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# How severe Space Weather can disrupt global supply chains

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## Abstract

Coronal mass ejections (CMEs) strong enough to create electromagnetic effects at latitudes below the auroral oval are frequent events that could soon have substantial impacts on electrical grids. Modern society's heavy reliance on these domestic and international networks increases our susceptibility to such a severe space weather event. Using a new high-resolution model of the global economy we simulate the economic impact of strong CMEs for 3 different planetary orientations. We account for the economic impacts within the countries directly affected as well as the post-disaster economic shock in partner economies linked by international trade. For a 1989 Quebec-like event the global economic impacts would range from USD 2.4 to 3.4 trillion over a year. Of this total economic shock about 50 % would be felt in countries outside the zone of direct impact, leading to a loss in global GDP of 3.9 to 5.6 %. The global economic damages are of the same order as wars, extreme financial crisis and estimated for future climate change.

## 1 Introduction

We are now midway in the current solar cycle and so far we have not experienced a single extreme geomagnetic storm. The probability of such an event is highly increased due to a peak in solar storms when sunspot numbers are maximal (Ramesh, 2010). Solar storms consist of three major components: solar flares, solar proton events and coronal mass ejections (CMEs). All of these cause “space weather” that affect humanity's technological systems and society, as well as Earth's atmosphere, climate, and potentially the biosphere. Fast CMEs ( $\approx 1000\text{--}2000\text{ km s}^{-1}$ ) are clouds of ejected plasma with embedded magnetic fields that can interact with Earth's magnetic field after an observed travel time as short as 15 h to create a geomagnetic storm (Cliver et al., 2004). Following this impact Earth's magnetic field can be disturbed worldwide for days (Bolduc, 2002), allowing more energetic solar and magnetospheric charged particles

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to find their way along the open magnetic field lines near the Earth's poles through the ionosphere and atmosphere to the surface. Many details of the associated physics are still unclear. However, currents and electric fields associated with enhanced particle precipitation can induce massive ground currents in electrical distribution networks which could result in large-scale power blackouts and permanent damage to electric transformers (Pirjola et al., 2000).

During a geomagnetic storm large time-varying currents are introduced in the ionosphere at auroral-like latitudes. These occur primarily on the night side where magnetic reconnection in the magnetotail accelerates particles along polar magnetic field lines. These time-varying auroral magnetic fields induce large sudden electric fields, voltage drops, and currents in power-lines and transformers, causing failures and possible reduction in electric power supply. Similar effects of geomagnetic storms due to sudden increased ionization on the dayside near the subpolar point due to a large X-ray flare from the Sun, or increased entry of energetic particles and interplanetary plasma due to magnetic reconnection on the dayside, are likely smaller and are ignored here. The strength of the AL and  $D_{st}$  indices, both due to variations in the magnetic field measurable on the ground due to space weather, depend strongly on the interplanetary event: when the interplanetary magnetic field  $B_z$  is directed southwards  $AL = -v_{sw}^2 B_s$ , where  $v_{sw}$  is the solar wind speed and  $B_s$  is a duration-weighted estimate of  $B_z$  (Muruyama et al., 1980).

The strength of the induced currents depends on a number of factors. They usually increase with geomagnetic latitude, transmission line length and voltage, but decrease with distance to the ocean and increased ground resistivity (Wei et al., 2013). Space weather events also cause auroras, usually in two small ovals around  $65 (\pm 5)$  degrees northern and southern latitude that vary in size, location and intensity during geomagnetic storms. The geographical distribution of the damage caused by a geomagnetic storm is very complex. Other observed consequences of geomagnetically induced currents (GICs) include damage to pipelines and telecommunication cables, accelerated corrosion, physical and electrical damage to satellites, and disruptions

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to radio navigation, which can particularly affect the transport and aviation sectors (Boteler et al., 1998).

