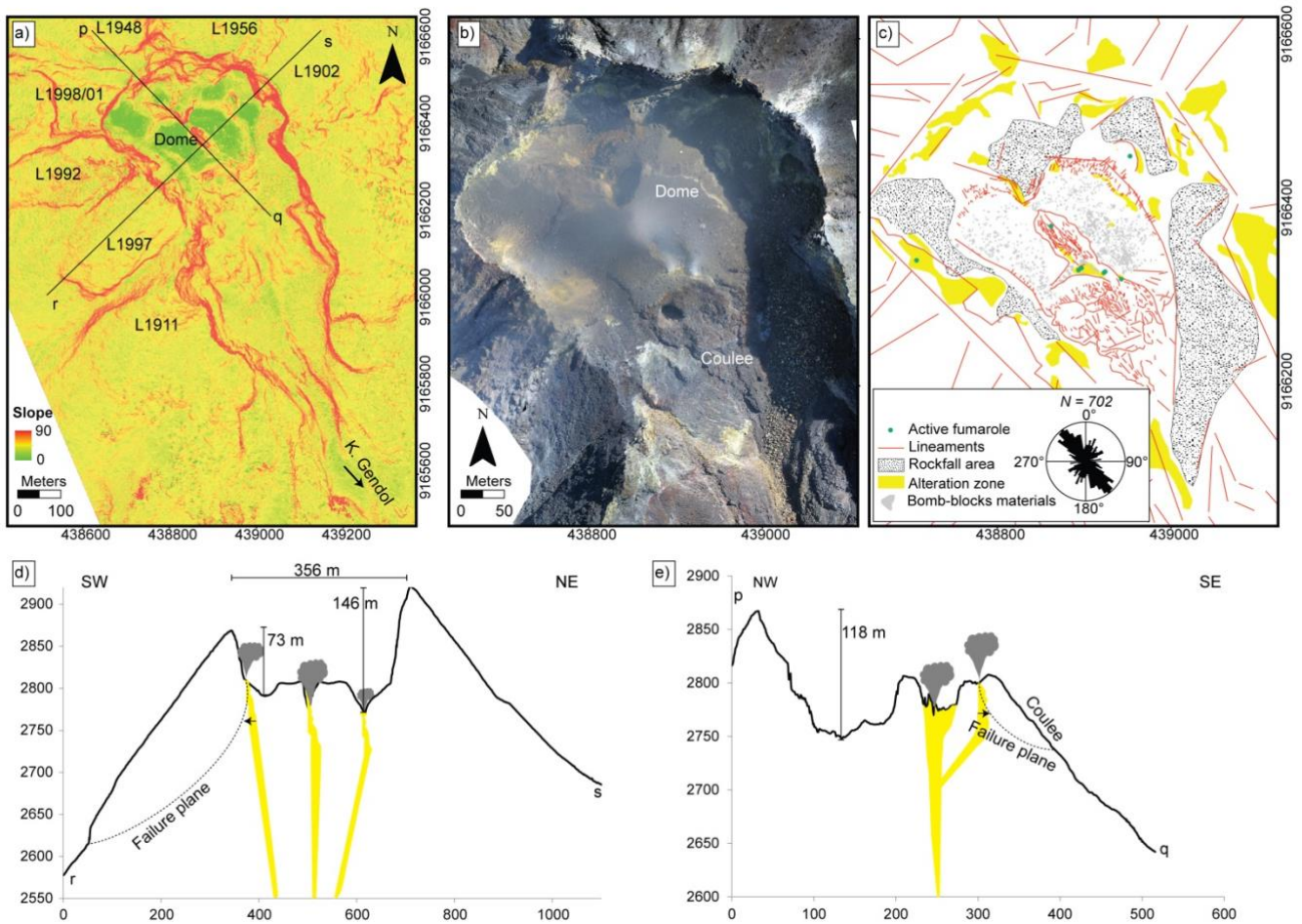


Figure 2. (a) Slope map of the Merapi summit shows that the Merapi flanks is very steep, especially the crater which has a slope of  $\sim 80^\circ$ . (b) Photomosaic of drone aerial images shows that the summit is highly fractured with  $N150^\circ E$  and highly altered. (c) Rock alteration is seen inside the crater. Cross section of line d) r-s and e) p-q shows that the crater has a maximum depth of 146 at the north east area. The current Merapi dome is located in the middle of the deep crater and it is too hazardous for direct investigation.