In 1989 Earth experienced its largest space weather event in several decades: a geomagnetic storm that caused a power blackout in Quebec that left millions of people without electricity for hours. It permanently damaged transformers in Canada, the USA and the UK, and disconnected other power transmission devices from California to Sweden (Erinmez, 2002; Lakhina et al., 2005). This storm caused damage across about  $120^\circ$  of longitude and  $5-10^\circ$  latitude and lasted for more than 12 h. The Quebec power grid went from normal operations to complete shutdown in 90 s. Temporal changes in the geomagnetic field of  $dB/dt = 1100 \text{ nT min}^{-1}$  were experienced and the strength of the storm, in terms of the Disturbance storm index which measures how much Earth's magnetic field is weakened, was estimated to be  $D_{st} = -640 \text{ nT}$ . Two other strong storms in the 20th century include a  $dB/dt = 5000 \text{ nT min}^{-1}$  storm in May 1921, the biggest geomagnetic event in the last century which led to aurora borealis over Samoa, and a fast CME in October 2003 which despite its low strength of  $D_{st} = -472 \text{ nT}$  caused effects at latitudes as low as South Africa where it incapacitated several large electrical transformers (Lakhina et al., 2005). Regions with latitudes below  $30^\circ \text{ S}$  were previously thought to stay free of damage.

Polar ice studies and anecdotal evidence suggest the most severe space weather event in the last 450 years was the Carrington event of September 1859 (Shea et al., 2006). That storm caused auroras visible within  $23^\circ$  of the equator in both hemispheres, e.g. in Honolulu, Havana, and Rome (Tsurutani et al., 2003). In the United States and Europe fires were started by arcing from currents induced in telegraph wires (Green et al., 2006). The strength of this storm has been estimated to be  $D_{st} = -1760 \text{ nT}$  or three times stronger than the 1989 event and four times stronger than the October 2003 storm (Lakhina et al., 2005; Tsurutani et al., 2012). In August 2013 a CME of Carrington size missed the Earth by a week, or  $90^\circ$  in heliographic longitude.

Although a solar maximum period might have a higher frequency of intense solar storms, there is no evidence that this will affect the intensity of any single event

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(Hapgood, 2012). Indeed, the 1859 event occurred outside solar maximum. The probability of a Carrington event (based on  $D_{st} < 850$  nT) per decade is estimated to be 12%, or a once in a century event like a 9.0 earthquake (Riley, 2012; Love, 2012). It has been estimated (Thompson et al., 2011) that  $dB/dt$  changes of 1000–4000 nT min<sup>-1</sup> ( $D_{st} = 2000–5000$  nT) for a storm occur every 100 years and  $dB/dt$  of 1000–6000 nT min<sup>-1</sup> ( $D_{st} = 3000–6500$  nT) every 200 years. These frequencies are comparable to other severe natural disasters such as large earthquakes and volcanic eruptions. Power grids typically experience problems when the rate of change of the magnetic field exceeds a 100–200 nT min<sup>-1</sup> (Wei et al., 2013). Occurring today, the Quebec 1989 event or the 1859 Carrington event would have profound impact on the daily lives of millions of people, both through direct effects and via the impacts to the globalized economic production system.

Little has been done on economic modeling of severe space weather so far and previous studies have mostly focused on the USA. It has been estimated in NAOS (2008); Showstack (2011) that a storm similar to that of 1859 or 1921 could cause damage of several trillion US dollars in the USA in the first year alone and that recovery could take years. The large transformers that could be vulnerable to a severe storm are produced infrequently, of order just 1–5 globally per year. The estimated damage to the power system in Quebec in 1989 is in the range of USD 2 billion, whilst the total damage is estimated to be around USD 13 billion (Kapfenmann, 2010; Boteler, 1998). Another study estimates that the economic losses in North America and Europe for a power blackout for 5 months caused by a Carrington-like event would be between USD 0.5 and 2.6 trillion (Lloyds, 2013; Wei et al., 2013). It has also been estimated that a North American power grid blackout would result in a GDP loss in the USA of about USD 30 billion per day, accumulating to over USD 10 trillion per year (Lloyds, 2013). None of the cost estimates consider indirect effects on supply chains, including those of global trade.

The complex and interconnected network of today’s globalized economy and infrastructure makes it difficult to predict the exact effects of a severe space weather event.

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Therefore we focus on the most economically important impact from such an event: the interruption of electrical distribution grids and failure of electric power transmission systems. We combine a simple physical model for disruption of power grids with the most comprehensive and most highly resolved economic input–output framework of the world economy to estimate the direct and indirect economic costs of severe space weather events with sizes between the Quebec 1989 and Carrington 1859 events.