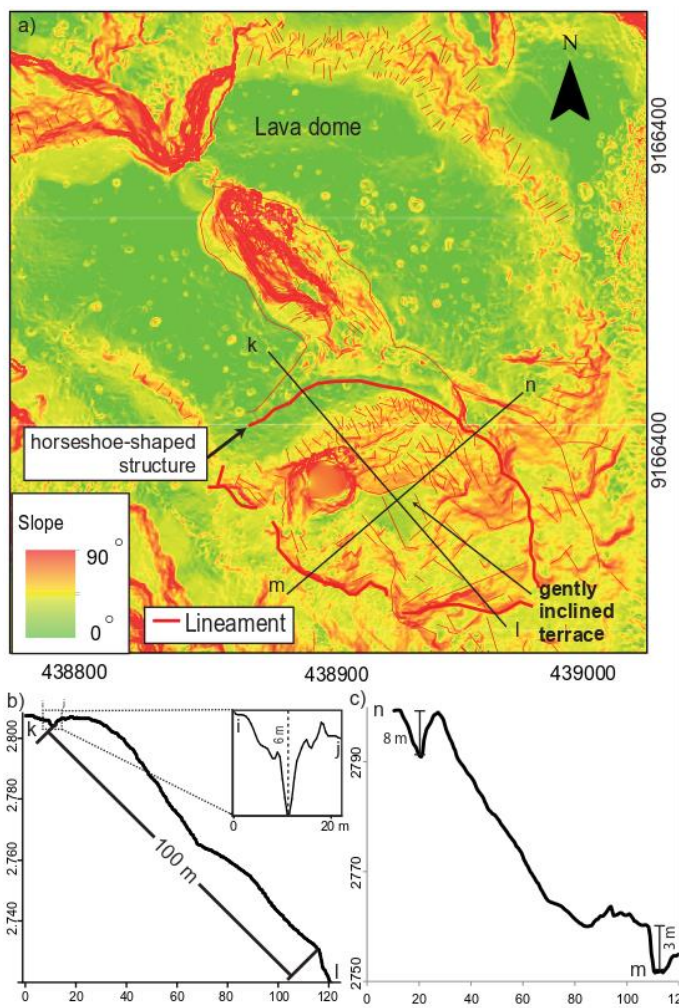


Figure 3.a) Detailed slope map of the Merapi lava dome shows that the top of the dome is relatively flat. The fissure and the north part of the dome are steeply inclined with slope of  $\geq 80^\circ$ . The southern block possibly consists of two different flow units which are separated by a gently inclined terrace. Cross section of line b) k-l and c) m-n shows that the horseshoe-shaped structure has maximum depth of 8 m.



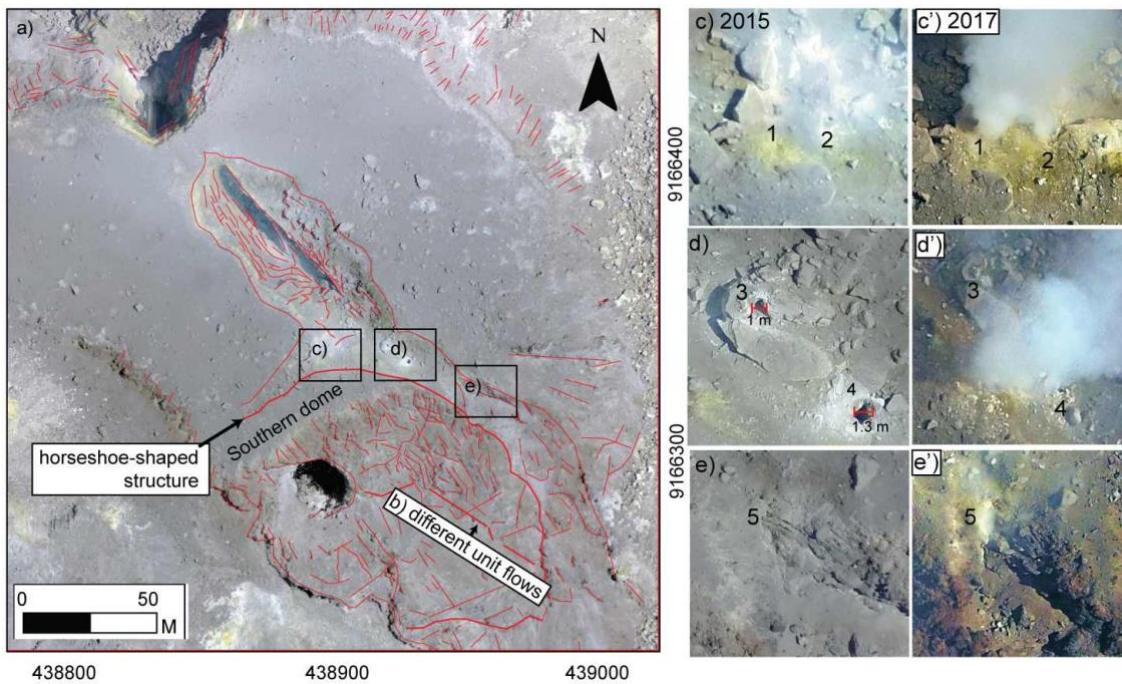


Figure 4. a) Photomosaic of UAV aerial images acquired in 2015 shows detailed structures at the Merapi lava dome. The dome is highly fractured at the fissure area, at the dome margin, and at the southern dome. b) A different unit flows and three fractures area (c, d, and e) are clearly identified by our photomosaic drone aerial images. Coordinates are in UTM. Zoomed images of drone images between 2015 and 2017 at those three fractures area show a mechanical weakening due to hydrothermal alteration, especially at the fracture number 5 (area e – e'). We estimate that the diameter of the 1<sup>st</sup>, 2<sup>nd</sup>, 3<sup>th</sup>, 4<sup>th</sup>, and 5<sup>th</sup> fractures is 0.7 m, 0.3 m, 1 m, 1.3 m, and 0.3 m, respectively.

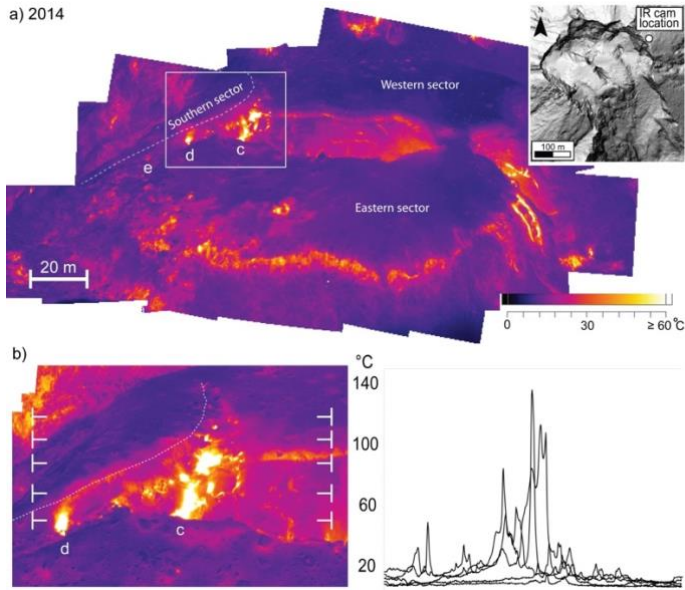


Figure 5. (a) Photomosaic of high resolution thermal image taken from the eastern flank (inset) shows the variation of apparent thermal variation of the Merapi dome in 2014. (b) We find high temperatures around  $\geq 140$  °C at the horseshoe-shaped structure and along the fracture area of c, d, and e as identified by our drone camera. High thermal pixel value may indicate a hydrothermal fluid activity that can progressively alter and ultimately weaken the dome.