The reduction of production capacities of the electricity sector for each country is

$$R(S, C, \phi_0) = F(S) A_C^{-1} \int dAG(S, C, \phi_0)$$

$$= F(S) A_C^{-1} \int dA e^{\frac{-(\phi - \phi_0)^2}{2\sigma_\phi(S)^2}} \times \left[ e^{\frac{-(\theta - \theta_0(S))^2}{2\sigma_\theta(S)^2}} + e^{\frac{-(\theta + \theta_0(S))^2}{2\sigma_\theta(S)^2}} \right]. \quad (1)$$

The quantity  $G(S, C, \phi_0)$  is the product of a Gaussian in longitude  $\phi$ , centered at longitude  $\phi_0$  (which corresponds to the time when the event occurred) with event-dependent standard deviation  $\sigma_\phi(S)$ , that depends on the event size  $S$ , times the sum of Gaussians in latitude that model the event-dependent auroral ovals centered at  $\pm\theta_0(S)$  with standard deviations  $\sigma_\theta(S)$ . Figure 1 illustrates the double-banded nature of the affected areas.

How should  $\theta_0(S)$ ,  $\sigma_\theta(S)$  and  $\sigma_\phi(S)$  vary with  $S$ ? Noting that the magnetic field  $B(r)$  at radial distance  $r$  from a long axial current  $I$  varies as  $B(r) = \mu_0 I / 2\pi r$ , where  $\mu_0$  is the permeability of free space, it is clear that the distance  $r$  at which the same value of  $B$  is observable increases linearly with  $I$ . Thus as a first approximation  $\sigma_\theta(S)$  and  $\sigma_\phi(S)$  should vary linearly with  $I$  and so with  $S$ . Observations show that  $\theta_0(S)$  decreases from non-zero values near 65° geographical latitude as  $S$  increases, since the aurora moves equatorward as geomagnetic storms and the associated currents intensify and then moves poleward as the driving currents decrease and the system recovers (Baumjohann et al., 1980). Thus the geographical area in which a certain level of damage occurs should vary as  $S^2$ , but should move equatorward as  $S$  increases.

We assume  $\sigma_\theta = 2^\circ \pm 1^\circ$  and  $\sigma_\phi = 20^\circ \pm 5^\circ$  for a storm similar to the Quebec 1989 Space Weather event. This storm has a footprint of about  $80^\circ$  of longitude and about  $8^\circ$  of latitude in both the northern and Southern Hemisphere. Since the  $D_{st}$  and AL values for the Quebec 1989 and Carrington 1859 events are believed to have differed by a factor 3, the geographical footprint of the Carrington event is expected to be a factor of 9 larger and the values of  $\sigma_\theta$  and  $\sigma_\phi$  each a factor of 3 larger. From the physical model  $\sigma_\theta = 6^\circ \pm 3^\circ$ , and  $\sigma_\phi = 60^\circ \pm 15^\circ$ . During the Carrington event auroras were observed as far south as  $20^\circ$  latitude with the latitudinal spread observed to be  $45^\circ$ , and the longitudinal domain was close to  $180^\circ$  (Green et al., 2006). Accordingly, the model is consistent with observations.

For each country the total impact of the storm is the quantity  $R$  that integrates the storm's effects as a function of geomagnetic latitude and longitude over the country's area. The storm is modeled like a flash-like impact. Outside the area of impact the damage in the electricity sector is zero. Storms weaker than the Carrington 1859 event but stronger than the Quebec 1989 event could result in around 10–20 damaged transformers in the US alone (OECD, 2011; MITRE, 2011; UK House of Commence Defence Committeee, 2012). Even the failure of a small number of transformers serving a highly populated area like the ones we choose in our scenario is enough to create prolonged power outage. We assume the storm causes damage which will last a year since the production and supply of a replacement transformer could take up to more than 12 months as could the restoration of a grid damaged over a huge area (Office of Energy Delivery & Electric Reliability, 2012).