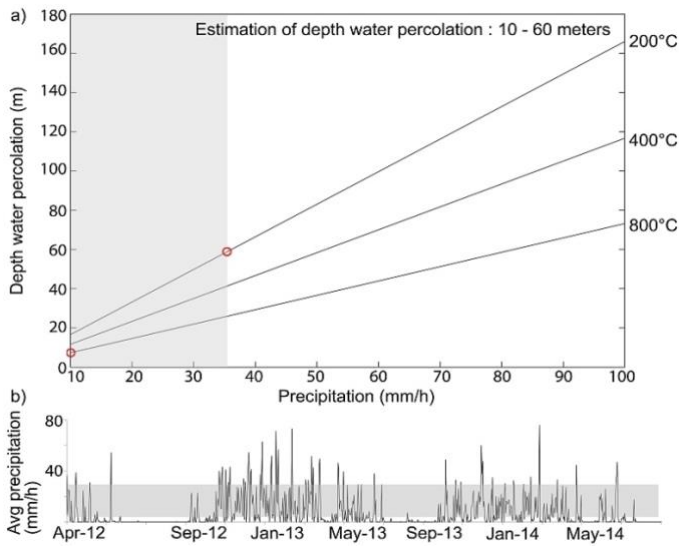


Figure 6 (a). Depth of rain water percolation as a function of rainfall intensity **at** Merapi is controlled by the temperature of the dome. By assuming the minimum and maximum temperature of the dome of 200 and 800 °C, respectively, the estimation of depth water percolation is 10 **to** 60 meters (red circles) during typical rainfall (grey area). (b) The typical intensity of rainfall from April 2012 to July 2014 was 10 – 35 mm/h (grey area) and was calculated based on average rain intensity from five observatories **near** Merapi (see Fig. 1a).