In order to quantify the economic impacts of a severe Space Weather event we simulate the consequences of major disasters by utilizing Leontief's input–output (IO) theory (Steenge et al., 2007; Leontief, 1996). IO analysis has been used extensively for investigating the repercussions of changes in one part of an economy on other parts of the same economy (see the recent articles by Lenzen et al., 2011 and Wiedmann et al., 2013). Input–output databases are routinely published by more than 100 national statistical bureaus in the world. More recently, a number of teams have assembled large-

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scale, detailed global Multi-Regional Input–Output (MRIO) databases, which contain the same set of data but integrated for all world regions or countries (Tukker et al., 2013). MRIO tables can be used in the same analytical manner as national input–output tables, for investigating effects that ripple along global supply-chain networks (Leontief et al., 1963). In this study we utilize the most detailed of these global MRIO database, distinguishing 187 countries with 25–400 sectors per country. The economic model captures more than 99.99% of global trade.

## 2 The model

Most often, productive activity in modern economies is assumed to be demand-driven, and the so-called demand-pull model is evoked, where an initial change vector  $\Delta \mathbf{y}$  in final demand  $\mathbf{y}$  ( $N \times 1$ , for example decreased household consumption caused by reduced electricity supply) causes flow-on effects that ripple through a complex upstream supply-chain network, and ultimately leads to a change  $\Delta \mathbf{x}$  in total output  $\mathbf{x}$  ( $N \times 1$ ) of an economy. The scalar  $N$  holds the number of sectors (industries and/or products) that are distinguished in the IO matrices. We distinguish  $P = 15909$  country-sector pairs (Lenzen et al., 2013, 2014) using data from 2011. The flow-on effects can be enumerated using an  $N \times N$  input–output transactions matrix  $\mathbf{T}$ , according to  $\Delta \mathbf{x} = (\mathbf{I} - \hat{\mathbf{T}}\hat{\mathbf{x}}^{-1})^{-1} \Delta \mathbf{y} = (\mathbf{I} - \mathbf{A})^{-1} \Delta \mathbf{y}$ , where  $\mathbf{I}$  denotes an  $N \times N$  identity matrix, the hat symbol “ $\hat{\cdot}$ ” denotes matrix diagonalisation, and  $\mathbf{A} = \mathbf{T}\hat{\mathbf{x}}^{-1}$  is the matrix of input coefficients. This relationship follows from the National Accounting Identity, which states that  $\mathbf{x} = \mathbf{T}\mathbf{x} + \mathbf{y} = \mathbf{A}\mathbf{x} + \mathbf{y}$ , where  $\mathbf{1} = \{1, 1, \dots, 1\}'$  is an  $N \times 1$  summation operator. A transaction matrix  $\mathbf{T}$  is essentially a square matrix with elements  $T_{ij}$  that represent the supply of products  $i$  for use in industry  $j$ . Matrices  $\mathbf{T}$  and  $\mathbf{A}$  thus include information on industrial interdependence and production structures in an economy, which can ultimately be used to trace flow-on effects of initial changes along supply chains that link all sectors in an economy.

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the final consumer, and hence have no damage-multiplying effect. Brazil is an equally interesting case here, because whilst being a net importer from the USA and most of Europe, it is a net exporter to China, Japan, Korea and Australia, and hence registers surplus production after a storm.

5 The economic model considers both direct impacts in international trade, such as the shared international power grids in Europe, and the indirect effects due to interrupted supply chains. In scenario 1, an American storm, indirect plus direct effects are calculated to reduce global consumption possibilities by 3.9% or USD 2.4 trillion. In scenario 2 (the European storm) direct and indirect effects are calculated to reduce global consumption possibilities by 5.6% or USD 3.4 trillion. In the Asia-centered storm of scenario 3 the storm is estimated to reduce global consumption possibilities by 5.0%, or USD 3.1 trillion.

#### 4 Discussion

15 This paper we concentrated on the possible impact of severe Space Weather events on the electric distribution system. Damage in the telecommunication sector (not shown) can also result as a consequence of solar activity but is due to solar flares (accompanied by x-rays and density perturbations in the ionosphere) and are more distributed in time and locations than the damage caused by CMEs. Accordingly, a different physical model than the one we use here is needed to account for damages in the telecommunication sector. A space weather event is substantially different from other natural disasters on Earth. Whilst hurricanes, earthquakes and tsunamis could cause direct human losses, a solar storm is likely to cause material damage only. Although radiation risks for astronauts and airline passengers on polar routes are described in the literature, no human losses as a consequence of a solar storm have been recorded, and therefore are not considered in this work, for example as a loss of labor.

25 A severe Space Weather event could be the worst natural disaster in modern history with global costs estimated to be over 5% of world GDP and impacts reaching

across every industry and every segment of society. Extreme space weather will impact severely on society's infrastructure – networks of trade, transport and production would need to adapt globally. In our modern globalized economy shocks to the production system in one country can cause large ripple effects in partner economies. Reduced inventories, increased shipping, the rise of just-in-time production and the acceleration of specialization and trade mean that the global economic production system, while more productive in total, is increasingly vulnerable to shocks. We have considered the possible impact of a century-scale space weather event on the global economy. The results indicate that total losses could be up to USD 3.4 trillion and impacts would affect sectors and populations well outside the direct area of impact. Changes in the intensity and timing of space weather event result in different global economic damage.

10 Global financial crises episodes lead to losses estimated between 2.95 and 4.54% of world GDP (Kappy et al., 2012). Economic impacts from climate change have been estimated to cost USD 125 billion yr<sup>-1</sup> (GHF, 2009). Our scenario estimates global GDP damage in between climate change and global financial crisis.

#### 5 Conclusions

20 For the first time a physical and an economic model have been combined to analyze the global economic impacts of severe space weather events affecting major global industrial regions like the Northeastern USA, central Europe, and Southeast Asia. Macroeconomic models, such as the input-output model we are using in this study, have been used for impact analysis for some time. Such models can be used specifically to provide an estimate of the system-wide impact including those of international trade and global supply chains. We find that a severe Space Weather event could lead to global economic damages of the same order as wars, extreme financial crisis and estimated for future climate change. But some countries may even benefit from the disaster in terms of higher domestic consumption possibilities. A lot of details of the dependencies between solar activity, geomagnetic activity, and failure of electric distribution systems

are still unclear. However, we provided a new physical model that relates the damages to the national power systems to the strength and size of a geomagnetic storm.

## Appendix A: Disaster impact method

5 Assume a disaster analysis setting as in Steenge et al. (2007). In its original form, this method allowed only for changes in consumption possibilities, i.e. reductions in final demand, and excess production available for final demand. There is no provision for situations in which the production loss is larger than total final demand, i.e. where  
10 intermediate demand has to be affected by the disaster. In this study this circumstance is dealt with by introducing sharing parameters, dividing the total damage to a sector between its deliveries to intermediate and to final demand. This way, a situation where damages to final demand are larger than total final demand can always be avoided by setting the share parameter appropriately.

15 Reductions in the production of a damaged sector only affect those intermediate sectors that receive a significant enough input from the damaged sector. Intermediate sectors that receive only marginal inputs from a damaged sector are assumed to be able to substitute for the reduced input, or slightly alter their production recipe otherwise, so they can keep producing at pre-disaster levels. The distinction between marginal and significant inputs is controlled by manually setting a threshold.

20 Let  $\mathbf{A}$  be an  $N \times N$  input coefficients matrix,  $\mathbf{y}$  ( $N \times 1$ ) final demand,  $\mathbf{x}$  ( $N \times 1$ ) total output, and  $\mathbf{I}$  a suitable identity matrix. Define  $\{\tilde{\mathbf{x}}, \tilde{\mathbf{y}}\}$  as the post-disaster quantities of  $\{\mathbf{x}, \mathbf{y}\}$ . As in Steenge et al. (2007), we ask that the post-disaster economy  $\{\tilde{\mathbf{x}}, \tilde{\mathbf{y}}\}$  is in balance:

$$\tilde{\mathbf{x}} = \mathbf{A}\tilde{\mathbf{x}} + \tilde{\mathbf{y}} \Leftrightarrow (\mathbf{I} - \mathbf{A})\tilde{\mathbf{x}} - \tilde{\mathbf{y}} = 0 \Leftrightarrow [\mathbf{I} - \mathbf{A} - \mathbf{I}] \begin{bmatrix} \tilde{\mathbf{x}} \\ \tilde{\mathbf{y}} \end{bmatrix} = 0.$$