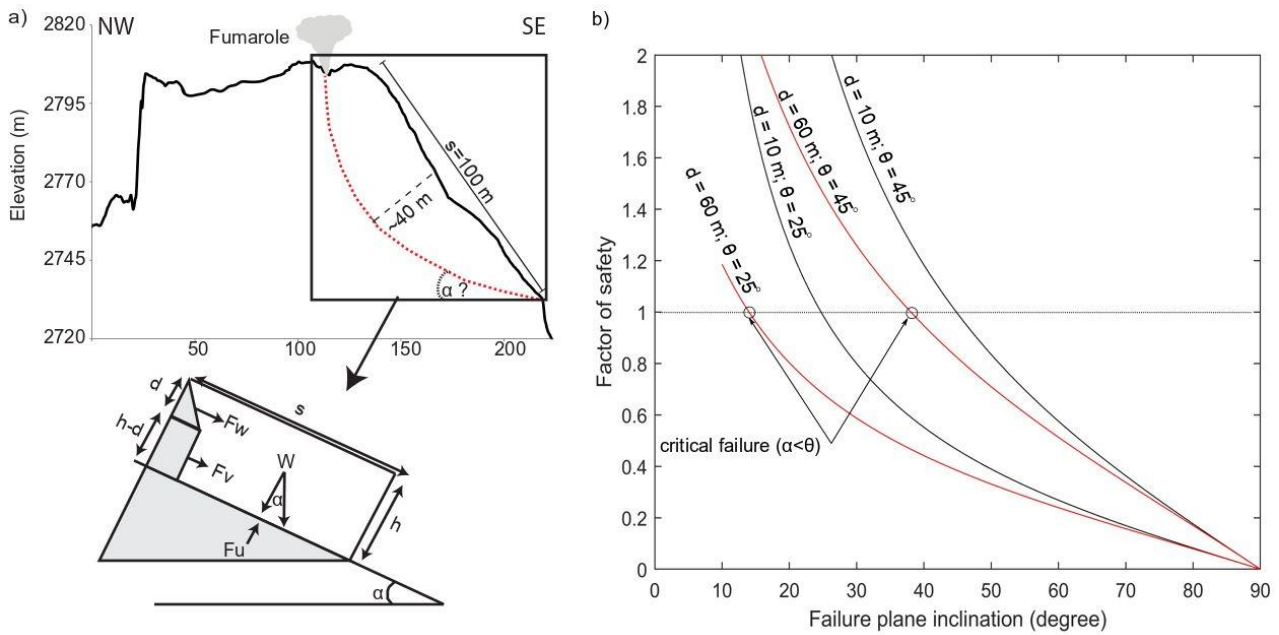


Figure 7. (a) Cross section of the Merapi lava dome shows that the horseshoe-shaped structure may develop into a translational fault with fracture spacing (s) up to 100 m and a hanging wall thickness of ~40 m. The stability of the unstable dome sector is influenced by the weight (W), water force ( $F_w$ ), vaporized water force ( $F_v$ ), and gas uplift force ( $F_u$ ) along the fault boundary during intense rainfall (Modified after Simmons et al. (2004)). (b) Analysis of factor of safety for the southern dome sector shows that deep water percolation (red lines) may reduce failure plane inclination and failure may occur even if the failure plane is inclined gently below the friction angle ( $\alpha < \theta$ ).

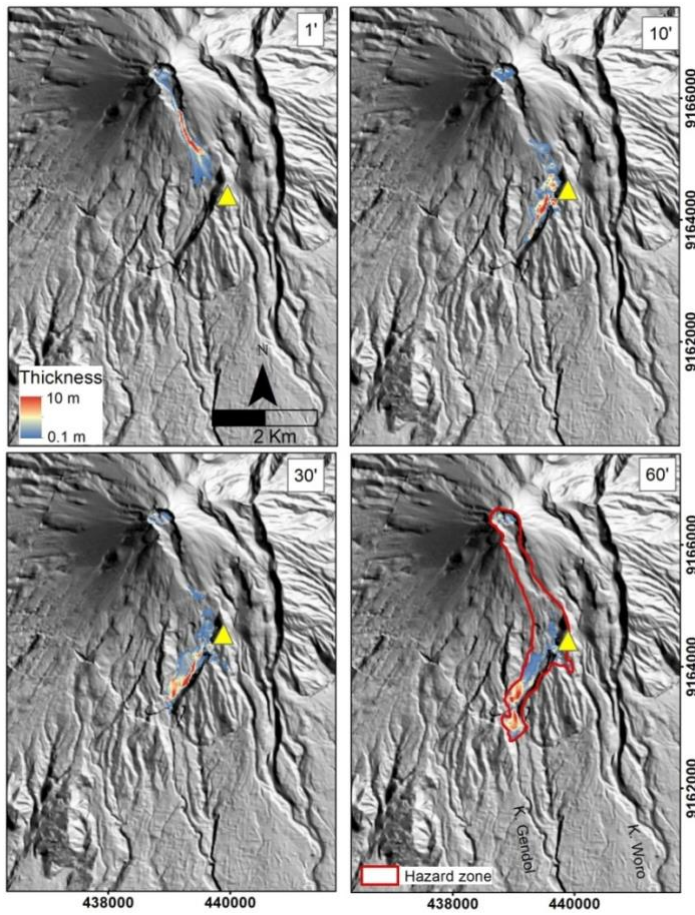


Figure 8. Result of numerical simulation of the pyroclastic block-and-ash flow that may form due to collapse of the delineated southern dome sector after 1, 10, 30, and 60 minutes. The block-and-ash flow are deflected by the Kendil hills (yellow triangle) within ~1 minute after the collapse. The red outline indicates the total inundation zone as a result of the deposition of the block-and-ash flow material. Coordinates are in UTM meters.