25 Introduce damage parameters  $\Gamma$  so that  $\tilde{x}_i = (1 - \Gamma_i)x_i$ .  $\Gamma_i$  is the relative production loss of sector  $i$ .  $1 - \Gamma_i$  is the relative remaining capacity of sector  $i$ . The following approach  
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let part of the production loss affect intermediate demand, so that the loss affecting final demand is never larger than total final demand itself:  $y_i - \tilde{y}_i = \lambda_i(x_i - \tilde{x}_i) = \lambda_i\Gamma_i x_i \Leftrightarrow$   
5  $\tilde{y}_i = y_i - \lambda_i\Gamma_i x_i$ .  $\lambda_i$  is the factor that splits the production loss  $x_i - \tilde{x}_i$  into a fraction  $y_i - \tilde{y}_i$  imposed on final demand, and the remainder on intermediate demand. Assuming a constant production recipe  $\mathbf{A} = \text{const.}$ , a reduction in only one intermediate input  $T_{ij}$  from a damaged sector  $i$  means that the entire production of sector  $j$  must go down in the same proportion as the reduced input  $i$ . Here the production of those sectors  $j$  is reduced, where input  $i$  formed a significant contribution of sector  $j$ 's production recipe. Where this is not the case, sectors  $j$  are allowed to operate at pre-disaster levels of  
10 output.

The loss of production of the electricity sector(s) is represented by  $\Delta x_{el} = \gamma_{el} \cdot x_{el}$ . This loss affects power supply to households and to other industrial, non-electricity sectors, and is distributed according to  $D_{el} = y_{el}/x_{el} = \Delta y_{el}/\Delta x_{el}$ . We assume the fraction of production capacity lost in non-electricity sectors to be  $\gamma_{i \neq el} = \Delta y_{el}/y_{el} = \Delta T/T_{el}$ . We  
15 generally find

$$\frac{\Delta T}{T} = \frac{\Delta x - \Delta y}{x - y} = \frac{\frac{\Delta y}{D} - \Delta y}{y/D - y} = \frac{\Delta y \left(\frac{1}{D} - 1\right)}{y \left(\frac{1}{D} - 1\right)} = \frac{\Delta y}{y}.$$

20 Thereafter final and intermediate demand is relative (not necessarily absolute) curtailed equally. If  $\Delta x = \Delta y$ , i.e.  $\Delta T = 0$ , one would minimize total economic damage, because the indirect impacts on supply chains and international trade are missing from the assessment of the damage caused by the geomagnetic storm.

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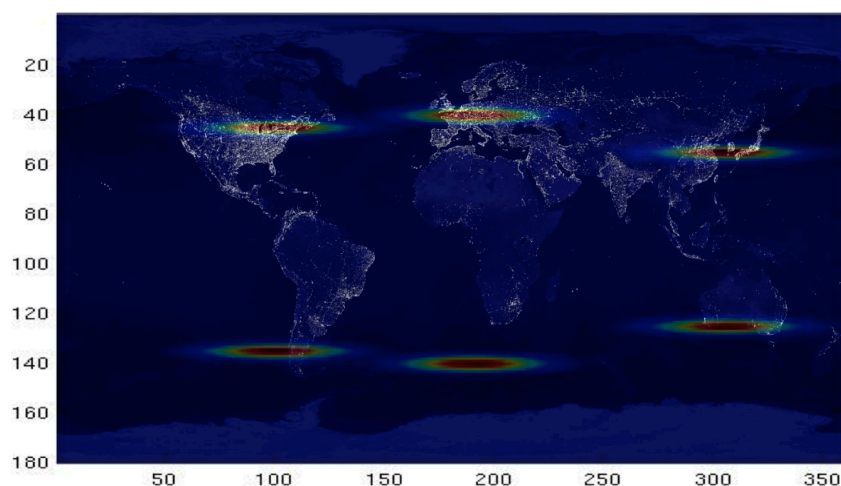
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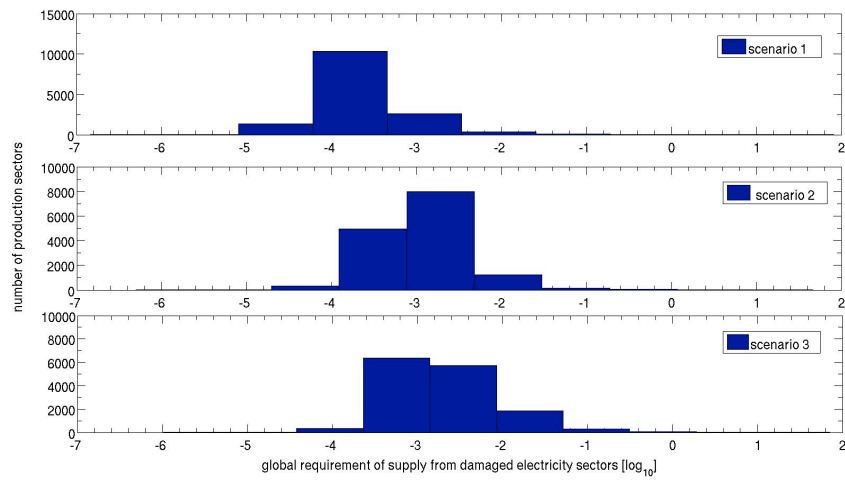
- Shea, M. A., Smart, D. F., McCracken, K. G., Dreschhoff, G. A. M., and Spence, H. E.: Solar proton events for 450 years: the Carrington event in perspective, *Adv. Space Res.*, 38, 232–238, 2006.
- Showstack, R.: Threat of severe space weather to the U.S. electrical grid explored at conference, *EOS T. Am. Geophys. Un.*, 92, 374–375, 2011.
- The human impact report on climate change: The anatomy of a silent crisis, Geneva, available at: <http://www.ghf-ge.org/human-impact-report.pdf> (last access: 4 June 2014), 2009.
- The MITRE Corporation: JASON Report, “Impacts of Severe Space Weather on the Electric Grid”, The MITRE Corporation JASON Program Office, 7515 Colshire Drive McLean, Virginia 22102, USA, 2011.
- Thomson, A. W. P., Dawson, E. B., and Reay, S. J.: Quantifying extreme behavior in geomagnetic activity, *Space Weather*, 9, S10001, doi:10.1029/2011SW000696, 2011.
- Tsurutani, B. T., Gonzalez, W. D., Lakhina, G. S., and Alex, S.: The extreme magnetic storm of 1–2 September 1859, *J. Geophys. Res.*, 108, 1268, doi:10.1029/2002JA009504, 2003.
- Tsurutani, B. T., Verkhoglyadova, O. P., Mannucci, A. J., Lakhina, G. S., and Huba, L. D.: Extreme changes in the dayside ionosphere during a Carrington-type magnetic storm, *Space Weather Space Clim.*, 5, A05, doi:10.1051/swsc/2012004, 2012.
- Tukker, A. and Dietzenbacher, E.: Global multiregional input–output frameworks: an introduction and outlook, *Econom. Syst. Res.*, 25, 1–19, 2013.
- UK House of Commons Defence Committee, Developing Threats: Electro-Magnetic Pulses (EMP), JASON Report, available at: <http://go.nature.com/edi8qf> (last access: 4 June 2014), 2012.
- UNSD: National Accounts Main Aggregates Database, New York, USA, United Nations Statistics Division, available at: <http://unstats.un.org/unsd/snaama/Introduction.asp> (last access: 4 June 2014), 2011.
- US National Academy of Sciences, Severe Space weather Events – Understanding Societal and Economic Impacts, a workshop report, Washington, DC: The National Academies Press, 2008.
- Wei, L. H., Homeier, N., and Gannon, J. L.: Surface electric fields for North America during historical geomagnetic storms, *Adv. Space Res.*, 11, 451–462, 2013.
- Wiedmann, T. O., Wiedmann, T. O., Schandl, H., Lenzen, M., Moran, D., Suh, S., West, J., and Kanemoto, K.: The material footprint of nations, *P. Natl. Acad. Sci. USA*, doi:10.1073/pnas.1220362110, in press, 2013.