Table 1. Detail parameters to calculate depth water percolation (eq.1) and factor of safety (eq. 2)

| Parameters                                  | Value                                     | Source   |
|---|---|--|
| Thermal diffusivity ( $K_R$ )               | $1.4 \times 10^{-6} \text{ m}^2/\text{s}$ | Taron et al. (2007)                                    |
| Heat capacity of rock ( $c_R$ )             | 918 J/kg K                                | Taron et al. (2007); Simmons et al. (2004)             |
| Heat capacity of water ( $c_W$ )            | 4187 J/kg K                               | Taron et al. (2007); Simmons et al. (2004)             |
| Rain duration ( $t$ )                       | 3 hours                                   | Assumption   |
| Rain intensity ( $i$ )                      | 10 – 100 mm/h                             | Data observation                                       |
| Thermal to cool fracture ( $\Delta T_R$ )   | 200 – 800° C                              | Thermal datasets and from Matthews and Barclay (2004). |
| Thermal to vaporized water ( $\Delta T_W$ ) | 100°C                                     | Assumption   |
| Fracture spacing (s)                        | 100 m                                     | Digital Elevation Model                                |
| Dome sector thickness (h)                   | ~40 m                                     | Estimation   |
| Density of rock ( $\rho_r$ )                | 2242 kg/m <sup>3</sup>                    | Tiede et al. [2005]                                    |
| Density of water ( $\rho_w$ )               | 1000 kg/m <sup>3</sup>                    | Taron et al. (2007); Simmons et al. (2004)             |
| Density of gas ( $\rho_g$ )                 | 0.75 kg/m <sup>3</sup>                    | Girona et al. (2015)                                   |
| Cohesive strength (Cs)                      | 10000 kN/m <sup>2</sup>                   | Mayer et al. (2014); Pola et al. (2014)                |
| Friction angle ( $\theta$ )                 | 25° - 45°                                 | Husein et al. (2015); Simmons et al. (2004)            |
| Gravitational acceleration (g)              | 9.8 m/s <sup>2</sup>                      |  |

Table 2. Detail input parameters in Titan2D simulation

| Parameters                      | Input data             | Source  |
|---------------------------------|------------------------|---|
| Topography model                | Updated DEM            | Drone photogrammetry + TLS + DEM from Gerstnecker et al. (2005) |
| Number of flux source           | 1                      | Assumption  |
| Duration (s)                    | 3600 s                 | Maximum time computation  |
| Volume                          | 300.000 m <sup>3</sup> | DEM + failure plane inclination from FS analysis                |
| Initial velocity                | 0                      | Assumption  |
| Internal coulomb friction angle | 30°                    | Charbonnier et al. (2012)                                       |
| Bed coulomb friction angle      |                        | Charbonnier et al. (2012)                                       |
| Zone 1 :> 2426                  | 28°                    | Zone 1 :> 2426  |
| Zone 2 : 2053 – 2425            | 27°                    | Zone 2 : 2053 – 2425  |
| Zone 3 : 1680 – 2052            | 26°                    | Zone 3 : 1680 – 2052  |
| Zone 4 : 1555 – 1679            | 24°                    | Zone 4 : 1555 – 1679  |
| Zone 5 : 1431 – 1554            | 22°                    | Zone 5 : 1431 – 1554  |
| Zone 6 : 1306 – 1430            | 20°                    | Zone 6 : 1306 – 1430  |
| Zone 7 : 1182 – 1305            | 18°                    | Zone 7 : 1182 – 1305  |
| Zone 8 : 0 - 1181               | 16°                    | Zone 8 : 0 - 1181   |