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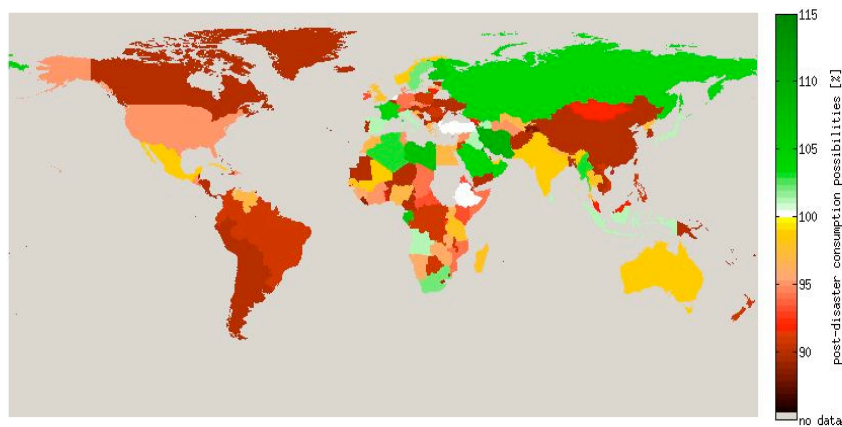
**Figure 1.** Earth at night with Quebec-like events over the Americas (scenario 1), Europe and the southern ocean (scenario 2), and East Asia and Australia (scenario 3). The red area has the highest storm intensity normalized to 1. The storm’s intensity has a Gaussian falloff.

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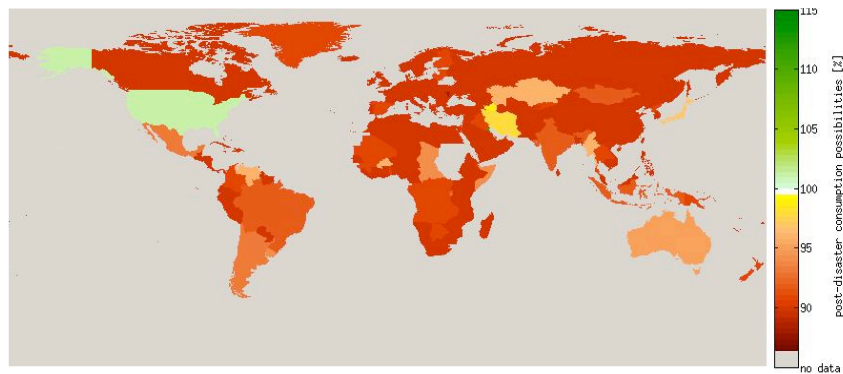
**Figure 2.** Global production requirement of supply from damaged electricity sectors. The threshold for scenario 1 is  $2 \times 10^{-5}$ , for scenario 2  $5 \times 10^{-5}$ , and for scenario 3  $10^{-6}$ .

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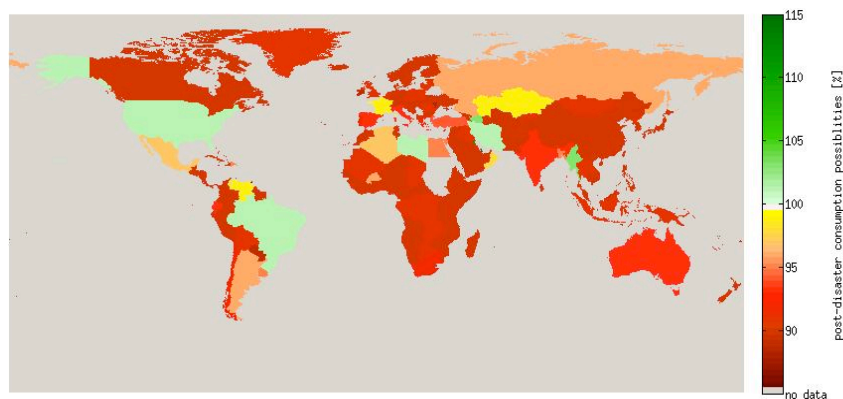
**Figure 3.** Effects of scenario 1, a Quebec 1989-like event centered over the Americas. Globally, the storm would reduce total consumption possibilities by 3.9% though the effect is uneven: it is most severe in countries directly affected and their economic partners, while other countries (e.g. Russia, Saudi Arabia, France, and Egypt) may gain consumption possibilities in the post-disaster economy.

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**Figure 4.** Storm scenario 2, a Quebec-like event centered over Europe. Due to Europe’s participation in many global supply chains, a disaster in Europe would be felt not just in the continent itself but in nearly all other countries in the world as well. The US is a notable exception: that economy could experience a slight increase in consumption possibilities in the post-disaster economy.

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**Figure 5.** In scenario 3, an Australasian storm, the effects again are seen most strongly in the directly impacted countries and their trade partners.

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